

Article

Dating Oceanic Subduction in the Jurassic Bangong–Nujiang Oceanic Arc: A Zircon U–Pb Age and Lu–Hf Isotopes and Al-in-Hornblende Barometry Study of the Lameila Pluton in Western Tibet, China

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Abstract: The subduction and close of the Mesozoic Bangong–Nujiang Ocean (BNO) led to a collision of the Lhasa and Qiangtang blocks, which formed the backbone of the Tibetan Plateau (the largest and highest plateau on Earth). However, the detailed subduction processes (in particular, the oceanic subduction processes) within the BNO are still not clear. Here, we focus on the plutonic complex of the oceanic arc in the Bangong–Nujiang suture (BNS) and report field observations on zircon U–Pb ages, Lu-Hf isotopes, and the Al-in-hornblende barometry of quartz diorites from the Lameila pluton in western Tibet. Zircon from the quartz diorites yielded a LA-ICP-MS U-Pb age of 164 Ma. The zircon showed very positive $\varepsilon_{Hf}(t)$ values from 10.5 to 13.9, suggesting the Lameila pluton was likely sourced from the depleted-mantle wedge, which is in contrast with contemporary (164–161 Ma) volcanic rocks in the region that had negative $\varepsilon_{Hf}(t)$ values of -7.4 to -16.2 and a magma source from partial melting of subducted sediments. The Lameila pluton showed a temperature-corrected Al-in-hornblende pressure of 3.9 ± 0.8 kbar, corresponding to an emplacement depth of 13 ± 3 km. Therefore, the thickness of the Jurassic oceanic arc crust must have doubled since the initial growth of the oceanic arc on the BNO crust, with a crustal thickness of 6.5 km during the Middle Jurassic. In combination with previous works on volcanic rocks, this study further supports a two-subduction zone model in association with the BNO during the Middle Jurassic, namely, a north-dipping BNO-Qiangtang subduction zone and an oceanic subduction zone within the BNO. The latter oceanic subduction zone produced the depleted-mantle-derived Lameila pluton and the subducted sediment-derived volcanic rocks in the fore arc.

Keywords: Al-in-hornblende; Lameila pluton; oceanic arc; Bangong-Nujiang; Tibet

1. Introduction

The Tibetan Plateau is the largest plateau on Earth, with an average elevation above 4 km [1] and crustal thickness of 60–80 km [2] over an area of 2.5 million km². It was formed by the sequential amalgamation of continental or oceanic arc blocks (Figure 1) over several orogenic cycles since the Paleozoic, and further enhanced by the collision and continuing convergence of the India and Asia continents since the Cenozoic [3–5]. Knowledge of the pre-Cenozoic tectonics, particularly the Lhasa–Qiangtang collision that followed the closure of the Bangong–Nujiang Ocean (BNO), are critical in understanding the formation processes of the Tibetan Plateau. While much progress has been made recently in understanding the processes of the oceanic-continental subduction and the subsequent



collision between the Lhasa and Qiangtang blocks in the past decades [6–16], the oceanic subduction process within the BNO is still unclear.

Evidence for the oceanic subduction and oceanic arc development come mainly from studies of volcanic rocks with boninitic geochemical characteristics. Zhang et al. [17] reported the occurrence of high-Mg quartz gabbro/diabase within the Bangong–Nujiang suture (BNS) near Dingqing, and proposed an oceanic subduction within the BNO. According to Shi et al. [18], the boninitic andesites in the BNS near Rutog record the fore arc volcanism associated with oceanic subduction. Recent studies on the 164–161 Ma high-Mg andesites from Shiquanhe [19] and Daruco [20] have suggested a Jurassic oceanic arc developed within the BNO. Apart from these significant findings, the configuration of this arc, such as its crustal thickness and the source(s) for the arc magmatism are still unclear.

In this work, we report a mineralogical and isotopic study on the plutonic part of the BNO oceanic arc in western Tibet. Field observations, zircon U–Pb ages and Lu–Hf isotopes, in combination with the Al-in-hornblende barometry of quartz diorites from the Lameila pluton, provide robust new evidence for the development and configuration of the Jurassic oceanic arc within the BNO.

2. Geological Background

2.1. Tectonic Evolution of the Bangong–Nujiang Ocean

The Mesozoic BNO separates the Qiangtang block to the north and the Lhasa block to the south (Figure 1). Rifting, which eventually led to formation of the oceanic basin, started at the Permian–Triassic boundary (Figure 2a). The start of rifting is evidenced by the initiation of separate sedimentary evolutions in the Qiangtang and Lhasa blocks, respectively, from the Triassic onward [21,22]. This is also supported by the oldest age (260–240 Ma) of mafic rocks from the ophiolites in the BNS [9,23,24]. The BNO rapidly developed and became a 5000 \pm 1020 km wide ocean by the Late Triassic, as recorded by the 45.1° \pm 9.2° paleolatitude gap between the Qiangtang and Lhasa blocks [25–27]. The subduction of the BNO beneath the Qiangtang block gave birth to the Jurassic–Cretaceous continental arc along the southern margin of the Qiangtang block (Figure 2b) [13,15,28,29], and caused ophiolite obduction near Dongqiao during the Middle Jurassic [30]. Synchronous to the oceanic-continental subduction, oceanic subduction within the BNO also commenced no later than the Middle Jurassic, and consequently a Jurassic oceanic arc was formed south of the Qiangtang continental arc (Figure 2b) [19,20,31,32]. As these subductions continued, the ocean basin narrowed rapidly. During the Middle Jurassic, the width of the BNO was approximately 2600 \pm 710 km (23.4° \pm 6.4° paleolatitude gap between the Qiangtang and Lhasa blocks) [26].

Whether or not a south-dipping subduction of the BNO occurred is highly debatable (Figure 2c). The main evidence for a south-dipping subduction of the BNO is the occurrence of Early Cretaceous magmatic rocks with depleted isotopic characteristics in the northern margin of the Lhasa block. These magmatic rocks are suggested to be the result of the south-dipping BNO subduction [7,33], however the first-order structures along the BNS do not support this inference [8,34,35]. Alternatively, the Cretaceous magmatic rocks in the northern margin of the Lhasa block could have been formed by the north-dipping subduction of the Indus–Yarlung Tethyan Ocean [36].

The BNO closed during the Early Cretaceous or Late Cretaceous, and the seawater became shallow such that the sedimentary facies changed from marine to non-marine between 125 and 118 Ma [36]. The magmatism changed from low-temperature arc-type to high-temperature bimodal suites (high-Nb basalt and rhyolite with an age of approximately 90 Ma) locally along the BNS at the Early–Late Cretaceous boundary (Figure 2d) [37].

2.2. Geology of the Study Area

The study area is located near Shiquanhe in western Tibet (Figure 1). This area consists of three first-order tectonic units including the south margin of the Qiangtang block, the BNS, and the north margin of the Lhasa block (Figure 3). The Qiangtang block exposes Permian to Triassic bioclastic

limestones intercalated with slates that were intruded by the Jurassic Larelaxin pluton with an age of 168–160 Ma [38,39]. This pluton is the westernmost part of the Jurassic–Cretaceous arc in the south margin of the Qiangtang block [29]. The BNS separates the Qiangtang block to the north and the Lhasa block to the south (Figure 1), and consists of two belts of ophiolitic mélange near Rutog (the west part of the Bangong–Nujiang ophiolite belt, BNOB) and Shiquanhe (the west part of the Shiquanhe-Nam Co ophiolite belt, SNOB) that were imbricated by south-directed thrusts [8,35,40]. Zircon U–Pb ages of the Rutog ophiolitic mélange span from 232 to 167 Ma [31,41], and those of the Shiquanhe ophiolitic mélange range between 193 and 161 Ma [42]. Jurassic flysch crops out in the northern part of the study area, whereas Cretaceous clastic sediments bearing volcanic rocks and limestone crop out in the southern part of the study area. The northern margin of the Lhasa block exposes Permian–Triassic limestone interlayered with shale and the above-mentioned Cretaceous strata. The BNS and Lhasa block were intruded by Late Cretaceous granites.

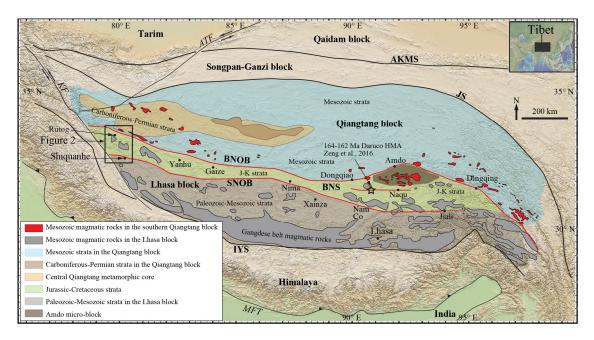


Figure 1. Geologic map of the Tibetan plateau. Base map is from GeoMapApp (www.geomapapp. org). Exposures of central Qiangtang metamorphic core, Carboniferous-Permian and Mesozoic strata in the Qiangtang block are modified after Kapp et al. [43], magmatic rocks in the southern Qiangtang block are modified after Liu et al. [15], and other geologic units are modified after Yin and Harrison [5]. Abbreviations: ATF = Altyn Tagh fault, KF = Karakoram fault, MFT = Main Frontal Thrust, IYS = Indus–Yarlung suture, BNS = Bangong–Nujiang suture, BNOB = Bangong–Nujiang ophiolite belt, SNOB = Shiquanhe–Nam Co ophiolite belt, JS = Jinsha suture, AKMS = Ayimaqin–Kunlun Mutztagh suture. Inset shows the location of Tibet. Black box shows the study area of this paper.

The boninitic andesites and 167 Ma supra-subduction zone (SSZ) type ophiolites [18,31] west of Rutog and the 161 Ma high-Mg andesites [44] north of Shiquanhe are the surface traces of the Jurassic oceanic arc within the BNS. The 164 Ma Lameila pluton newly recognized in this work lies 20 km to the WWN of the Shiquanhe high-Mg andesites. This pluton trends WWN and is 20 m long and 1–3 km wide (Figure 3). It was covered by Jurassic flysch and Cretaceous sedimentary rocks. The Lameila pluton and Jurassic flysch were thrust over the Cretaceous strata by the north-dipping Jiagang thrust [8].

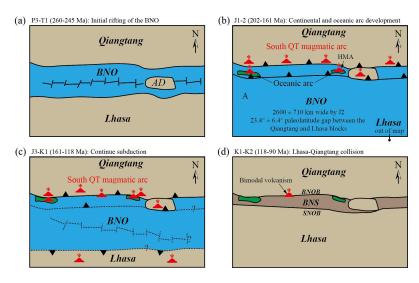


Figure 2. Simplified sketch map showing the possible evolution models of the Bangong–Nujiang Ocean. AD = Amdo micro-block; please see Figure 1 for other abbreviations. (**a**): Initial rifting of the Bangong–Nujiang Ocean at the Permian–Triassic boundary; (**b**): Continental and oceanic arc development during the Early and Middle Jurassic; (**c**): Continue subduction of the Bangong–Nujiang Ocean during the Middle Jurassic to Early Cretaceous; (**d**): The Lhasa–Qiangtang collision between 118 and 90 Ma.

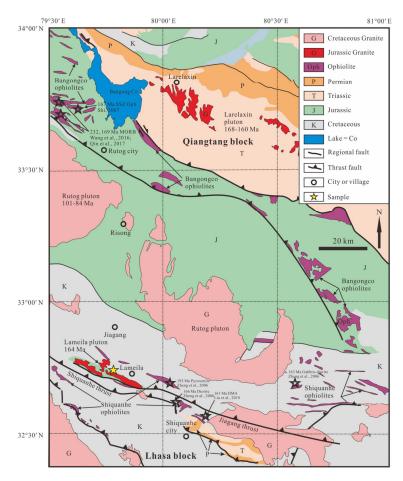


Figure 3. Geologic map of the Shiquanhe area modified after Liu et al. [39]. The sample location for this study is shown as a yellow star, and published ages of ophiolites and their locations are shown as unfilled stars.

3. Samples and Analytical Methods

3.1. Sampling and Petrography

The Lameila pluton consisted mainly of middle- to fine-grained massive quartz diorites that were grayish white in color (Figure 4a). Dark-colored amphibole-rich xenoliths of diorites are in sharp contact with the host quartz diorites (Figure 4b). We collected one quartz diorite sample (18TG49, 32°44'37.77" N, 79°46'40.84" E, H 4604 m) for zircon separation and three quartz diorite samples (18TG50, 18TG51 and 18TG52) for the Al-in-hornblende barometry study. The quartz diorites showed granitic texture, and consisted of plagioclase (55%–60%) + hornblende (10%–15%) + K-feldspar (6%–10%) + quartz (8%–10%) + biotite (1%–2%) + magnetite + zircon. The plagioclase and hornblende were euhedral to subhedral, and most of the K-feldspar were subhedral; anhedral quartz occurred interstitial to the above minerals. Magnetite grains were enclosed in plagioclase, hornblende, and K-feldspar (Figure 4c). The hornblende grains were brown or dark green in color, and showed no signs of zoning patterns in both optical and electron backscatter electron images (Figure 4c,d). Zircon grains could locally be found enclosed in hornblende and plagioclase. Some hornblende crystals were partly altered to chlorite or biotite (Figure 4c). The above petrographic observations suggest that the Lameila quartz diorites contained the full buffering assemblages required for Al-in-hornblende barometry.

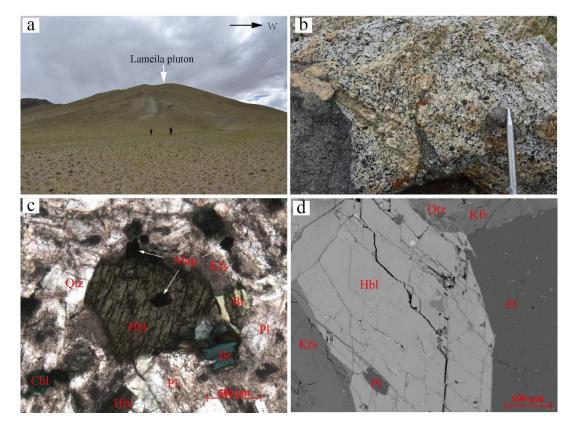


Figure 4. Field (a,b), photomicrograph (c), and backscattered electron (BSE) (d) images of the Lameila pluton. Abbreviations: Qtz = quartz, Hbl = hornblende, Pl = plagioclase, Kfs = K-feldspar, Bt = biotite, Mag = magnetite, Chl = chlorite.

3.2. In-Situ LA-ICP-MS U–Pb Dating and Trace Elements of Zircon

Zircon separation and cathodoluminescence (CL) images were carried out at the Guangzhou Tuoyan Testing Technology Co., Ltd. (Guangzhou, China). Approximately 5 kg of a rock sample was crushed and sieved for separating zircon grains using standard magnetic and heavy liquid separation procedures. About 150 grains were mounted in epoxy and polished to expose the center

of crystals. In situ zircon U–Pb isotopic analyses were conducted using a RESOlution 193 nm ArF excimer laser (LA, Australian Scientific Instruments Pty Ltd., Canberra, Australia) coupled with a ThermoiCAP Qc Inductively Coupled Plasma-Mass Spectrometer (ICP-MS, Thermo Fisher Scientific Inc., Waltham, MA, USA) at the State Key Laboratory of Geological Processes and Mineral Resources, China University of Geosciences (Wuhan). The ablation system operated at a wavelength of 193 nm using a spot diameter of 33 μ m at a 10 Hz repetition rate for 40 s, and ablation pits were about 20 μ m deep. Helium was used as carrier gas to transport the ablated materials to the ICP-MS. Each analysis consisted of a 15 s background acquisition, 40 s sample data acquisition, and 45 s washout delay at the end. Every 10 sample analyses were followed by analysis of one glass standard (NIST 612) and two zircon standards 91500 [45]. Analytical results were calculated using ICPMSDataCal software (Version number, Liu et al., Wuhan, China) [46]. Concordia U–Pb diagrams and weighted mean ²⁰⁶Pb/²³⁸U age calculations were made using Isoplot v2.3 (Berkeley Geochronological Center, Berkeley, CA, USA) [47].

3.3. In-Situ LA-MC-ICP-MS Lu-Hf Isotopes of Zircon

In situ zircon Lu–Hf isotope analyses were carried out using the RESOlution-S155 193 nm ArF excimer laser, which was attached to a Nu Plasma II MC-ICP-MS at the State Key Laboratory of Geological Processes and Mineral Resources, China University of Geosciences (Wuhan). A stationary spot was used for the present analyses, with a beam diameter of 33 μ m. Helium was used as the carrier gas to transport the ablated materials to the MC-ICP-MS. Atomic masses of 172, 173, 175–180, and 182 were simultaneously measured in static-collection mode. The ${}^{176}Lu/{}^{175}Lu = 0.02655$ and ${}^{176}Yb/{}^{173}Yb =$ 0.7965 ratios were used to correct the isobaric interferences of ¹⁷⁶Lu and ¹⁷⁶Yb on ¹⁷⁶Hf, respectively [48]. Penglai zircon was selected as the standard during the analysis, whose analytical results (¹⁷⁶Hf/¹⁷⁷Hf $= 0.282896 \pm 38$, mean square weighted deviation (MSWD) = 0.89, n = 25) during our one-day analyses were in good agreement with the published data within error (0.282906 ± 0.000010) [49]. To calculate the $\varepsilon_{Hf}(t)$ values, we adopted a decay constant for ${}^{176}Lu$ of 1.867×10^{-11} year⁻¹ [50] and chondritic present-day values of ¹⁷⁶Lu/¹⁷⁷Hf (0.0336) and ¹⁷⁶Hf/¹⁷⁷Hf (0.282785) [51]. Depleted-mantle Hf model ages (T_{DM}) were calculated using the measured ${}^{176}Lu/{}^{177}$ Hf ratios of zircon, assuming that the depleted-mantle reservoir had a ¹⁷⁶Hf/¹⁷⁷Hf of 0.283250 at present day, with a ¹⁷⁶Lu/¹⁷⁷Hf value of 0.0384 [52]. The mantle extraction model age (T_{DM}^{C}) for the source rocks of the magmas was calculated by projecting initial ¹⁷⁶Hf/¹⁷⁷Hf ratios of the zircons to the depleted-mantle model growth line using a mean ¹⁷⁶Lu/¹⁷⁷Hf value (0.015) for average continental crust [53].

3.4. EPMA Amphibole Chemistry

Mineral compositions were determined at the State Key Laboratory of Geological Processes and Mineral Resources, China University of Geosciences (Wuhan), with a JEOL JXA-8100 Electron Probe Micro Analyzer (EPMA, JEOL Ltd., Tokyo, Japan) equipped with four wavelength-dispersive spectrometers (WDSs). The samples were coated with a thin conductive carbon film prior to analysis. The precautions suggested by Zhang and Yang [54] were used to minimize the difference of carbon film thickness between samples and standards and obtain an approximately uniform ~20 nm coating. During the analysis, an accelerating voltage of 15 kV, a beam current of 20 nA and a 1- μ m spot size were used to analyze the minerals. Data were corrected online using a modified ZAF (atomic number, absorption, fluorescence) correction procedure. The peak counting time was 10 s for Na, Mg, Al, Si, K, Ca, and Fe, and 20 s for Ti and Mn. The background counting time was one-half of the peak counting time on the high- and low-energy background positions. The following standards were used: sanidine (K), pyropegarnet (Fe, Al), diopsode (Ca, Mg), jadeite (Na), rhodonite (Mn), olivine (Si), and rutile (Ti). Precision of the EPMA analysis was calculated from counting statistics and was generally better than $\pm 1\%$ for contents of >10 wt %, and better than $\pm 5\%$ for contents of >0.5 wt %.

4. Results

4.1. Zircon U–Pb Geochronology and Trace Elements

Zircon grains in sample 18TG49 from the Lameila pluton were euhedral, with sizes that were 100–200 μ m long and 80–120 μ m wide. They showed oscillatory zoning patterns (Figure 5), and had low Th (68–318 ppm) and U (129–460 ppm) concentrations with high Th–U ratios (0.50–0.72). The U–Pb geochronological and trace element data are listed in Tables 1 and 2. Twenty-seven of the 30 total analyses yielded concordant ²⁰⁶Pb/²³⁸U and ²⁰⁷Pb/²³⁵U ages, with ²⁰⁶Pb/²³⁸U ages between 158 and 169 Ma. The weighted mean age was 164.0 ± 1.1 Ma (n = 27, MSWD = 1.3, Figure 6a,b). The oscillatory zoning patterns, Th and U concentrations, and typical magmatic zircon rare earth element (REE) patterns (Figure 6c) imply that this age represents the time of magma emplacement.

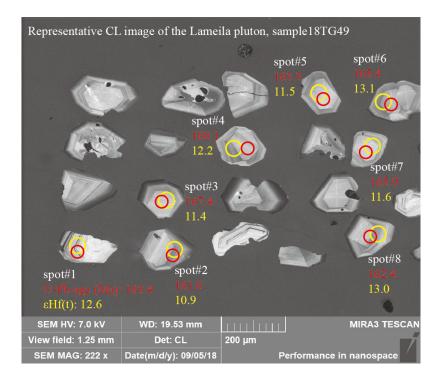


Figure 5. Cathodoluminescence (CL) image of zircons from the Lameila pluton. Red and yellow circles represent analytical spots for LA-ICP-MS zircon U–Pb and LA-MC-ICP-MS zircon Lu–Hf isotope measurements, and the analytical data are also shown.

	Th	U				Isotope	Ratios				Is	otope Age	es (Ma)			Disculture
Analysis	(ppm)	(ppm)	Th/U	²⁰⁷ Pb/ ²⁰⁶ Pb	1s	²⁰⁷ Pb/ ²³⁵ U	1s	²⁰⁶ Pb/ ²³⁸ U	1s	²⁰⁷ Pb/ ²⁰⁶ Pb	1s	²⁰⁷ Pb/ ²³⁵ U	1s	²⁰⁶ Pb/ ²³⁸ U	1s	Discordant (%)
TG-49-1	77.6	131	0.59	0.05031	0.00381	0.17434	0.01198	0.02568	0.00052	209.3	171.3	163.2	10.4	163.4	3.3	99%
TG-49-2	164	278	0.59	0.05090	0.00265	0.17982	0.00949	0.02539	0.00039	235.3	115.7	167.9	8.2	161.6	2.5	96%
TG-49-3	83.8	155	0.54	0.04984	0.00328	0.17730	0.01076	0.02630	0.00049	187.1	155.5	165.7	9.3	167.4	3.1	99%
TG-49-4	89.7	163	0.55	0.04776	0.00348	0.17494	0.01119	0.02642	0.00044	87.1	162.9	163.7	9.7	168.1	2.8	97%
TG-49-5	117	191	0.61	0.04956	0.00300	0.17455	0.00979	0.02566	0.00038	176.0	137.9	163.4	8.5	163.3	2.4	99%
TG-49-6	156	224	0.70	0.04879	0.00297	0.17458	0.01024	0.02600	0.00039	200.1	75.9	163.4	8.9	165.4	2.5	98%
TG-49-7	67.5	135	0.50	0.04839	0.00368	0.17626	0.01280	0.02607	0.00044	116.8	170.3	164.8	11.0	165.9	2.8	99%
TG-49-8	114	195	0.58	0.04969	0.00300	0.17621	0.01065	0.02547	0.00042	189.0	142.6	164.8	9.2	162.2	2.6	98%
TG-49-10	136	240	0.57	0.05038	0.00303	0.18133	0.01004	0.02624	0.00036	213.0	140.7	169.2	8.6	167.0	2.3	98%
TG-49-11	142	263	0.54	0.04984	0.00329	0.17727	0.01167	0.02548	0.00038	187.1	155.5	165.7	10.1	162.2	2.4	97%
TG-49-12	221	326	0.68	0.05038	0.00270	0.17300	0.00914	0.02486	0.00034	213.0	124.1	162.0	7.9	158.3	2.2	97%
TG-49-13	82.6	144	0.57	0.05000	0.00400	0.17835	0.01272	0.02614	0.00051	194.5	177.8	166.6	11.0	166.4	3.2	99%
TG-49-14	318	439	0.72	0.04824	0.00238	0.17222	0.00882	0.02548	0.00032	109.4	124.1	161.3	7.6	162.2	2.0	99%
TG-49-15	128	225	0.57	0.04745	0.00294	0.16620	0.00978	0.02534	0.00042	77.9	135.2	156.1	8.5	161.3	2.7	96%
TG-49-16	125	215	0.58	0.05128	0.00385	0.18500	0.01305	0.02612	0.00041	253.8	169.4	172.4	11.2	166.2	2.6	96%
TG-49-17	217	378	0.57	0.04843	0.00276	0.17221	0.00931	0.02569	0.00035	120.5	129.6	161.3	8.1	163.5	2.2	98%
TG-49-18	137	237	0.58	0.04818	0.00315	0.17349	0.01014	0.02604	0.00041	109.4	148.1	162.4	8.8	165.7	2.6	98%
TG-49-19	217	324	0.67	0.04392	0.00267	0.16182	0.00982	0.02637	0.00042	error	error	152.3	8.6	167.8	2.6	90%
TG-49-20	199	352	0.57	0.04723	0.00283	0.16394	0.00929	0.02495	0.00040	61.2	137.0	154.1	8.1	158.8	2.5	97%
TG-49-21	67.6	129	0.52	0.04950	0.00406	0.17970	0.01408	0.02575	0.00064	172.3	181.5	167.8	12.1	163.9	4.0	97%
TG-49-22	311	459	0.68	0.04772	0.00219	0.16739	0.00723	0.02535	0.00037	87.1	103.7	157.2	6.3	161.4	2.3	97%
TG-49-23	213	350	0.61	0.04815	0.00248	0.17252	0.00879	0.02579	0.00041	105.6	127.8	161.6	7.6	164.2	2.6	98%
TG-49-24	257	379	0.68	0.04773	0.00263	0.16848	0.00896	0.02557	0.00034	87.1	125.9	158.1	7.8	162.8	2.1	97%
TG-49-27	103	176	0.58	0.05028	0.00381	0.18149	0.01243	0.02619	0.00046	209.3	171.3	169.3	10.7	166.6	2.9	98%
TG-49-28	219	331	0.66	0.04438	0.00236	0.15994	0.00824	0.02607	0.00037	error	error	150.6	7.2	165.9	2.3	90%
TG-49-29	123	217	0.57	0.04769	0.00320	0.17719	0.01204	0.02664	0.00039	83.4	151.8	165.6	10.4	169.5	2.5	97%
TG-49-30	186	302	0.62	0.04790	0.00280	0.17132	0.00971	0.02591	0.00041	94.5	133.3	160.6	8.4	164.9	2.6	97%

 Table 1. LA-ICP-MS zircon U–Pb dating for the Lameila pluton.

Table 2. Zircon trace element results for the Lameila pluton.

Spot	Ti	La	Ce	Pr	Nd	Sm	Eu	Gd	Tb	Dy	Но	Er	Tm	Yb	Lu
TG-49-1	5.89	0.5	5.1	0.14	0.66	0.44	0.23	3.62	1.53	23.21	10.45	58.87	14.69	161.1	40.03
TG-49-2	7.02	0.01	7.89	0.02	0.51	0.96	0.69	8.75	4.06	66.63	30.52	170	40.01	415	101
TG-49-3	5.87	0.01	4.54	0.02	0.4	0.98	0.59	6.56	2.54	39.26	18.04	103	25.39	277.0	70.97
TG-49-4	1.56	0	4.63	0.01	0.47	0.93	0.53	6.44	2.65	41.27	18.82	108	26.52	286	73.25
TG-49-5	1.81	0	5.68	0.04	0.49	1.02	0.64	7.64	3.4	50.83	22.36	124	29.57	314	77.54
TG-49-6	3.96	0.01	6.75	0.05	0.82	1.42	0.89	10.54	4.07	60.54	26.81	147	34.56	363	89.11
TG-49-7	3.07	0.01	3.87	0.03	0.32	0.78	0.45	5.01	2.09	33.26	15.22	88.5	21.63	239	60.21
TG-49-8	8.04	0.01	5.64	0.03	0.45	0.84	0.58	7.62	3.15	50.55	22.63	129	31.06	332	81.54
TG-49-10	2.76	0	6.55	0.02	0.43	0.97	0.54	7.39	3.32	52.94	24.4	139.7	33.93	363.0	88.88
TG-49-11	4.84	0	7.28	0.03	0.44	0.98	0.64	7.58	3.28	58.31	26.5	154	37.16	397	96.92
TG-49-12	7.3	0.01	9.09	0.04	0.66	1.24	0.89	12.89	5.28	82.71	36.28	200	46.44	483	115
TG-49-13	4.4	0.01	4.98	0.02	0.35	1.01	0.36	5.81	2.42	38.97	17.51	102.7	25.96	286	71.75
TG-49-14	8.23	0.02	11.99	0.03	0.69	1.62	1.1	15.4	6.81	107	47.3	259	58.79	609	143
TG-49-15	7.23	0.01	6.47	0.03	0.42	1.38	0.72	8.01	3.42	56.47	25.6	151	36.42	394	97.83
TG-49-16	2.09	0.02	6.39	0.04	0.59	1.26	0.74	8.79	3.58	55.17	24.96	145	34.48	378	95.16
TG-49-17	5.06	0.01	10.24	0.03	0.67	1.28	0.76	10.62	5.09	85.17	39.26	224	52.43	554	134
TG-49-18	11.3	6.15	20.91	1.4	6	2.1	0.94	9.67	3.71	59.55	27.09	154	37.35	404	101
TG-49-19	3.06	0.02	9.2	0.03	0.69	1.57	0.95	12.29	5.1	80.69	36.14	209	49.3	527	130
TG-49-20	7.06	0.01	9.58	0.03	0.63	1.38	0.92	11.54	5.2	83.23	37.23	214	50.63	545	132
TG-49-21	5.58	0.02	4.33	0.02	0.42	0.73	0.39	5.6	2.43	35.71	16.05	94.67	24.02	268	68.25
TG-49-22	8.62	0.01	13.4	0.05	0.72	1.36	0.98	13.83	6.78	113	50.14	276	63.66	656	156
TG-49-23	7.88	0.01	10.16	0.04	0.6	1.12	0.77	11.21	5.09	84.17	37.73	211	49.78	517	126
TG-49-24	8.28	0.02	11.04	0.03	0.6	1.52	0.89	12.78	5.72	94.2	41.68	228	52.92	548	132
TG-49-27	10.25	0.08	5.27	0.04	0.46	1.08	0.61	6.56	2.8	44.17	19.46	112	27.12	287	72.33
TG-49-28	5.81	0.01	10.01	0.05	0.79	1.49	0.85	11.44	5.1	84.35	37.16	205	47.97	493	118
TG-49-29	3.45	0.01	6.05	0.04	0.45	1.14	0.74	7.56	3.27	53.26	24.71	140	33.63	356	89.29
TG-49-30	5.14	0.01	8.73	0.02	0.49	1.11	0.63	9.91	4.44	73.33	32.92	181	42.8	446	108

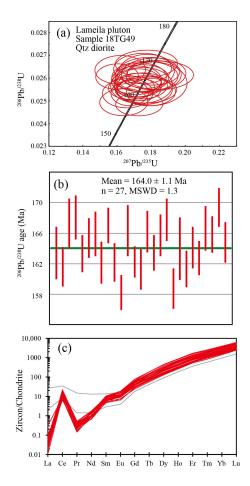


Figure 6. Zircon U–Pb ages (**a**,**b**) and chondrite-normalized [55] rare earth element patterns (**c**). The error bars in (**b**) refer to one sigma. The two outliers (spots TG-49-1 and TG-49-18 in Table 2) shown as gray lines in (**c**) are likely the result of tiny inclusions of mineral or fluid with slightly higher light rare earth elements (LREE) than the host zircon.

4.2. In-Situ Zircon Lu-Hf Isotopes

Zircons from the Lameila pluton had 176 Hf/ 177 Hf ratios ranging from 0.282972 to 0.283068 (Table 3), the initial 176 Hf/ 177 Hf ratios calculated using the corresponding zircon ages were 0.282966–0.283063, and the $\epsilon_{\rm Hf}(t)$ values ranged from 10.5 to 13.9 (Figure 7). These had very young T_{\rm DM} ages of 270–410 Ma, with an average of 340 Ma. The T_{\rm DM}{}^{\rm C} ages were between 320 and 540 Ma, with an average age of 440 Ma.

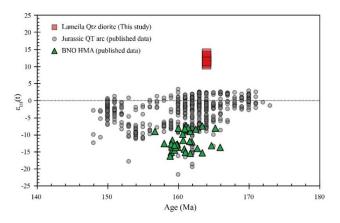


Figure 7. Plot of $\varepsilon_{Hf}(t)$ versus age for the Lameila pluton. Data for the Jurassic Qiangtang arc are from Liu et al. [15]; data for the BNO high-Mg andesite are from Liu et al. [19] and Zeng et al. [20].

Spot	Age (Ma)	¹⁷⁶ Hf/ ¹⁷⁷ Hf	2σ	¹⁷⁶ Yb/ ¹⁷⁷ Hf	2σ	¹⁷⁶ Lu/ ¹⁷⁷ Hf	2σ	(¹⁷⁶ Hf/ ¹⁷⁷ Hf) _i	ε _{Hf} (t)	T _{DM} (Ma)	T _{DM} C (Ma)
18TG-49-1	164	0.283028	0.000013	0.006345	0.000118	0.000327	0.000006	0.283027	12.6	312	405
18TG-49-2	164	0.282985	0.000020	0.035704	0.000358	0.001775	0.000015	0.282980	10.9	386	512
18TG-49-3	164	0.282998	0.000016	0.033158	0.000295	0.001620	0.000015	0.282993	11.4	367	482
18TG-49-4	164	0.283019	0.000015	0.032713	0.000422	0.001657	0.000020	0.283014	12.2	336	433
18TG-49-5	164	0.283000	0.000016	0.031823	0.000122	0.001578	0.000010	0.282995	11.5	363	477
18TG-49-6	164	0.283047	0.000014	0.035129	0.000156	0.001746	0.000008	0.283042	13.1	296	371
18TG-49-7	164	0.283001	0.000016	0.024398	0.000150	0.001198	0.000005	0.282997	11.6	358	473
18TG-49-8	164	0.283042	0.000017	0.032354	0.000163	0.001599	0.000013	0.283037	13.0	302	381
18TG-49-10	164	0.282972	0.000015	0.039386	0.000686	0.001887	0.000033	0.282966	10.5	407	542
18TG-49-11	164	0.283068	0.000016	0.037275	0.000264	0.001773	0.000014	0.283063	13.9	265	323
18TG-49-12	164	0.282999	0.000018	0.030743	0.000285	0.001501	0.000012	0.282994	11.5	364	479

 Table 3. LA-MC-ICP-MS zircon Lu–Hf isotopes of the Lameila pluton.

4.3. Amphibole Chemistry and Stoichiometry

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Results of the EPMA analyses for amphibole chemistry are shown in Table 4. Locations for EPMA analysis were near the margins in contact with quartz or K-feldspar. The amphibole compositions of the three quartz diorite samples were broadly the same; therefore, we present the major element compositions and assess the dominance of atomic exchange reactions by treating all of the amphibole crystals as one population. All the amphiboles from the Lameila pluton displayed a narrow compositional range in SiO₂ (45.33-49.84 wt %, average = 47.13 wt %) and CaO (10.30-11.88 wt %, average = 11.23 wt %). Their Al₂O₃ contents were uniform and low, ranging between 6.32 and 8.81 wt % (average = 7.86 wt %), there are some variation for MgO (11.81-15.41 wt %, average = 12.96 wt %) and total FeO (12.52-17.25 wt %, average = 15.23 wt %) compositions (Table 4).

Spot	TG50-1	TG50-2	TG50-3	TG50-4	TG50-5	TG50-6	TG50-7	TG50-8	TG51-1	TG51-2	TG51-3
-											
SiO ₂ TiO ₂	46.05 0.57	47.05 0.63	48.14 1.10	46.62 0.63	47.02	46.96 0.70	46.44 0.76	46.06 0.63	47.32	46.64 0.59	49.84 0.74
				0.63 7.97	0.68				0.56		
Al ₂ O ₃	8.09	7.98	7.23		7.87	7.77	8.81	8.11	7.79	8.21	6.32
FeO _{total}	17.00	15.13	13.58	15.87	15.86	16.19	17.25	16.26	15.66	16.23	12.52
MnO	0.71	0.69	0.45	0.72	0.73	0.78	0.81	0.72	0.59	0.52	0.43
MgO	11.82	12.65	14.91	12.25	12.57	12.20	12.18	12.01	12.52	11.81	15.41
CaO	11.50	11.25	10.30	11.14	11.59	11.56	11.13	11.70	11.45	11.88	10.54
Na ₂ O	0.87	1.07	1.15	1.10	1.07	0.94	1.14	0.90	0.99	0.93	0.99
K_2O	0.55	0.46	0.22	0.51	0.32	0.47	0.30	0.57	0.29	0.58	0.16
Total (wt %)	97.16	96.91	97.08	96.81	97.71	97.57	98.82	96.96	97.17	97.39	96.95
			For	mulae base	ed on 23 ox	ygen and 1	13eCNK.				
T site											
Si	6.77	6.87	6.84	6.84	6.84	6.86	6.66	6.79	6.90	6.86	7.07
Al(IV)	1.23	1.13	1.16	1.16	1.16	1.14	1.34	1.21	1.10	1.14	0.93
C site											
Al(VI)	0.17	0.24	0.05	0.22	0.18	0.19	0.15	0.20	0.24	0.28	0.13
Ti	0.06	0.07	0.12	0.07	0.07	0.08	0.08	0.07	0.06	0.07	0.08
Fe ³⁺	0.96	0.84	1.37	0.90	0.86	0.82	1.24	0.81	0.83	0.61	1.14
Mg Fe ²⁺	2.59	2.75	3.16	2.68	2.72	2.66	2.60	2.64	2.72	2.59	3.26
	1.13	1.01	0.24	1.05	1.07	1.15	0.83	1.19	1.08	1.39	0.34
Mn	0.09	0.09	0.05	0.09	0.09	0.10	0.10	0.09	0.07	0.06	0.05
B site											
Ca	1.81	1.76	1.57	1.75	1.81	1.81	1.71	1.85	1.79	1.87	1.60
Na	0.19	0.24	0.32	0.25	0.19	0.19	0.29	0.15	0.21	0.13	0.27
A site											
Na	0.06	0.06	0.00	0.06	0.11	0.07	0.03	0.10	0.07	0.14	0.00
K	0.10	0.09	0.04	0.10	0.06	0.09	0.05	0.11	0.05	0.11	0.03
	Pressure (l	kbar), temp	oerature (°	C), and dep	oth (km) of	emplacem	ent at a cru	ıstal densit	ty of 2.95 g	/cm ³	
T _{O84}	621	628	687	629	635	638	644	629	619	624	640
T _{EL98}	666	633	642	644	644	638	701	659	624	636	569
P _{HZ86}	3.1	3.0	2.2	3.0	2.9	2.8	3.6	3.2	2.8	3.2	1.4
P _{JR89}	2.5	2.3	1.7	2.4	2.2	2.2	2.8	2.5	2.2	2.6	1.0
P _{S92}	3.7	3.5	2.8	3.5	3.4	3.4	4.1	3.7	3.4	3.8	2.0
P _{A97 S92}	3.6	3.5	2.9	3.5	3.4	3.3	3.9	3.6	3.3	3.6	2.3
P _{EL98}	4.6	6.3	4.7	6.0	4.8	4.9	5.6	4.4	5.8	5.3	5.1
P _{H87}	3.1	3.0	2.1	3.0	2.8	2.8	3.6	3.2	2.8	3.3	1.2
PAS95 084	4.0	3.8	2.7	3.8	3.6	3.6	4.3	4.0	3.6	4.1	2.2
Depth	13.4	12.9	9.0	12.9	12.4	12.1	14.6	13.4	12.3	13.8	7.4
Spot	TG51-4	TG51-5	TG51-6	TG51-7	TG51-8	TG52-1	TG52-2	TG52-3	TG52-4	TG52-5	TG52-6
SiO ₂	45.53	47.15	47.26	46.75	47.45	48.14	45.87	46.71	46.95	48.96	46.77
TiO ₂	0.64	0.72	0.71	0.57	0.79	0.98	0.84	0.83	0.68	1.00	0.64
Al ₂ O ₃	8.18	8.06	8.02	8.02	8.30	6.99	8.61	8.38	8.04	6.70	8.06
FeO _{total}	15.07	13.90	15.50	15.96	15.35	13.69	16.11	14.46	15.31	12.93	15.80
MnO	0.55	0.52	0.59	0.53	0.60	0.47	0.72	0.53	0.76	0.40	0.74
MgO	11.91	13.12	12.26	12.70	13.29	14.25	12.10	13.54	12.71	14.97	12.56
CaO	11.49	11.36	11.37	11.27	10.93	10.87	11.44	10.92	11.57	10.54	11.51
Na ₂ O	1.03	1.33	0.96	0.94	0.97	0.94	1.06	1.30	1.07	1.05	1.06
K ₂ O	0.60	0.49	0.98	0.94	0.97	0.94	0.59	0.41	0.46	0.19	0.41
Total (wt %)	95.00	96.65	97.12	97.22	98.05	96.56	97.34	97.08	97.55	96.74	97.55
10tai (Wt 70)	95.00	90.00	91.1Z	71.22	90.00	90.00	97.34	27.00	97.00	70.74	97.00

Spot	TG50-1	TG50-2	TG50-3	TG50-4	TG50-5	TG50-6	TG50-7	TG50-8	TG51-1	TG51-2	TG51-3
			For	mulae bas	ed on 23 o	kygen and	13eCNK				
T site											
Si	6.84	6.90	6.90	6.80	6.78	6.94	6.72	6.75	6.83	6.99	6.81
Al(IV)	1.16	1.10	1.10	1.20	1.22	1.06	1.28	1.25	1.17	1.01	1.19
C site											
Al(VI)	0.29	0.29	0.29	0.18	0.18	0.12	0.20	0.18	0.21	0.12	0.19
Ti	0.07	0.08	0.08	0.06	0.08	0.11	0.09	0.09	0.07	0.11	0.07
Fe ³⁺	0.60	0.62	0.74	1.03	1.19	1.06	0.89	1.07	0.81	1.13	0.90
Mg	2.67	2.86	2.67	2.75	2.83	3.06	2.64	2.92	2.76	3.19	2.73
Mg Fe ²⁺	1.29	1.09	1.15	0.91	0.64	0.59	1.08	0.68	1.05	0.41	1.03
Mn	0.07	0.06	0.07	0.07	0.07	0.06	0.09	0.06	0.09	0.05	0.09
B site											
Ca	1.85	1.78	1.78	1.76	1.67	1.68	1.79	1.69	1.80	1.61	1.79
Na	0.15	0.22	0.22	0.24	0.27	0.26	0.21	0.31	0.20	0.29	0.21
A site											
Na	0.15	0.16	0.05	0.02	0.00	0.00	0.10	0.06	0.11	0.00	0.09
Κ	0.12	0.09	0.08	0.09	0.07	0.04	0.11	0.08	0.09	0.03	0.08
	Pressure (kbar), temp	oerature (°	C). and dep	oth (km) of	emplacem	nent at a cr	ustal densi	ty of 2.95 g	/cm ³	
T _{O84}	632	640	639	620	647	673	656	654	635	674	629
T _{EL98}	641	623	622	655	662	612	682	671	645	595	653
P _{HZ86}	3.4	3.1	3.0	3.0	3.1	2.1	3.6	3.3	3.0	1.7	3.0
P _{IR89}	2.7	2.4	2.4	2.4	2.5	1.6	2.8	2.6	2.4	1.3	2.4
P _{S92}	3.9	3.6	3.6	3.5	3.6	2.6	4.1	3.8	3.6	2.4	3.6
PA97 592	3.7	3.5	3.5	3.5	3.6	2.8	3.9	3.7	3.5	2.6	3.5
P _{EL98}	5.7	6.6	6.6	5.4	5.8	5.0	5.2	6.3	5.2	5.2	5.1
P _{H87}	3.4	3.1	3.0	3.0	3.1	1.9	3.6	3.3	3.0	1.6	3.0
PAS95 084	4.2	3.8	3.8	3.8	3.8	2.7	4.2	3.9	3.8	2.4	3.8
Depth	14.1	13.0	12.9	12.9	13.0	9.0	14.3	13.4	12.9	8.0	13.0

Table 4. Cont.

Note: Subscripts for T and P represent initials of names of authors and years of papers, see text for details. T, C, B, and A site represent the corresponding site in amphibole formula. The preferred values for pressure are marked in bold.

The structural formulae and stoichiometry of amphiboles were calculated based on an anhydrous basis assuming 23 oxygen and 13 cations, excluding Ca, Na, and K (13eCNK). Fe³⁺ and Fe²⁺ contents were calculated assuming a charge of 46, as suggested by Leake et al. [56]. All of the Lameila amphiboles were calcic (Na_B = 0.13–0.32 atoms per formula unit (apfu), Ca_B + Na_B = 1.70–2.19 apfu) and are classified as magnesiohornblende (Mg# (Mg/(Mg + Fe²⁺)) = 0.65–0.93, Si = 6.66–7.07 apfu, Figure 8). Leake et al. [56] suggested that igneous amphiboles have a maximum of Si = 7.3 apfu, and those with Si > 7.3 apfu were crystallized under sub-solidus conditions in the presence of a fluid or had been altered by hydrothermal fluids; therefore, the low Si signature of the Lameila hornblende suggests a magmatic origin, which is in accordance with the petrographic observations.

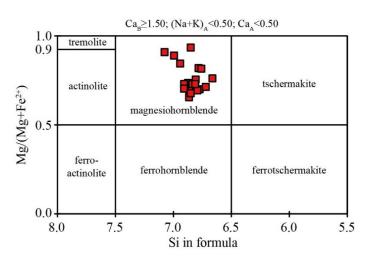


Figure 8. Classification of the Lameila amphiboles following Leake et al. [56].

5. Discussion

5.1. Key Control Factors on the Variation of Amphibole Compositions

The amphibole compositions are sensitive to physical–chemical conditions in magma systems, such as changing melt composition, pressure, temperature, volatile content, and co-crystallizing mineral phases [57]. Therefore, it is important to determine the dominating parameters that are responsible for the compositional variations of minerals. Variations in Al for the Lameila hornblendes are largely accommodated by the tetrahedral site (Al^{IV} = 0.93–1.34 apfu). The amount of Al in the C site (Al^{VI} = 0.05–0.29 apfu) is minor. Atom exchanges, such as the Al–Tschermak exchange, Ti–Tschermak exchange, the edenite exchange, and the plagioclase substitution are proposed to be significant mechanisms controlling compositional variations in amphiboles.

Elemental correlation diagrams of the Lameila amphiboles (Figure 9) imply that the contributions from both the pressure-sensitive Al–Tschermak exchange and the temperature-sensitive Ti–Tschermak exchange were negligible, as evidenced by the lack of covariations of Al^{VI} and Ti_C with Al^{IV} (Figure 9a,b). Instead, the slopes of the regression lines in the (Na + Ca)_A–Al^{IV} and the Ca_B–Al^{IV} diagrams (Figure 9b,c) suggest that about 55% of the Al^{IV} variations were accommodated by the temperature-sensitive edenite exchange and plagioclase substitution. The remaining proportions for the Al^{IV} and total Al variation may have been controlled by some other factors, such as fractional crystallization of the host magma, or changes in the redox state during the crystallization of the Lameila amphiboles (Figure 9d–f). The effects of the fractional crystallization processes and the changes of the redox state on Al variation may have somewhat overlapped with those of the edenite exchange and plagioclase substitution; however, the extent to which the former two factors contributed to the Al variations of the Lameila amphiboles cannot be quantitatively evaluated in the current work. Nevertheless, the existence of the temperature-sensitive edenite exchange and the plagioclase substitution in the Lameila samples demands temperature correction for the Al-in-hornblende barometer.

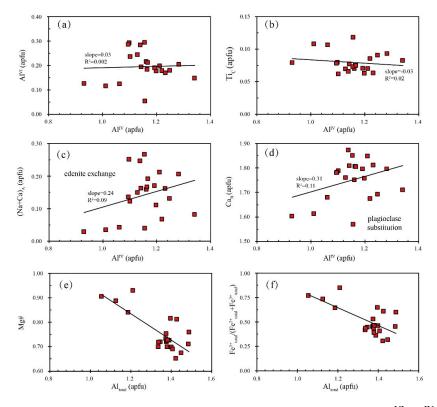


Figure 9. Composition variation diagrams for the Lameila amphiboles. (a) $Al^{VI}-Al^{IV}$ diagram; (b) Ti_C-Al^{IV} diagram; (c) $(Na + Ca)_A-Al^{IV}$ diagram; (d) Ca_B-Al^{IV} diagram; (e) Mg#-Al_{total} diagram; (f) $Fe^{3+}_{total}/Fe^{2+}_{total}-Al_{total}$ diagram.

5.2. Al-in-Hornblende Barometer and Depth of the Lameila Pluton Emplacement

Empirical calibration and experimental studies show that the total Al composition in calcic amphiboles increases linearly with crystallization pressure. This phenomenon forms the foundation of Al-based barometers for calcic amphiboles [57–61]. An important prerequisite for the applications of Al-in-hornblende barometers is an appropriate buffering assemblage (amphibole + plagioclase + K-feldspar/quartz, with medium to high oxygen fugacity). The lithology and petrography of the Lameila pluton (Figure 4) show that the above buffering assemblage was satisfied. In addition, the magma in which the Lameila amphiboles crystallized had a high oxygen fugacity. This inference is supported by the elevated Ce/Ce* and Eu/Eu* in the Lameila zircons (Figure 10) and the occurrence of primary magnetite grains within and along the margins of amphibole crystals (Figure 4c).

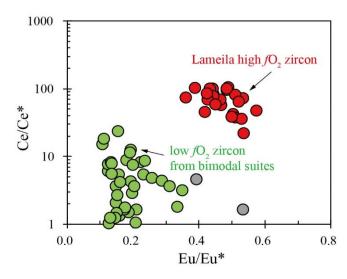


Figure 10. Plot of Ce/Ce* versus Eu/Eu* for zircons for the Lameila pluton. The data for bimodal volcanic suites are from Liu et al. [37].

We first calculated the emplacement pressure using Al-in-hornblende barometers without temperature correction. For this purpose, we employed several widely used barometers, including calibrations of P_{HZ86} [61], P_{H87} [60], P_{JR89} [59], P_{S92} [62], and P_{EL98} [58]. Results of the calculated pressures are shown in Table 4.

We then used the temperature-corrected calibration of Anderson and Smith [57] to evaluate the emplacement pressure. The temperatures for amphibole crystallization were estimated by employing several kinds of methods (Figure 11). The first ($T_{0.84}$) was the Ti-in-amphibole calibration of Otten [63], which was first reported by Helz [64]. The second (T_{EL98}) was the amphibole thermometer of Ernst and Liu [58]. Results of these two thermometers were further compared with temperatures (T_{Ti-in-zircon}) derived from the Ti-in-zircon thermometer [65]. Ranges of T_{O84} (619–687 °C, average = 642 °C) and T_{EL98} (569–701 °C, average = 641 °C) largely overlapped with each other, but the results of T_{EL98} were more scattered than T_{O84}. The T_{Ti-in-zircon} can record the temperature at which a zircon grain crystallizes in the host magma. Stages for zircon crystallization in an intermediate/felsic magma could be either early or late, and thus T_{Ti-in-zircon} of different zircon grains could record whole-stage temperatures of a magma system. Amphiboles crystallized early in dioritic magma during emplacement, and overlap between the ranges of $T_{\text{Ti-in-zircon}}$ and T_{O84} was the best estimation for the temperatures for amphibole crystallization. The trend of $T_{Ti-in-zircon}$ with increasingly younger ages was likely the result of heating by magma replenishment after the crystallization of zircon grains with older ages (Figure 11c). This batch of magma replenishment was recorded in the dark-colored amphibole-rich xenoliths of diorites that were in sharp contact with the host Lameila quartz diorites. The replenished magma probably delivered heat but not materials to the quartz diorites, which is evidenced by the linearly decreasing Mg with Al (Figure 9e). On this basis, the T_{O84} results are used in this study

for correction of the Al-in-hornblende barometer (P_{AS95_TO84}) proposed by Anderson and Smith [57]. The corrected P_{AS95_TO84} ranged from 3.6 to 4.3 kbar except for four outliers (2.2–2.7 kbar, Figure 11d). The outliers with elevated temperatures were likely the result of heating as recorded by the $T_{Ti-in-zircon}$ trend (Figure 11c). The average value of the majority was 3.89 ± 0.28 kbar (n = 18, MSWD = 0.11), which was interpreted as the pressure of the Lameila pluton emplacement.

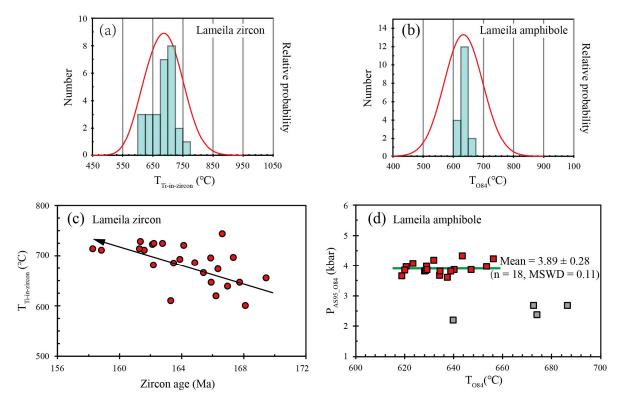


Figure 11. Plots of $T_{Ti-in-zircon}$ (**a**), T_{O84} (**b**), $T_{Ti-in-zircon}$ versus zircon age (**c**), and P_{AS95_O84} versus T_{O84} (**d**) for the Lameila pluton. Please see text for explanations. In panel d, data used for the calculation of the mean value are shown as red squares, and the outliers are shown as grey squares.

Uncertainties for the pressure estimation were the followings: (1) analytical uncertainties are better than 1%; (2) uncertainties with the Al-in-hornblende calibration are within 0.6 kbar; (3) the standard deviation for P_{AS95_TO84} results is 0.19 kbar. Taken all together, the uncertainties for the pressure would be better than 0.83 kbar. As a consequence, the pressure for the Lameila pluton is 3.89 ± 0.83 kbar. The average continental crust density is 2.700-2.875 g/cm³, and the oceanic crust is 2.700-2.950 g/cm³ [66]. The density of the oceanic arc crust in the BNO is not clear, but should be between the above two crusts. If we employ a conservative density of 2.950 g/cm³, the depth of the Lameila pluton emplacement would be 13.2 ± 2.8 km.

5.3. Mechanism for the Crustal Thickening of the Jurassic Oceanic Arcin the BNO

Previous studies on high-Mg volcanic rocks within the BNS suggest a Jurassic oceanic arc that formed during the intra-oceanic subduction of the BNO. However, the configuration of this oceanic arc, including the crustal thickness, is not clear. As mentioned in Sections 5.1 and 5.2, this oceanic arc in the BNO has grown significantly. The Al-in-hornblende barometer shows that the 164 Ma Lameila pluton in western Tibet intrudes in the upper plate of the Jurassic oceanic subduction system within the BNO at a depth of about 13 km. This emplacement depth is the minimum crustal thickness of the Jurassic oceanic arc within the BNO, since the thickness of the crust underneath the emplacement level of the Lameila pluton is unknown. This depth of emplacement implies that the Jurassic oceanic arc or the oceanic arc with shift and the initial development of this oceanic arc on the oceanic crust of

the BNO. According to studies on both modern and paleo-oceanic arcs around the world, arc crust thicknesses depend on the thickness of pre-arc basement, tectonic extension or shortening, and arc maturity (magmatic addition and evolution) [67]. The pre-arc basement crust consists of mainly mafic oceanic crust and minor sedimentary cover. Studies on the ophiolites show that the composition of the oceanic crust is dominated by regular N-MORB and E-MORB type mafic rocks, and that oceanic crust related to OIB (oceanic island basalt) or oceanic plateaus are not recognized in the Shiquanhe area [10,68]. Therefore, the thickness of pre-arc oceanic crust is most likely comparable to that of the average oceanic crust (6.5 km) [69]. Rocks with levels higher than the Lameila pluton have been removed, thus much of the Lameila pluton is exposed at the surface (Figure 4a), and the outcrops of sedimentary rocks with pre-arc ages are extremely limited in the vicinity of the Lameila pluton and the Shiquanhe high-Mg andesite (Figure 3). Therefore, the pre-arc sedimentary covers on the floor of the BNO may be quite thin and thus their contribution to pre-arc crustal thickness is negligible. The effect of tectonic shortening on the thickening of the Jurassic oceanic arc in the BNO could not be quantitatively evaluated in this work due to the very limited preservation of structural traces.

Alternatively, magmatism may play a considerable role in the thickening of the Jurassic oceanic arc within the BNO. Studies on modern oceanic arcs, such as the Izu-Bonin-Marina [70–73], the Lesser Antilles [74–76], the Tonga-Kermadec [77–79], and the Aleutian [80–82] arcs show a variation of crustal thickness (9–35 km) within a single arc or between different arcs. In addition, an important finding of these studies is that the thick part of an arc is always related to high magmatic flux. This phenomenon implies that magmatic addition is a dominant mechanism responsible for oceanic arc crustal thickening. The lithological assemblage of the Jurassic oceanic arc in the BNS comprises only magmatic rocks, such as the Lameila quartz diorites, the Shiquanhe and Daruco high-Mg andesites, and the Dingqing high-Mg quartz gabbro/diabase. Pending additional studies of detailed tectonic evolution, we tentatively propose the crustal thickening of the Jurassic oceanic arc in the BNO to be a result of magmatic addition.

5.4. Oceanic Arc Development and Two Contrasting Sources for Magmatism

The east–west trending BNS in the Tibetan Plateau is a belt more than 1200 km long consisting of scattered ophiolitic fragments that are associated with thick sequences of Jurassic flysch. The BNS is extraordinarily wide, especially near Amdo in eastern Tibet and between Rutog and Shiquanhe in western Tibet (Figure 1). Besides the ophiolitic fragments and the Jurassic flysch, there are also micro-continental blocks such as the Amdo micro-block, and possibly oceanic arc blocks within the BNS. The mechanism incorporating these micro-blocks into the BNS is unclear.

In contrast to the well-exposed Amdo micro-block, surface traces of the oceanic arc rocks in the BNS are highly obscure. Previous recognition of BNS oceanic arc rocks has been mainly based on geochemical evidence from the volcanic rocks that are associated with ophiolitic fragments (SSZ type ophiolite), and suggests that an oceanic subduction occurred within the BNO [10,31,83–85]. Importantly, the discoveries of boninitic rocks near Dingqing [17], Daruco [20], Shiquanhe [19], and Rutog [18] in the BNS provide further evidence of the existence of an oceanic arc within the BNO. However, the extremely localized outcrops of these volcanic rocks make it hard to trace them in the field, and thus hamper further studies on the development and configuration of this arc. The new data from the Lameila pluton of this study have three implications for the tectonic evolution of the BNO. Firstly, this study recognizes the intrusive part of the oceanic arc, and in combination with the previously reported extrusive rocks, further confirms the existence of a Jurassic oceanic arc within the BNO. Secondly, the emplacement depth of the Lameila pluton suggests that the crust of the oceanic arc within the BNO was significantly thickened during the end of the middle Jurassic. Thirdly, the zircon Lu–Hf isotopes of the Lameila pluton, together with that of the previously reported high-Mg andesites within the BNS, imply diverse sources for the oceanic arc magmatism as shown below.

Studies on the high-Mg and esites from Shiquanhe (Figure 3) and Daruco (Figure 1) show arc-type geochemical compositions and enriched Nd ($\varepsilon_{Nd}(t) = -8.7$ to -13.4) and Hf ($\varepsilon_{Hf}(t) = -7.4$ to -16.2)

isotopes. It has been proposed that these high-Mg andesites are formed by the melting of subducted sediments (Figure 12) and subsequent interaction with overlying mantle peridotites during subduction initiation [19,20]. In contrast with the enriched isotopes of the above high-Mg andesites, zircon Lu–Hf isotopes of the Lameila pluton were depleted, with $\varepsilon_{Hf}(t)$ ranging from 10.5 to 13.9. The depleted Lu–Hf isotopes of the Lameila pluton clearly indicate that they were formed by the partial melting of a depleted-mantle source, most likely the depleted-mantle wedge peridotites above the oceanic subduction zone in the BNO (Figure 12). Synchronous with this oceanic arc in the BNO, a Jurassic continental arc (Figure 1) in the south margin of the Qiangtang block has also developed, as suggested by many previous researchers [13,15,16,29,38,39]. These two arcs suggest the occurrence of two subduction zones during the Middle Jurassic, namely, an oceanic–continental subduction zone between the BNO and the Qiangtang block in the north, and an oceanic subduction zone within the BNO in the south (Figure 1) were accreted onto the southern margin of the Qiangtang block during the continental block (Figure 1) were accreted onto the southern margin of the Qiangtang block during the continental subduction of the BNO, or during the Lhasa-Qiangtang collision. The accretion of the oceanic arc and micro-block could lead to a wide Bangong–Nujiang suture in central Tibet (Figure 2d).

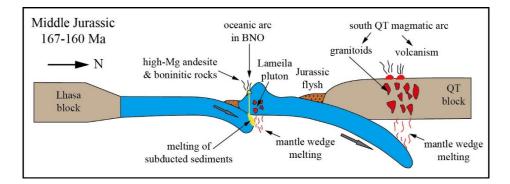


Figure 12. Tectonic model showing the Middle Jurassic Bangong–Nujiang Ocean (BNO) subduction, including oceanic subduction within the BNO and the subduction of the BNO beneath the south margin of the Qiangtang (QT) block. Associated with these two subduction zones are two arcs, one oceanic arc developed within the BNO and the other formed in the south margin of the Qiangtang block.

6. Conclusions

This work reports a study of the plutonic part of the Jurassic oceanic arc within the Bangong–Nujiang suture in western Tibet. The Lameila pluton consists mainly of quartz diorites with a zircon U–Pb age of 164 Ma. In addition, the zircon $\varepsilon_{Hf}(t)$ values of 10.5–13.9 indicate that the Lameila pluton was most likely sourced from the depleted-mantle wedge. Temperature-corrected Al-in-hornblende pressure was 3.9 ± 0.8 kbar, and the corresponding depth of emplacement was 13 ± 3 km. The thickness of the Jurassic oceanic arc crust must have doubled since the initial growth of the oceanic arc on the Bangong–Nujiang Ocean crust, with a thickness of 6.5 km during the Middle Jurassic.

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