



# Article Characteristics of Sasakian Manifolds Admitting Almost \*-Ricci Solitons

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**Abstract:** This article presents some results of a geometric classification of Sasakian manifolds (SM) that admit an almost \*-Ricci soliton (RS) structure (g,  $\omega$ , X). First, we show that a complete SM equipped with an almost \*-RS with  $\omega \neq$  const is a unit sphere. Then we prove that if an SM has an almost \*-RS structure, whose potential vector is a Jacobi vector field on the integral curves of the characteristic vector field, then the manifold is a null or positive SM. Finally, we characterize those SM represented as almost \*-RS, which are \*-RS, \*-Einstein or \*-Ricci flat.

**Keywords:** sasakian manifold; (almost) \*-Ricci soliton; \*-Einstein manifold; infinitesimal contact transformation; conformal vector field

MSC: 53C25; 53C21; 53D15

## 1. Introduction

Nonlinear systems that support solitons can generate self-similarity and fractals on successively smaller scales. Very often, the solitons of a given system are all self-similar to each other, see [1].

Several authors study Ricci solitons (RS) in contact metric geometry, where Sasakian manifolds (SM) are especially important. SM, which are odd-dimensional analogues of K'ahler manifolds, are used to build geometrical examples, e.g., manifolds of special holonomy, Einstein manifolds (EM), K'ahler manifolds and orbifolds. The geometry of SM (in particular,  $\eta$ -EM and Sasaki–Einstein manifolds) attracts much attention from mathematicians. We refer the reader to [2–4] for the latest developments in the theory of these manifolds.

Many tools for K'ahler manifolds have analogues for contact manifolds, among them the \*-Ricci tensor, introduced in [5] for almost Hermitian manifolds. In [6], the \*-Ricci tensor was applied to a real hypersurface in a non-flat complex space form. Namely, for an almost contact manifold  $M^{2n+1}(\varphi, \zeta, \eta, g)$  and any vector fields  $Y_1, Y_2, Y_3$  on  $M^{2n+1}$ , we get

$$\operatorname{Ric}_{g}^{*}(Y_{1}, Y_{2}) = \frac{1}{2}\operatorname{trace}_{g}\left\{Y_{3} \to R(Y_{1}, \varphi Y_{2})\varphi Y_{3}\right\}.$$
(1)

If  $\operatorname{Ric}_{g}^{*}$  vanishes identically then such  $M^{2n+1}$  is called \*-*Ricci flat*. A \*-*RS* structure has been introduced in [7], using  $\operatorname{Ric}_{g}^{*}$  instead of the Ricci tensor in the equation of RS,

$$\pounds_X g + 2\operatorname{Ric}_g^* = 2\,\omega\,g. \tag{2}$$

Here, *f* is the Lie derivative. If  $\omega$  in (2) belongs to  $C^{\infty}(M)$  and is not a constant, then we get an *almost* \*-*RS*, the triple  $(g, \omega, X)$ . An almost \*-RS is *shrinking* for  $\omega > 0$ , *steady* for  $\omega = 0$  and *expanding* for  $\omega < 0$ . Observe that a \*-RS is \*-*EM* (i.e., the \*-Ricci tensor and the metric tensor are homothetic) if X is a Killing vector field (KVF), see [8]. Thus, it is a generalization of a \*-EM. In particular, if  $X = \nabla f$  (the gradient of  $f \in C^{\infty}(M)$ ) in (2),



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**Copyright:** © 2023 by the authors. Licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (https:// creativecommons.org/licenses/by/ 4.0/). then we obtain a *gradient almost* \*-*RS*. Some generalizations of RS were studied by several authors, see [9–15].

Ghosh-Patra [16] first studied \*-RS and gradient almost \*-RS on a contact metric manifold, in particular, on an SM and  $(k, \mu)$ -contact manifolds. Later on, \*-RS were studied on almost contact manifolds in [15,17–20]. Wang [20] proved that "if the metric of a Kenmotsu 3-manifold represents a \*-RS, then the manifold is locally the hyperbolic space  $\mathbb{H}^3(-1)$ ", and it was proved in [18] that "if a non-Kenmotsu  $(k, \mu)$ '-almost Kenmotsu manifold has a \*-RS structure, then it is locally isometric to  $\mathbb{H}^{n+1}(-4) \times \mathbb{R}$  under some restrictions".

There is a natural question: "*under what conditions a (gradient) almost RS is a unit sphere*"? Several affirmative answers to this question on Riemannian and contact metric manifolds, are given in [9,10,12–14,21,22]. An almost \*-RS is a generalization of \*-RS and \*-EM; thus we ask a question: "under what conditions a (gradient) almost \*-RS is a unit sphere?"

In [16], they answered this question by showing that "*if a complete SM admits a gradient almost \*-RS, then it is a unit sphere*". Here, we consider non-gradient almost \*-RS in the case of SM, and we are looking for conditions under which SM having an almost \*-RS structure is a unit sphere. Our main achievement is the following.

**Theorem 1.** If a complete SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  of dimension greater than three has an almost \*-RS structure  $(g, \omega, X)$  with  $\omega \neq \text{const}$ , then it is the unit sphere  $\mathbb{S}^{2n+1}$ .

The following question arises: "under what conditions an almost RS is an RS, or, EM"? Several answers to this question can be found in [14,22–24]. An almost \*-RS generalizes \*-RS and \*-EM; so the following question arises:

"when an almost \*-RS is a \*-RS, for example, an \*-EM?"

In [16], they found such a condition on an SM, and this question on a Kenmotsu manifold was studied in [15]. In this regard, we are interested in characterizing those SM represented as almost \*-RS, which are \*-RS or are \*-EM. In the following theorem, we assume that the potential vector field is a Jacobi vector field on the integral curves of the Reeb vector field.

**Theorem 2.** If an SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  has an almost \*-RS structure  $(g, \omega, X)$  such that X is a Jacobi field on the  $\zeta$ -integral curves, then  $(g, \omega, X)$  is a \*-RS on M.

Now, we recall some definitions, e.g., [3]. A SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  is an  $\eta$ -EM if

$$\operatorname{Ric}_{g} = \alpha_{1} g + \alpha_{2} \eta \otimes \eta,$$

where functions  $\alpha_1$  and  $\alpha_2$  belong to  $C^{\infty}(M)$ . For *K*-contact manifolds (for example, SM) of dimension  $\geq 4$ ,  $\alpha_1$ ,  $\alpha_2$  are real constants, see [25], thus, the scalar curvature is constant. An  $\eta$ -Einstein SM is a *null-SM* when  $\alpha_1 = -2$  and  $\alpha_2 = 2n + 2$ , and is a *positive-SM* when  $\alpha_1 > -2$ , see [3]. Using ([16], Theorem 8), we get the following consequence of Theorem 2.

**Corollary 1.** If an SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  has an almost \*-RS structure  $(g, \omega, X)$  such that X is a Jacobi vector field on the  $\zeta$ -integral curves, then M is positive SM and X is KVF (and g is \*-EM), or M is null-SM and  $\varphi$  is invariant under X.

**Remark 1.** Observe from [24] that if an SM has an almost RS structure, whose potential field is a Jacobi vector field on the  $\zeta$ -integral curves, then the manifold is null-SM with the expanding RS.

A vector field *Y* on a contact manifold  $(M, \eta)$  that preserves the contact form  $\eta$  is called an *infinitesimal contact transformation*, see [26,27], i.e.,

$$\mathcal{L}_{Y}\eta = \nu \,\eta,\tag{3}$$

**Theorem 3.** Let an SM represent an almost \*-RS  $(g, \omega, X)$  such that X is an infinitesimal contact transformation. Then X leaves  $\varphi$  invariant, and the manifold is  $\eta$ -EM of scalar curvature  $(\omega + 4n)n$ .

Assuming compactness we get one more sufficient condition for g to be \*-Ricci flat (e.g., \*-EM).

**Theorem 4.** Let a compact SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  admit an almost \*-RS  $(g, \omega, X)$  such that X is an infinitesimal contact transformation. Then the soliton is steady  $(\omega = 0)$ , X is an infinitesimal automorphism, and g is \*-Ricci flat of scalar curvature  $4n^2$ .

Finally, we show that the property "potential vector field and the characteristic vector field are parallel" provides "\*-Ricci flat" (e.g., \*-EM).

**Theorem 5.** Let an SM  $M^{2n+1}(\varphi, \zeta, \eta, g)$  admit an almost \*-RS  $(g, \omega, X)$  such that X is parallel to  $\zeta$ . Then the soliton is steady ( $\omega = 0$ ), V is KVF, and g is \*-Ricci flat of scalar curvature  $4n^2$ .

**Remark 2.** Observe from [14] that if a compact SM has an almost RS structure  $(g, \omega, X)$  such that X is an infinitesimal contact transformation, then the soliton is shrinking  $(\omega = 2n)$ , X is strict and leaves  $\varphi$  invariant, and the manifold is EM of scalar curvature 2n(2n + 1). By Theorem 4, the soliton is steady  $(\omega = 0)$ , X is strict and leaves  $\varphi$ ,  $\zeta$  and  $\eta$  invariant, and the manifold is \*-EM of scalar curvature  $4n^2$ . In addition, note that an SM considered as an RS or an almost RS with X parallel to  $\zeta$ , is an EM of scalar curvature 2n(2n + 1), and X is a KVF, see [28].

Within the framework of Theorems 3–5, we not only find sufficient conditions for \*-EM, but also characterize the structural tensor fields and the potential vector field.

## 2. Preliminaries

Here, we recall some properties of SM, see [2,26]. We suppose that all manifolds are smooth and connected. The curvature tensor *R* on a Riemannian manifold (M, g) is given by

$$R(Y_1, Y_2) = [\nabla_{Y_1}, \nabla_{Y_2}] - \nabla_{[Y_1, Y_2]}, \quad Y_1, Y_2 \in \mathcal{X}(M),$$
(4)

where  $\nabla$  is the Levi-Civita connection and  $\mathcal{X}(M)$  is the space of vector fields on *M*. They define the Ricci operator *Q* by

$$g(QY_1, Y_2) = \operatorname{Ric}_{g}(Y_1, Y_2) = \operatorname{trace}_{g} \{Y_3 \to R(Y_3, Y_1)Y_2\}, \quad Y_1, Y_2, Y_3 \in \mathcal{X}(M)$$

is a symmetric (1, 1)-tensor. Then  $r = \text{trace}_g Q$  is the scalar curvature. We get the following formula (follows from twice contracted second Bianchi identity, see, e.g., [29]):

$$\frac{1}{2}g(Y_1, \nabla r) = (\operatorname{div} Q)(Y_1) = \sum_i g((\nabla_{E_i} Q)Y_1, E_i), \quad Y_1 \in \mathcal{X}(M),$$
(5)

where  $\{E_i\}$  is any local orthonormal basis on *M*.

For the Lie derivative of  $f \in C^{\infty}(M)$  along  $Y_1 \in \mathcal{X}(M)$  we have

$$\pounds_{Y_1} f = Y_1(f) = g(Y_1, \nabla f),$$

where the Hessian of f is defined by

$$\operatorname{Hess}_{f}(Y_{1},Y_{2}) = \nabla^{2} f(Y_{1},Y_{2}) = (\nabla_{Y_{1}} \nabla f) Y_{2} = Y_{1} Y_{2}(f) - (\nabla_{Y_{1}} Y_{2})(f), \quad Y_{1},Y_{2} \in \mathcal{X}(M).$$

The hessian  $\text{Hess}_f$  is symmetric and  $g(\nabla_{Y_1} \nabla f, Y_2) = g(\nabla_{Y_2} \nabla f, Y_1)$  is true, so it is a section of  $T^*M \otimes T^*M$ . A vector field Y on (M, g) is conformal if

$$\mathcal{L}_Y g = f g$$

for some  $f \in C^{\infty}(M)$ . Such *Y* is called homothetic if f = constant, and *Y* is said to be KVF if f = 0.

A manifold  $M^{2n+1}$  is a contact manifold if  $\eta \wedge (d\eta)^n \neq 0$  for some global 1-form  $\eta$ . In this case,  $\zeta$  satisfying  $d\eta(\zeta, \cdot) = 0$  and  $\eta(\zeta) = 1$  is called a characteristic vector field. Polarization of  $d\eta$  on  $\mathcal{D}_{\equiv} = \ker \eta$  allows to find a (1,1)-tensor  $\varphi$  and a Riemannian structure g satisfying

$$\varphi^2 = -id_{TM} + \eta \otimes \zeta, \quad \eta = g(\zeta, \cdot), \tag{6}$$

$$d\eta(\cdot, \cdot) = g(\cdot, \varphi \cdot). \tag{7}$$

They call such  $M^{2n+1}(\varphi,\zeta,\eta,g)$  a contact metric manifold with associated metric *g*. By (6), we get

$$\varphi(\zeta) = 0$$
,  $\eta \circ \varphi = 0$ , and  $\operatorname{rank}(\varphi) = 2n$ .

Further, a contact metric manifold is called SM if the following condition is valid, see [3,26]:

$$[\varphi Y_1, \varphi Y_2] = -2 d\eta (Y_1, Y_2) \zeta, \quad Y_1, Y_2 \in \mathcal{X}(M).$$

The curvature tensor of an SM satisfies

$$R(Y_1, Y_2)\zeta = \eta(Y_2)Y_1 - \eta(Y_1)Y_2, \quad Y_1, Y_2 \in \mathcal{X}(M).$$
(8)

Recall that if  $\zeta$  is KVF, then *M* is called a *K*-contact manifold. A SM is *K*-contact, the converse is valid in dimension 3 only, see ([26], p. 87). SM satisfy the following, see ([26], p. 113):

$$\nabla_{Y_1}\zeta = -\varphi Y_1, \quad Y_1 \in \mathcal{X}(M), \tag{9}$$

$$Q\zeta = 2\,n\,\zeta.\tag{10}$$

According to ([16], Lemma 2.1):

$$\nabla_{\zeta} Q = Q \varphi - \varphi Q;$$

therefore,  $\nabla_{\zeta} Q = 0$ , as  $\varphi$  and the Ricci operator commute on an SM, see [26]. Using (9) we find the derivative of (10) in the direction of  $Y_1 \in \mathcal{X}(M)$ ,

$$(\nabla_{Y_1}Q)\zeta = Q\varphi Y_1 - 2n\,\varphi Y_1. \tag{11}$$

The \*-Ricci tensor  $\operatorname{Ric}_{g}^{*}$  on an SM has the view (see [16], Lemma 5):

$$\operatorname{Ric}_{g}^{*}(Y_{1}, Y_{2}) = \operatorname{Ric}_{g}(Y_{1}, Y_{2}) - (2n-1)g(Y_{1}, Y_{2}) - \eta(Y_{1})\eta(Y_{2}), \quad Y_{1}, Y_{2} \in \mathcal{X}(M).$$
(12)

Thus, we can write the \*-RS structure (2) on an SM as

$$(\pounds_V g)(Y_1, Y_2) + 2\operatorname{Ric}_g(Y_1, Y_2) = 2(\omega + 2n - 1)g(Y_1, Y_2) + 2\eta(Y_1)\eta(Y_2)$$
(13)

for all  $Y_1, Y_2 \in \mathcal{X}(M)$ .

**Remark 3.** Contracting the derivative of (8) using a local orthonormal basis  $\{E_i\}_{1 \le i \le 2n+1}$  on M and using (9), we get the following:

$$(\operatorname{div} R)(Y_1, Y_2)\zeta - g(R(Y_1, Y_2)\varphi E_i, E_i) = -2g(\varphi Y_1, Y_2).$$

Contracting the second Bianchi identity, they reduce the above equation to the following:

$$g((\nabla_{Y_1}Q)Y_2 - (\nabla_{Y_2}Q)Y_1, \zeta) - g(R(Y_1, Y_2)\varphi E_i, E_i) = -2g(\varphi Y_1, Y_2).$$
(14)

Using (11), from (14) we obtain

$$g(R(Y_1, Y_2)\varphi E_i, E_i) = g(Q\varphi Y_1, Y_2) + g(\varphi QY_1, Y_2) - 2(2n-1)g(\varphi Y_1, Y_2)$$

Replacing  $Y_2$  by  $\varphi Y_2$  in the preceding equation and using the definition (1), we represent the \*-Ricci tensor of an SM in the form (12).

To prove Theorems 1 and 2, we need the following three results.

**Proposition 1.** If an SM admits an almost \*-RS structure, then we get the following:

$$(\pounds_X \nabla)(Y_1, \zeta) + 2 Q \varphi Y_1 = 2(2n-1) \varphi Y_1 + Y_1(\omega)\zeta + \zeta(\omega)Y_1 - \eta(Y_1) \nabla \omega, \quad Y_1 \in \mathcal{X}(M).$$
(15)

**Proof.** Using the derivative of (13) for an arbitrary  $Y_3 \in \mathcal{X}(M)$  and (9), gives

$$(\nabla_{Y_3} \pounds_X g)(Y_1, Y_2) + 2 (\nabla_{Y_3} \operatorname{Ric}_g)(Y_1, Y_2) = 2 Y_3(\omega) g(Y_1, Y_2) - 2 \eta(Y_1) g(Y_2, \varphi Y_3) - 2 \eta(Y_2) g(Y_1, \varphi Y_3)$$
(16)

for  $Y_1, Y_2 \in \mathcal{X}(M)$ . Following Yano ([30] p. 23), for  $Y_1, Y_2, Y_3 \in \mathcal{X}(M)$  we write

$$(\pounds_X \nabla_{Y_3} g - \nabla_{Y_3} \pounds_X g - \nabla_{[X,Y_3]} g)(Y_1,Y_2) = -g((\pounds_X \nabla)(Y_3,Y_1),Y_2) - g((\pounds_X \nabla)(Y_3,Y_2),Y_1).$$

Inserting (16) in the above relation and using  $\nabla g = 0$  yields the following:

$$g((\pounds_X \nabla)(Y_3, Y_1), Y_2) + g((\pounds_X \nabla)(Y_3, Y_2), Y_1) + 2(\nabla_{Y_3} \operatorname{Ric}_g)(Y_1, Y_2) = 2\{Y_3(\omega)g(Y_1, Y_2) - \eta(Y_1)g(Y_2, \varphi Y_3) - \eta(Y_2)g(Y_1, \varphi Y_3)\}.$$

Due to the symmetry  $(\pounds_X \nabla)(Y_1, Y_2) = (\pounds_X \nabla)(Y_2, Y_1)$ , using cyclical permutations of  $Y_1, Y_2, Y_3$  in the preceding equation, we get

$$g((\pounds_X \nabla)(Y_1, Y_2), Y_3) = (\nabla_{Y_3} \operatorname{Ric}_g)(Y_1, Y_2) - (\nabla_{Y_1} \operatorname{Ric}_g)(Y_2, Y_3) - (\nabla_{Y_2} \operatorname{Ric}_g)(Y_3, Y_1) + Y_1(\omega) g(Y_2, Y_3) + Y_2(\omega) g(Y_3, Y_1) - Y_3(\omega) g(Y_1, Y_2) - 2 \eta(Y_2) g(\varphi Y_1, Y_3) - 2 \eta(Y_1) g(\varphi Y_2, Y_3).$$
(17)

Since the Ricci operator *Q* is self-adjoint, using  $\nabla_{\zeta} Q = 0$  and (11), and replacing  $Y_2$  by  $\zeta$  in (17), we get (15).  $\Box$ 

**Proposition 2.** Let an SM represent an almost \*-RS and  $\zeta$  leave  $\omega$  invariant. Then we get

$$R(Y_{1}, Y_{2})\nabla\omega - g(R(Y_{1}, Y_{2})\nabla\omega, \zeta)\zeta = 4\{(\nabla_{Y_{1}}Q)Y_{2} - (\nabla_{Y_{2}}Q)Y_{1}\} - 2\{Y_{1}(\omega)Y_{2} - Y_{2}(\omega)Y_{1}\} + 2\{Y_{1}(\omega)\eta(Y_{2})\zeta - Y_{2}(\omega)\eta(Y_{1})\zeta\} + 2(\omega - 2)\{2g(Y_{1}, \varphi Y_{2})\zeta + \eta(Y_{1})\varphi Y_{2} - \eta(Y_{2})\varphi Y_{1}\} + \eta(Y_{2})\nabla_{Y_{1}}\nabla_{\zeta}\nabla\omega - \eta(Y_{1})\nabla_{Y_{2}}\nabla_{\zeta}\nabla\omega + 2g(\varphi Y_{2}, Y_{1})\nabla_{\zeta}\nabla\omega + g(\varphi Y_{2}, \nabla_{Y_{1}}\nabla\omega)\zeta - g(\varphi Y_{1}, \nabla_{Y_{2}}\nabla\omega)\zeta + g(\zeta, \nabla_{Y_{1}}\nabla\omega)\varphi Y_{2} - g(\zeta, \nabla_{Y_{2}}\nabla\omega)\varphi Y_{1}, \quad Y_{1}, Y_{2} \in \mathcal{X}(M).$$
(18)

**Proof.** Applying the Lie derivative for  $R(Y_1, \zeta)\zeta = Y_1 - \eta(Y_1)\zeta$  for all  $Y_1 \in \mathcal{X}(M)$  (follows from (8)) along X and using (8), yields

$$(\pounds_X R)(Y_1,\zeta)\zeta + R(Y_1,\zeta)\pounds_X\zeta + \eta(\pounds_X\zeta)Y_1 + (\pounds_Xg)(Y_1,\zeta)\zeta + g(Y_1,\pounds_X\zeta)\zeta = 0.$$
(19)

Replacing  $(Y_1, Y_2)$  by  $(Y_1, \zeta)$  in (13) and using (10), we acquire the following:

$$(\pounds_X g)(Y_1, \zeta) = 2\omega \eta(Y_1).$$

Plugging it into the Lie derivative of  $\eta(Y_1) = g(Y_1, \zeta)$  and  $\eta(\zeta) = 1$ , we find  $\eta(\pounds_X \zeta) = -\omega$  and  $(\pounds_X \eta)(\zeta) = \omega$ . Thus, in view of (8), Equation (19) gives us

$$(\pounds_X R)(Y_1,\zeta)\zeta = 2\omega\{Y_1 - \eta(Y_1)\zeta\}, \quad Y_1 \in \mathcal{X}(M).$$

$$(20)$$

By conditions,  $\zeta(\omega) = 0$  is valid, thus (15) reduces to the following:

$$(\pounds_X \nabla)(Y_1, \zeta) + 2 Q \varphi Y_1 = 2(2n-1) \varphi Y_1 + Y_1(\omega)\zeta - \eta(Y_1) \nabla \omega, \quad Y_1 \in \mathcal{X}(M).$$
(21)

Using  $\zeta$ -derivative of (21), gives

$$(\nabla_{\zeta} \pounds_X \nabla)(Y_1, \zeta) = g(Y_1, \nabla_{\zeta} \nabla \omega)\zeta - \eta(Y_1) \nabla_{\zeta} \nabla \omega,$$
(22)

where equalities  $\nabla_{\zeta} Q = \nabla_{\zeta} \zeta = \nabla_{\zeta} \varphi = 0$  for an SM were used. On the other hand, using  $Y_1 = \zeta$  in (21), then differentiating along  $Y_1$  and using (9) and the symmetry of  $\pounds_X \nabla$ , we get

$$(\nabla_{Y_1}(\pounds_X \nabla))(\zeta, \zeta) = 2\,(\pounds_X \nabla)(\varphi Y_1, \zeta) - \nabla_{Y_1} \nabla \omega.$$
<sup>(23)</sup>

Next, differentiating  $g(\zeta, \nabla \omega) = 0$  along  $Y_1 \in \mathcal{X}(M)$  and using (6), gives  $(\varphi Y_1)(\omega) = g(\zeta, \nabla_{Y_1} \nabla \omega)$ ; therefore, it suffices to combine (6), (10), (21) and (23) to arrive at the result

$$(\nabla_{Y_1} \pounds_X \nabla)(\zeta, \zeta) = 4 Q Y_1 - 4(2n-1)Y_1 - 4 \eta(Y_1)\zeta + 2g(\zeta, \nabla_{Y_1} \nabla \omega)\zeta - \nabla_{Y_1} \nabla \omega.$$
(24)

We need the following commutation result, see ([30], p. 23):

$$(\pounds_X R)(Y_1, Y_2)Y_3 = (\nabla_{Y_1} \pounds_X \nabla)(Y_2, Y_3) - (\nabla_{Y_2} \pounds_X \nabla)(Y_1, Y_3).$$
(25)

Next, replacing both  $Y_2$  and  $Y_3$  by  $\zeta$  in (25) and then plugging the values of  $(\pounds_X R)(Y_1, \zeta)\zeta$ ,  $(\nabla_{\zeta} \pounds_X \nabla)(Y_1, \zeta)$  and  $(\nabla_{Y_1} \pounds_X \nabla)(\zeta, \zeta)$  from (20), (22) and (24), respectively, we get

$$\nabla_{Y_1} \nabla \omega = 4 Q Y_1 - 2 \{ \omega + 2(2n-1) \} Y_1 + 2(\omega-2) \eta(Y_1) \zeta + g(\zeta, \nabla_X \nabla \omega) \zeta + \eta(Y_1) \nabla_\zeta \nabla \omega$$
<sup>(26)</sup>

for  $Y_1 \in \mathcal{X}(M)$ . Using (9) in the  $Y_2$ -derivative of (26), we acquire

$$\begin{split} \nabla_{Y_2} \nabla_{Y_1} \nabla \omega &= 4 \big\{ (\nabla_{Y_2} Q) Y_1 + Q(\nabla_{Y_2} Y_1) \big\} - 2Y_2(\omega) \big\{ Y_1 - \eta(Y_1)\zeta \big\} - 2 \big\{ \omega + 2(2n-1) \big\} \nabla_{Y_2} Y_1 \\ &+ 2(\omega-2) \big\{ \eta(\nabla_{Y_2} Y_1)\zeta - g(Y_1, \varphi Y_2)\zeta - \eta(Y_1)\varphi Y_2 \big\} + \big\{ \eta(\nabla_{Y_2} Y_1) \\ &- g(Y_1, \varphi Y_2) \big\} \nabla_{\zeta} \nabla \omega + \eta(Y_1) \nabla_{Y_2} \nabla_{\zeta} \nabla \omega - g(\varphi Y_2, \nabla_{Y_1} \nabla \omega)\zeta \\ &+ g(\zeta, \nabla_{Y_2} \nabla_{Y_1} \nabla \omega)\zeta - g(\zeta, \nabla_{Y_1} \nabla \omega)\varphi Y_2. \end{split}$$

Since  $\text{Hess}_{\omega}$  is symmetric and  $\varphi$  is skew-symmetric, using (26) and the above equation in (4) completes the proof of (18).  $\Box$ 

Recall that the contact metric structure commutes with the Ricci operator, i.e.,  $Q\varphi = \varphi Q$ , e.g., [26]. Its covariant derivative and (5) provide the following.

**Lemma 1** (see [12]). *For an SM M and all*  $Y_1 \in \mathcal{X}(M)$ *, we have* 

(i) 
$$\sum_{i=1}^{2n+1} g((\nabla_{\varphi Y_1} Q) \varphi E_i, E_i) = 0$$
, (ii)  $\sum_{i=1}^{2n+1} g((\nabla_{\varphi E_i} Q) \varphi Y_1, E_i) = -\frac{1}{2} Y_1(r)$ ,

where  $\{E_i\}_{1 \le i \le 2n+1}$  is a local orthonormal basis on M.

#### 3. Proof of Results

**Proof of Theorem 1.** Since the characteristic vector field is KVF,  $\pounds_{\zeta} g = 0 = \pounds_{\zeta}$  Ric<sub>g</sub> is valid. Applying this to the Lie derivative of (13) along  $\zeta$ , and using  $\pounds_{Y_1} \pounds_{Y_2} g - \pounds_{Y_2} \pounds_{Y_1} g = \pounds_{[Y_1, Y_2]} g$ , e.g., [30], we get

$$\pounds_{[X,\zeta]} g = -2\zeta(\omega) g.$$
<sup>(27)</sup>

By (27),  $[X, \zeta]$  is a conformal vector field. This gives us the following alternatives: (I)  $[X, \zeta]$  is non-homothetic, (II)  $[X, \zeta]$  is homothetic.

Okumura [31] proved that "if a complete SM of dimension > 3 has a non-Killing conformal vector field, then it is a unit sphere". Applying this theorem for (I), we conclude that (M, g) is a unit sphere.

We will finish the proof by showing a contradiction for case (II). Sharma [32] proved the following: "*a homothetic vector field on an SM (more generally, K-contact manifold) is necessarily a KVF*". So, (27) implies that  $\zeta$  leaves  $\omega$  invariant. Thus, (18) holds. Contracting it over  $Y_1$  and then using (5), (8), we obtain

$$\operatorname{Ric}_{g}(Y_{2},\nabla\omega) = 4 g(\varphi Y_{2},\nabla_{\zeta}\nabla\omega) + \eta(Y_{2}) \operatorname{div}(\nabla_{\zeta}\nabla\omega) - g(\zeta,\nabla_{Y_{2}}\nabla_{\zeta}\nabla\omega) - 2 g(Y_{2},\nabla r) + 2(2n-1) g(Y_{2},\nabla\omega),$$
(28)

where we used trace<sub>g</sub>  $\varphi = 0 = \varphi \zeta$ , the skew-symmetry of  $\varphi$  and symmetry of Hess<sub> $\omega$ </sub>. Differentiating the equality  $g(\zeta, \nabla_{\zeta} \nabla \omega) = 0$  along  $Y_2 \in \mathcal{X}(M)$  and using (9), gives

$$g(\zeta, \nabla_{Y_2} \nabla_{\zeta} \nabla \omega) = g(\varphi Y_2, \nabla_{\zeta} \nabla \omega).$$

Thus, (28) can be rewritten as

$$\operatorname{Ric}_{g}(Y_{2},\nabla\omega) = 3 g(\varphi Y_{2},\nabla_{\zeta}\nabla\omega) + \eta(Y_{2}) \operatorname{div}(\nabla_{\zeta}\nabla\omega) - 2 Y_{2}(r) + 2(2n-1) Y_{2}(\omega)$$
(29)

for all  $Y_2 \in \mathcal{X}(M)$ . Next, recall the following result for SM, see [26]:

$$R(\varphi Y_2, \varphi Y_1)Y_3 = R(Y_2, Y_1)Y_3 + g(Y_2, Y_3)Y_1 - g(Y_1, Y_3)Y_2 - g(Y_2, Y_3)Y_1 + g(Y_1, Y_3)Y_2$$

for all  $Y_1, Y_2, Y_3 \in \mathcal{X}(M)$ . Substituting  $Y_1 = \varphi Y_1$  and  $Y_2 = \varphi Y_2$  in (18) and using the last formula, in view of (6),  $\varphi \zeta = 0$  and the skew-symmetry of  $\varphi$ , we obtain

$$R(Y_{1}, Y_{2})\nabla\omega - g(R(Y_{1}, Y_{2})\nabla\omega, \zeta)\zeta = 4 \left\{ (\nabla_{\varphi Y_{1}}Q)\varphi Y_{2} - (\nabla_{\varphi Y_{2}}Q)\varphi Y_{1} \right\} + Y_{2}(\omega) \left\{ 2Y_{1} - \eta(Y_{1})\zeta \right\} - Y_{1}(\omega) \left\{ 2Y_{2} - \eta(Y_{2})\zeta \right\} + 4(\omega - 2) g(Y_{1}, Y_{2})\zeta + 2 g(Y_{1}, Y_{2})\nabla_{\zeta}\nabla\omega + 2 \eta(Y_{2}) g(\zeta, \nabla_{Y_{1}}\nabla\omega)\zeta - 2 \eta(Y_{1}) g(\zeta, \nabla_{Y_{2}}\nabla\omega)\zeta - g(Y_{2}, \nabla_{Y_{1}}\nabla\omega)\zeta - g(Y_{1}, \nabla_{Y_{2}}\nabla\omega)\zeta + g(\zeta, \nabla_{Y_{2}}\nabla\omega)Y_{1} - g(\zeta, \nabla_{Y_{1}}\nabla\omega)Y_{2}.$$
(30)

On the other hand, contracting (30) over  $Y_1$  and applying (8), (5), (29) and Lemma 1, we find

$$\eta(Y_2)\operatorname{div}(\nabla_{\zeta}\nabla\omega) + 2(n-1)\{Y_2(\omega) - g(\varphi Y_2, \nabla_{\zeta}\nabla\omega)\} = 0, \tag{31}$$

where we have used the symmetry of Hess<sub> $\omega$ </sub> and that  $\zeta$  leaves  $\omega$  invariant. Next, replacing  $Y_2$  by  $\varphi Y_2$  in (31), noting that (6) and using  $g(\zeta, \nabla_{\zeta} \nabla \omega) = 0$ , we acquire

$$2(n-1)\{g(\varphi Y_2, \nabla \omega) + g(Y_2, \nabla_{\zeta} \nabla \omega)\} = 0, \quad Y_2 \in \mathcal{X}(M).$$
(32)

Furthermore, differentiating  $g(\zeta, \nabla \omega) = 0$  along  $Y_2 \in \mathcal{X}(M)$  and using (9), we achieve

$$g(\zeta, \nabla_{Y_2} \nabla \omega) = g(\varphi Y_2, \nabla \omega).$$

Thus, (32) for n > 1 gives us  $g(\varphi Y_2, \nabla \omega) = 0$ ; consequently,  $\nabla \omega = 0$ . Hence,  $\omega$  is constant – a contradiction with the conditions of the theorem.  $\Box$ 

**Proof of Theorem 2.** Applying Proposition 1 to the well-known formula:

$$\nabla_{Y_1}\nabla_{Y_2}X - \nabla_{\nabla_{Y_1}Y_2}X - R(Y_1, X)Y_2 = (\pounds_X \nabla)(Y_1, Y_2),$$

see ([30], p. 23), we acquire

$$\nabla_{Y_1}\nabla_{\zeta}X - \nabla_{\nabla_{Y_1}\zeta}X - R(Y_1, X)\zeta = -2Q\varphi Y_1 + 2(2n-1)\varphi Y_1 + Y_1(\omega)\zeta + \zeta(\omega)Y_1 - \eta(Y_1)\nabla\omega.$$
(33)

By conditions, X is a Jacobi vector field on the  $\zeta$ -integral curves, see [33], i.e.,

$$\nabla_{\zeta}\nabla_{\zeta}X + R(X,\zeta)\zeta = 0.$$

Using  $Y_1 = \zeta$  in (33) and  $\nabla_{\zeta} \zeta = 0$  (that is a consequence of (9)), we achieve the equality  $\nabla \omega = 2\zeta(\omega)\zeta$ , or, using the exterior derivative,

$$d\omega = 2\zeta(\omega)\eta$$

Applying exterior derivative, the Poincaré lemma ( $d^2 = 0$ ), and the wedge product with  $\eta$ , we acquire  $\zeta(\omega) \eta \wedge d\eta = 0$ ; thus,  $\zeta(\omega) = 0$ , as  $\eta \wedge d\eta$  is nowhere zero on a contact manifold. Thus,  $d\omega = 0$ , i.e.,  $\omega = constant$ .  $\Box$ 

Corollary 4 follows from ([16], Theorem 8) and our Theorem 2.

**Proof of Theorem 3.** Using  $d \circ \pounds_X = \pounds_X \circ d$  (*d* commutes with the Lie derivative) and applying the operator *d* to (3), gives

$$(\pounds_X d\eta)(Y_1, Y_2) = \frac{1}{2} \Big\{ Y_1(\nu) \,\eta(Y_2) - Y_2(\nu) \,\eta(Y_1) \Big\} + \nu \,d\eta(Y_1, Y_2) \tag{34}$$

for a function  $\nu \in C^{\infty}(M)$  and any  $Y_1, Y_2 \in \mathcal{X}(M)$ . Applying the Lie derivative of (7) in the *X*-direction and using (2), (3) and (34), we obtain

$$2(\pounds_X \varphi)(Y_1) + 2(2\omega - \nu + 2(2n-1)) \varphi Y_1 = 4Q\varphi Y_1 + \eta(Y_1) \nabla \nu - Y_1(\nu) \zeta.$$
(35)

From the first equality of (6), for  $Y_1 \in \mathcal{X}(M)$  we get

$$(\pounds_X \varphi)(\varphi Y_1) + \varphi(\pounds_X \varphi)(Y_1) = (\pounds_X \eta)(Y_1)\zeta + \eta(Y_1)\pounds_X \zeta.$$
(36)

As a result of (10), \*-RS Equation (2) gives us

$$(\pounds_X g)(Y_1, \zeta) = 2\omega \eta(Y_1)$$

Taking into account this, as well as (3), it suffices to show that

$$g(\pounds_X \zeta, Y_1) = (\nu - 2\omega) \eta(Y_1). \tag{37}$$

By direct calculation using  $\varphi \zeta = 0$  we get  $(\mathcal{L}_X \varphi)(\zeta) = 0$ ; hence, (35) gives  $\nabla \nu = \zeta(\nu) \zeta$ . Thus, by (9) we get

$$\operatorname{Hess}_{\nu}(Y_{1}, Y_{2}) = Y_{1}(\zeta(\nu)) \eta(Y_{2}) - \zeta(\nu) g(\varphi Y_{1}, Y_{2}).$$
(38)

Since  $\varphi$  is skew-symmetric and Hess<sub> $\nu$ </sub> is symmetric, by (7) and (38), for  $Y_1, Y_2$  orthogonal to  $\zeta$  we achieve

$$\zeta(\nu)\,d\eta(Y_1,Y_2)=0$$

Thus,  $\zeta(\nu) = 0$ , as  $d\eta$  is nonzero; hence,  $\nabla \nu = 0$ . Thus,  $\nu$  is constant. Combining (2), (37) and  $\eta(\zeta) = 1$ , we get  $\nu = \omega$ . Substituting this and (10) in (35), gives

$$(\pounds_X \varphi)(\varphi Y_1) = \varphi(\pounds_X \varphi)(Y_1) = -2 QY_1 + (\omega + 2(2n-1))Y_1 - (\omega - 2) \eta(Y_1)\zeta, \quad (39)$$

where the equality  $Q\varphi = \varphi Q$  (for an SM) has been used, see ([26], p. 116). Now, by (3), (36), (37), (39) and  $\nu = \omega$ , we obtain the relation

$$2\operatorname{Ric}_{g} = (\omega + 2(2n-1))g - (\omega - 2)\eta \otimes \eta.$$
(40)

Hence, (M, g) is an  $\eta$ -EM of scalar curvature  $n(\omega + 4n)$ . Substituting (40) into (35), we find  $\mathcal{E}_X \varphi = 0$ ; thus, *X* leaves  $\varphi$  invariant.  $\Box$ 

**Proof of Theorem 4.** The volume form  $\Omega$  on a contact metric manifold satisfies  $\Omega = \eta \land (d\eta)^n \neq 0$ ; therefore, its Lie derivative in the *X*-direction and (3) give

$$\pounds_X \Omega = (n+1)\nu \,\Omega.$$

Proceeding, we obtain the equality  $\mathcal{L}_X \Omega = (\operatorname{div} X)\Omega$ , from which we deduce  $\operatorname{div} X = (n+1)\nu$ . Applying the Divergence theorem (for compact *M*), this gives  $\nu = 0$ ; consequently,  $\omega = 0$  by the proof of Theorem 3. Thus, (3) and (37), respectively, follows from the condition that *X* leaves  $\eta$  and  $\zeta$  invariant. Moreover, Equation (40) becomes

$$\operatorname{Ric}_{g} = (2n-1)g + \eta \otimes \eta. \tag{41}$$

Thus, from (12) we conclude that (M, g) is \*-Ricci flat of zero \*-scalar curvature  $r^*$ . Using (2), we find that X is KVF. Applying Theorem 3, completes the proof.

**Proof of Theorem 5.** By conditions,  $X = \sigma \zeta$ , where  $\sigma$  is a non-zero smooth function. By the skew-symmetry of  $\varphi$  and (9), we obtain

$$(\pounds_X g)(Y_1, Y_2) = g(\nabla_{Y_1} X, Y_2) + g(\nabla_X Y_2, Y_1) = Y_1(\sigma)\eta(Y_2) + Y_2(\sigma)\eta(Y_1).$$
(42)

Thus, (13) becomes

$$Y_1(\sigma)\eta(Y_2) + Y_2(\sigma)\eta(Y_1) + 2\operatorname{Ric}_g(Y_1, Y_2) = 2(\omega + 2n - 1)g(Y_1, Y_2) + 2\eta(Y_1)\eta(Y_2).$$
(43)

Using  $Y_2 = \zeta$  in (43) and (10), yields

$$Y_1(\sigma) = (2\omega - \zeta(\sigma))\eta(Y_1), \quad Y_1 \in \mathcal{X}(M).$$
(44)

Again, using  $Y_1 = \zeta$  and  $Y_2 = \zeta$  in (43) and applying (10), we get  $\zeta(\sigma) = \omega$ ; hence, from (44) it follows that  $Y_1(\sigma) = \zeta(\sigma) \eta(Y_1)$ , for  $Y_1 \in \mathcal{X}(M)$ . By the above argument, we get that  $\sigma$  is constant. By (42), X is KV, and using (44) we get  $\omega = 0$ . The above reduces (43) to (41), and therefore, g is \*-Ricci flat (follows from (12)) and of constant scalar curvature  $4n^2$ .  $\Box$ 

### 4. Conclusions

A modern geometrical concept of an almost \*-RS can be important in differential geometry and theoretical physics. We study the interaction of this structure on a smooth manifold with the well-known Sasakian structure and prove some results of geometric classification. Theorem 1 contains a condition for a complete SM equipped with an almost \*-RS to be a unit sphere. Theorems 2–5 having local character, contain conditions, ensuring that an SM equipped with an almost \*-RS structure is a \*-RS, e.g., a \*-EM. Using the fact that an almost \*-Ricci tensor on an SM can be written as (12), an almost \*-RS reduces to the form (13), which is less general than the almost  $\eta$ -RS equation for an SM:

$$\frac{1}{2}\mathcal{L}_X g + \operatorname{Ric}_g + \omega g + \delta \eta \otimes \eta = 0.$$
(45)

In connection with the above, the following question arises: "are our Theorems 1–5 true under the condition (45) instead of (13), where  $\omega$  and  $\delta$  are arbitrary smooth functions on *M*?"

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