

Supplementary Materials

S1

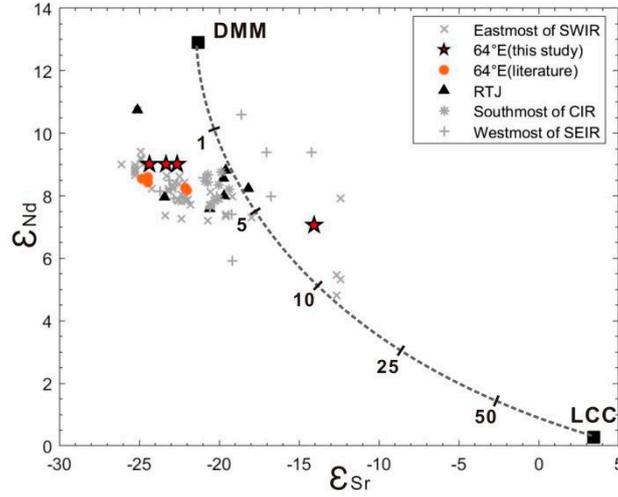


Figure S1. Plots of ϵ_{Sr} vs. ϵ_{Nd} showing data for DMM and LCC end-members. Data sources are the same as in Figure 5. Isotopic data sources for the depleted MORB mantle (DMM) are from [1,2] and for the lower continental crust (LCC) composition they are from [3,4]. From this figure, we considered that the DMM was variously contaminated by the LCC. Furthermore, most of samples were below the trend line, which implies the existence of a third component.

PCA S1. Principal Component Analysis (PCA)

The MORBs from 64°E are proposed to represent a ternary mixing among a DMM end-member, an EM2 end-member, and an LCC end member, and similar observations were made for the E-SWIR, RTJ, S-CIR, and W-SEIR basalts. In order to test these hypotheses, a principal component analysis (PCA) was performed on the Pb isotope data of the 64°E samples (this study and data from the literature).

As only three end-members are involved in that mixing processes, the data should plot on a single plane in the three-dimensional data ($^{206}\text{Pb}/^{204}\text{Pb}$ vs. $^{207}\text{Pb}/^{204}\text{Pb}$ vs. $^{208}\text{Pb}/^{204}\text{Pb}$). Furthermore, this plane is supposed to pass through the three end-members. This means that no more than two vectors (principal components) are necessary to explain the selected data.

To perform the PCA, the covariance matrix is calculated on the reduced data:

$$\text{cov}(data) = \begin{bmatrix} 0.0249 & 0.0031 & 0.0249 \\ 0.0031 & 0.0026 & 0.0074 \\ 0.0249 & 0.0074 & 0.0380 \end{bmatrix} \quad (1)$$

This matrix has three eigenvalues and three associated eigenvectors:

$$\lambda_1 = 0.0584; \quad \lambda_2 = 0.0065; \quad \lambda_3 = 0.0007 \quad (2)$$

$$\vec{v}_1 = \begin{pmatrix} 0.6002 \\ 0.1386 \\ 0.7878 \end{pmatrix}; \quad \vec{v}_2 = \begin{pmatrix} -0.7682 \\ 0.3743 \\ 0.5194 \end{pmatrix}; \quad \vec{v}_3 = \begin{pmatrix} 0.2229 \\ 0.9169 \\ -0.3311 \end{pmatrix} \quad (3)$$

Those three eigenvectors define the three-dimensional coordinate system in which our data can be plotted (Figure 10a).

The data variability is mainly expressed by the first two principal components (98.9%), thus showing the extent of inhomogeneity of the mantle: the mantle source

gradually undergoes binary or ternary mixing from the 61°E of SWIR eastward to the RTJ.

The second covariance matrix is

$$\text{cov}(data) = \begin{bmatrix} 0.0368 & 0.0029 & 0.0359 \\ 0.0029 & 0.0015 & 0.0046 \\ 0.0359 & 0.0046 & 0.0450 \end{bmatrix} \quad (4)$$

The associated eigenvalues and eigenvectors are

$$\lambda_1 = 0.0774; \quad \lambda_2 = 0.0050; \quad \lambda_3 = 0.0009 \quad (5)$$

$$\vec{v}_1 = \begin{pmatrix} 0.6635 \\ 0.0704 \\ 0.7448 \end{pmatrix}; \quad \vec{v}_2 = \begin{pmatrix} 0.7385 \\ -0.2212 \\ -0.6370 \end{pmatrix}; \quad \vec{v}_3 = \begin{pmatrix} -0.1199 \\ -0.9727 \\ 0.1988 \end{pmatrix} \quad (6)$$

When the Pb isotope data from E-SWIR, RTJ, S-CIR, and W-SEIR are considered (Figure 10b), the first two principal components explain 98.9% of the total dispersion. The remaining 1.1% may be explained by the analytic uncertainties.

The third covariance matrix calculated on the reduced data for these data is

$$\text{cov}(data) = \begin{bmatrix} 0.0538 & 0.0046 & 0.0457 \\ 0.0046 & 0.0017 & 0.0057 \\ 0.0457 & 0.0057 & 0.0770 \end{bmatrix} \quad (7)$$

The three eigenvalues and associated eigenvectors are

$$\lambda_1 = 0.1130; \quad \lambda_2 = 0.0183; \quad \lambda_3 = 0.0013 \quad (8)$$

$$\vec{v}_1 = \begin{pmatrix} 0.6126 \\ 0.0652 \\ 0.7877 \end{pmatrix}; \quad \vec{v}_2 = \begin{pmatrix} 0.7891 \\ 0.0074 \\ -0.6143 \end{pmatrix}; \quad \vec{v}_3 = \begin{pmatrix} 0.0459 \\ -0.9978 \\ 0.0469 \end{pmatrix} \quad (9)$$

The data all plot on a plane described by the eigenvectors v_1 and v_2 . This is in agreement with what was expected for a ternary mixing process. Data may now be represented on a two-dimensional graph (Figure 11). This and the triangular shape of the data in the figure confirm that only three end-members (DMM, LCC, and EM2) are necessary to explain the dispersion of the data in the Pb space. Figure 11 also shows that the mantle source beneath E-SWIR, RTJ, S-CIR, and W-SEIR have a certain similarity.

References

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