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## Article

# Acoustic Radiation by 3D Vortex Rings in Air 

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#### Abstract

Acoustic radiation emitted by three-dimensional (3D) vortex rings in air has been investigated on the basis of the unsteady Navier-Stokes equations. Power series expansions of the unknown functions with respect to the initial vorticity which is supposed to be small are used. In such a manner the system of the Navier-Stokes equations is reduced to a parabolic system with constant coefficients at high derivatives. The initial value problem is as follows. The vorticity is defined inside a toroid at $t=0$. Other gas parameters are assumed to be constant throughout the whole space at $t=0$. The solution is expressed by multiple integrals which are evaluated with the aid of the Korobov grids. Density oscillations are analyzed. The results show that the frequency band depends on the initial size of the vortex ring and its helicity. The presented data may be applied to the study of a flow in a wake region behind an aerodynamic body.


Keywords: vortex ring; acoustic radiation; Navier-Stokes equations

## 1. Introduction

Vortical structures (vortex rings and cylindrical vortices) play an important role in the sound radiation of gaseous flows. These structures convert a fraction of their rotational energy into sound waves [1]. The case of the three-dimensional (3D) vortex is of special interest. In this case we have three velocity components $\left(v_{r}, v_{\varphi}, v_{\theta}\right)$, with $r, \varphi, \theta$ being spherical coordinates. The aim of the present investigation is to determine the frequency band of acoustic radiation emitted by 3D vortex rings in air.

The analysis is based on the unsteady Navier-Stokes equations. A new method for solving the equations is set forth. It uses power series expansion of the unknown functions with respect to the initial vorticity which is supposed to be small. By applying this procedure the Navier-Stokes equations are reduced to a parabolic system with constant coefficients. As a result we get the solution in the form of a power series, with multiple integrals being appropriate coefficients. The first term of the series is the main one that determines the properties of acoustic radiation at small vorticity. General questions of sound generation by vortex structures were considered in [2,3]. Acoustic radiation by a solitary vortex ring in incompressible and weakly compressible fluid was analyzed in [4]. There are investigations of the acoustic radiation during the interaction of two vortex rings [5,6]. Linear problems of sound production by vortices are considered in [1]. The influence of vortex rings on properties of turbulent flows is confirmed by numerous experiments [7,8] and by computations [9,10]. Nevertheless, further investigations are necessary in the field of acoustic radiation by a solitary vortex ring in viscous heat-conducting gas. The topic of this paper is the study of the frequency band of acoustic radiation by a 3D solitary vortex ring in air and the evolution of the radiation.

## 2. Governing Equations

We use the Helmholtz decomposition of the velocity field as a potential part and a solenoidal one:

$$
\begin{align*}
v(x, t) & =-\frac{\nabla}{4 \pi} \int_{R^{3}} \frac{s(\xi, t)}{|x-\xi|} d \xi+\frac{\nabla}{4 \pi} \times \int_{R^{3}} \frac{\Omega(\xi, t)}{|x-\xi|} d \xi, \\
s & =\nabla v, \Omega=\nabla \times v, \nabla=\left\{\frac{\partial}{\partial x_{1}}, \frac{\partial}{\partial x_{2}}, \frac{\partial}{\partial x_{3}}\right\} \tag{1}
\end{align*}
$$

Taking into account Equation (1), the Navier-Stokes equations in dimensionless form can be written as follows:

$$
\begin{gather*}
\frac{\partial \Omega_{i}}{\partial t}=v \Delta \Omega_{i}+\frac{3}{4} \varepsilon_{i j l}\left(\frac{\partial v_{l}}{\partial x_{m}}+\frac{\partial v_{m}}{\partial x_{l}}\right) \frac{\partial^{2} h}{\partial x_{j} \partial x_{m}}-v_{j} \frac{\partial \Omega_{i}}{\partial x_{j}}+\Omega_{m} \frac{\partial v_{i}}{\partial x_{m}}+f_{l i}, \\
\frac{\partial w}{\partial t}=-v_{j} \frac{\partial w}{\partial x_{j}}+s, \\
\frac{\partial s}{\partial t}=\frac{1}{\gamma} \exp (h) \Delta w+\frac{4}{3} v \Delta s-\left(\frac{1}{\gamma} \exp (h)+0.5 v s\right) \Delta h-v_{j} \frac{\partial s}{\partial x_{j}}+1.5 v \frac{\partial v_{i}}{\partial x_{j}} \frac{\partial^{2} h}{\partial x_{i} \partial x_{j}}+f_{2},  \tag{2}\\
\frac{\partial h}{\partial t}=\frac{\gamma}{\operatorname{Pr}} \eta \Delta h-(\gamma-1) s-v_{j} \frac{\partial h}{\partial x_{j}}+f_{3}, \\
\Delta=\frac{\partial^{2}}{\partial x_{j} \partial x_{j}}, w=-\log \rho, h=\log T, v=\mu / \rho, \eta=\lambda / \rho, i=1,2,3 ; j=1,2,3 ; l=1,2,3 ; m=1,2,3
\end{gather*}
$$

Here, $\varepsilon_{i j l}$ is the antisymmetrical tensor; $\rho, T, v$ are the dimensionless density, temperature, velocity (divided by $\rho_{0}, T_{0}, c_{0}$, respectively); $\mu, \nu, \lambda, c$ are the viscosity, kinematic viscosity, heat conductivity, and low-frequency sound speed; $\gamma$ is the adiabatic exponent; $\operatorname{Pr}$ is the Prandtl number. The functions $f_{1 i}, f_{2}, f_{3}$ are non-linear terms with respect to the first derivatives over coordinates. Subscript " 0 " refers to the initial state. The system Equation (2) was made dimensionless by using the characteristic length $l_{0}=v_{0} / c_{0}$ and the characteristic time $t_{0}=v_{0} / c_{0}^{2}$.

### 2.1. Initial Value Problem

At the initial instant the vorticity has non-zero values only within a gaseous toroid [11]:

$$
\begin{gather*}
x_{1}=\left(r_{00} u_{1} \sin u_{2}+R_{c}\right) \cos u_{3}, x_{2}=\left(r_{00} u_{1} \sin u_{2}+R_{c}\right) \sin u_{3}, x_{3}=r_{00} u_{1} \cos u_{2}, \\
0<u_{1}<1, \quad 0<u_{2}, u_{3}<2 \pi, \tag{3}
\end{gather*}
$$

$r_{00}, R_{c}$ being the initial dimensions of the vortex ring, namely $r_{00}$ is the radius of its cross-section, $R_{c}$ is the radius of the ring. The problem is considered under the assumption that the dimensionless initial vorticity $\omega_{0}$ is small: $\omega_{0} \ll 1$. The initial conditions are

$$
\begin{equation*}
\Omega_{1}=-\omega_{0} \sin u_{3}, \Omega_{2}=\omega_{0} \cos u_{3}, \Omega_{3}=\alpha \omega_{0} \tag{4}
\end{equation*}
$$

inside the initial toroid,

$$
\begin{equation*}
w(x, 0)=s(x, 0)=h(x, 0)=0 \tag{5}
\end{equation*}
$$

in the whole space. The parameter $\alpha$ refers to helicity. The value of $\alpha$ is $\alpha=0$ (no helicity) and $\alpha=1$ (there is the presence of helicity).

### 2.2. Solution to the Problem

Equation (2) represents a non-linear parabolic system. We seek the solution of the parabolic system as a power series expansion. Taking into account Equation (1) and the initial conditions, we get:

$$
\begin{align*}
& \Omega_{i}(x, t)=\varepsilon \Omega_{i}^{(1)}(x, t)+\varepsilon^{2} \Omega_{i}^{(2)}(x, t)+\varepsilon^{3} \Omega_{i}^{(3)}(x, t)+\ldots \\
& v_{i}(x, t)=\varepsilon v_{i}^{(1)}(x, t)+\varepsilon^{2} v_{i}^{(2)}(x, t)+\varepsilon^{3} v_{i}^{(3)}(x, t)+\ldots \tag{6}
\end{align*}
$$

Inserting Equation (6) into Equations (1) and (2) gives

$$
\begin{align*}
& w(x, t)=\varepsilon^{2} w^{(1)}(x, t)+\varepsilon^{3} w^{(2)}(x, t)+\varepsilon^{4} w^{(3)}(x, t)+\ldots \\
& s(x, t)=\varepsilon^{2} s^{(1)}(x, t)+\varepsilon^{3} s^{(2)}(x, t)+\varepsilon^{(4)} s^{(3)}(x, t)+\ldots  \tag{7}\\
& h(x, t)=\varepsilon^{2} h^{(1)}(x, t)+\varepsilon^{3} h^{(2)}(x, t)+\varepsilon^{4} h^{(3)}(x, t)+\ldots
\end{align*}
$$

We have for the lowest-order functions:

$$
\begin{gather*}
\frac{\partial \Omega_{i}^{(1)}}{\partial t}=\Delta \Omega_{i}^{(1)}, \\
\frac{\partial w^{(1)}}{\partial t}=s^{(1)}, \\
\frac{\partial s^{(1)}}{\partial t}=\frac{1}{\gamma} \Delta w^{(1)}+\frac{4}{3} \Delta s^{(1)}-\frac{1}{\gamma} \Delta h^{(1)}+\psi_{3}^{(1)},  \tag{8}\\
\frac{\partial h^{(1)}}{\partial t}=\frac{\gamma}{\operatorname{Pr}} \Delta h^{(1)}-(\gamma-1) s^{(1)}+\psi_{4}^{(1)}, \\
\psi_{3}^{(1)}=\frac{\partial v_{m}^{(1)}}{\partial x_{j}} \frac{\partial v_{j}^{(1)}}{\partial x_{m}}, \psi_{4}^{(1)}=\frac{1}{2} \gamma(\gamma-1) D_{m j} v^{(1)} D_{m j} v^{(1)}, D_{m j} v^{(1)}=\frac{\partial v_{m}^{(1)}}{\partial x_{j}}+\frac{\partial v_{j}^{(1)}}{\partial x_{m}}
\end{gather*}
$$

From Equation (1) one deduces:

$$
\begin{gather*}
v^{(1)}(x, t)=-\frac{0.03125}{\pi^{5 / 2} t^{3 / 2}} \int_{R^{3}} d \xi \int_{0}^{\infty} d r^{\prime} \int_{0}^{\pi} \sin \theta^{\prime} d \theta^{2} \int_{0}^{2 \pi} d \varphi^{\prime} \Omega^{(1)}(\xi, 0) \times n \exp \left(-\frac{0.25}{t}\left|x+x^{\prime}-\xi\right|^{2}\right),  \tag{9}\\
n=\left\{\sin \theta^{\prime} \cos \varphi^{\prime}, \sin \theta^{\prime} \sin \varphi^{\prime}, \cos \theta^{\prime}\right\}, x^{\prime}=\left\{r^{\prime} \sin \theta^{\prime} \cos \varphi^{\prime}, r^{\prime} \sin \theta^{\prime} \sin \varphi^{\prime}, r^{\prime} \cos \theta^{\prime}\right\}
\end{gather*}
$$

The system Equation (8) consists of three homogeneous parabolic equations with respect to $\Omega_{i}^{(1)}$ and a non-homogeneous parabolic subsystem. All equations have constant coefficients at higher derivatives. The solution to the subsystem can be obtained with the aid of the Fourier transform.

The first Equation (8) yields

$$
\begin{equation*}
\Omega_{i}^{(1)}(x, t)=\frac{0.125}{(\pi t)^{3 / 2}} \int_{R^{3}} \Omega_{i}^{(1)}(\xi, 0) \exp \left(-\frac{0.25}{t}|x-\xi|^{2}\right) d \xi \tag{10}
\end{equation*}
$$

Equation (9) allows us to determine the terms $\psi_{3}^{(1)}, \psi_{4}^{(1)}$. The Fourier transform of the homogeneous parabolic subsystem in Equation (8) gives

$$
\begin{gather*}
\frac{d \tilde{w}^{(1)}}{d t}=\tilde{s}^{(1)}, \\
\frac{d \tilde{s}^{(1)}}{d t}=-\frac{k^{2}}{\gamma} \tilde{w}^{(1)}-\frac{4}{3} k^{2} \tilde{s}^{(1)}+\frac{k^{2}}{\gamma} \tilde{h}^{(1)},  \tag{11}\\
\frac{d \tilde{h}^{(1)}}{d t}=-\frac{\gamma}{\operatorname{Pr}} k^{2} \tilde{h}^{(1)}-\gamma(\gamma-1) \tilde{s}^{(1)}
\end{gather*}
$$

The wavy line denotes the Fourier transform, $k$ being the wave number.
The characteristic equation of the system in Equation (11) is [12]

$$
\begin{equation*}
f^{3}+k^{2}\left(\frac{4}{3}+\frac{\gamma}{\operatorname{Pr}}\right) f^{2}+k^{2}\left(\frac{4 \gamma}{3 \operatorname{Pr}} k^{2}+1\right) f+\frac{k^{4}}{\operatorname{Pr}}=0 \tag{12}
\end{equation*}
$$

At $0<k<k_{*}, \quad k_{*} \approx 1$ for air, the roots of Equation (12) are

$$
\begin{equation*}
f_{1}=\sigma_{1}(k), \quad f_{2,3}=\sigma_{2}(k) \pm i \omega_{r}(k) ; \quad \sigma_{1}, \sigma_{2}<0 \tag{13}
\end{equation*}
$$

At $k>k_{*}$ all roots are real ones and decay rapidly in a very short time, so we do not take this case into account. The dispersion curve $\omega_{r}(k), 0<k<k_{*}$ has two branches. We consider only the branch that refers to smaller values of the attenuation coefficients $\sigma_{1}, \sigma_{2}$.

The fundamental solution matrix $A$ of the subsystem in Equation (11) is given by

$$
\begin{gather*}
A=\left[a_{j 1}\right], a_{j 1}(k, t-\tau)=c_{j 1}(k) \exp \left\{\sigma_{1}(k)(t-\tau)\right\}, \\
a_{j 2}(k, t-\tau)=c_{j 2}(k) \exp \left\{\sigma_{2}(k)(t-\tau)\right\} \cos \left\{\omega_{r}(k)(t-\tau)\right\},  \tag{14}\\
a_{j 3}(k, t-\tau)=c_{j 3}(k) \exp \left\{\sigma_{2}(k)(t-\tau)\right\} \sin \left\{\omega_{r}(k)(t-\tau)\right\}
\end{gather*}
$$

Here $c_{j m}(k)$ must be defined from the initial conditions. Our goal is to investigate the density evolution. We get for the function $w^{(1)}$ :

$$
\begin{equation*}
w^{(1)}(x, t)=\frac{1}{(2 \pi)^{3 / 2}} \int_{0}^{t} d \tau \int_{R^{3}} d k \int_{R^{3}} d \xi \exp \{i k(x-\xi)\}\left\{a_{12} \psi_{3}^{(1)}(\xi, \tau)+a_{13} \psi_{4}^{(1)}(\xi, \tau)\right\} \tag{15}
\end{equation*}
$$

Let us introduce a new variable:

$$
\begin{equation*}
X=\xi-x, \quad X=\left\{R_{3} \sin \theta_{3} \cos \varphi_{3}, R_{3} \sin \theta_{3} \sin \varphi_{3}, R_{3} \cos \theta_{3}\right\} \tag{16}
\end{equation*}
$$

The change of the variables yields:

$$
\begin{equation*}
w^{(1)}(x, t)=\left(\frac{2}{\pi}\right)^{1 / 2} \int_{0}^{t} d \tau \int_{0}^{\infty} k d k \int_{0}^{\infty} \sin \left\{k R_{3}\right\} R_{3} d R_{3} \int_{0}^{\pi} \sin \theta_{3} d \theta_{3} \int_{0}^{2 \pi} d \varphi_{3}\left\{a_{12} \psi_{3}^{(1)}(x+X, \tau)+a_{13} \psi_{4}^{(1)}(x+X, \tau)\right\} \tag{17}
\end{equation*}
$$

The density deviation from its initial value can be written as:

$$
\begin{equation*}
\left(\rho_{d}-\rho_{0}\right) / \rho_{0} \approx-w \approx-\omega_{0}^{2} w^{(1)} \tag{18}
\end{equation*}
$$

Here the subscript " $d$ " denotes a dimensional value.
The function $w^{(1)}$ does not depend on $\omega_{0}$ and neither does the frequency of density oscillations. The functions $w^{(n)}(x, t), s^{(n)}(x, t), h^{(n)}(x, t), n>1$ can be obtained analogously. Thus, the Navier-Stokes equations have been reduced to a parabolic system with constant coefficients at the derivatives. As a result, we get a power series in $\omega_{0}$. The coefficients of the series are known functions of $x, t$ (multiple integrals). The first terms of the series can be used for the analysis of the frequency band of the density oscillations in the case of small vorticity.

## 3. Results and Discussion

Equation (17) was used for the investigation of density evolution. The multiple integral was evaluated with the aid of the Korobov grids [13]. The density field has been studied in the neighborhood of a vortex ring. The parameters of the ring are as follows: $R_{\mathrm{c}}=0.15 \mathrm{~cm} ; r_{00}=0.03 \mathrm{~cm}$, the ratio of the radius of the ring cross-section to the ring radius is equal to 0.2 . The value $\omega_{0}=0.000045$
was used in computations. The density values were analyzed within the initial domain of the toroid (at the point $x_{1}=r_{00}, x_{2}=x_{3}=0$ ) as well as outside it (at the point $x_{1}=x_{2}=0, x_{3}=0.046 \mathrm{~cm}$ ).

Figures 1 and 2 show the dependence of the density on time at the point outside the initial toroid for two cases: (i) there is no helicity (Figure 1); (ii) the helicity is present (Figure 2).


Figure 1. Density oscillations for the case of no helicity.


Figure 2. Density oscillations for the case of present helicity.

As seen, density oscillations arise. High-frequency oscillations ( $f=266 \mathrm{kHz}$ ) are modulated by a low-frequency signal $(f=3.3 \mathrm{kHz})$. First of all, the amplitude of the oscillations increases, then decreases. Later on, the amplitude of the low-frequency oscillations grows and the picture becomes a chaotic one. Figures 3 and 4 show the density oscillations as they decay.


Figure 3. Density oscillations at the final stage of the process (no helicity).


Figure 4. Density oscillations at the final stage of the process (helicity is present).

Figures 5 and 6 represent the density oscillations inside the initial domain of the toroid.


Figure 5. Density oscillations inside the initial vortex ring (no helicity).


Figure 6. Density oscillations inside the initial vortex ring (helicity is present).

One can see that the high-frequency component of the oscillations does not depend on helicity. The presence of helicity has an effect only on the amplitude of the oscillations increasing it. The value of high frequency remains constant during the whole process excluding the final stage of the process (see Figures 3 and 4). The value of high frequency corresponds satisfactorily to the experimental data in [14]. In these experiments an unstable vortex ring appeared behind a shock wave reflected from a concave body. The radius of the ring was equal to 0.1 cm , the frequency of oscillations was $170-220 \mathrm{kHz}$.

## 4. Conclusions

A new method is set forth which allows us to investigate the frequency of acoustic radiations by 3D vortex rings. As shown, helicity has no effect on the frequency of the high-frequency component. The results may be of interest for aeroacoustics.

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## Author Contributions

Fedor V. Shugaev proposed the topic of the present investigation; Dmitri Y. Cherkasov performed the computations for the case of no helicity; Oxana A. Solenaya performed the computations for the case of present helicity.

## Conflicts of Interest

The authors declare no conflict of interest.

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