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\mathcal{L}_p -Norm-like Affine Projection Sign Algorithm for Sparse System to Ensure Robustness against Impulsive Noise

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Abstract: An improved affine projection sign algorithm (APSA) was developed herein using a \mathcal{L}_p norm-like constraint to increase the convergence rate in sparse systems. The proposed APSA is robust against impulsive noise because APSA-type algorithms are generally based on the \mathcal{L}_1 -norm minimization of error signals. Moreover, the proposed algorithm can enhance the filter performance in terms of the convergence rate due to the implementation of the \mathcal{L}_p -norm-like constraint in sparse systems. Since a novel cost function of the proposed APSA was designed for maintaining the similar form of the original APSA, these have symmetric properties. According to the simulation results, the proposed APSA effectively enhances the filter performance in terms of the convergence rate of sparse system identification in the presence of impulsive noises compared to that achieved using the existing APSA-type algorithms.

Keywords: adaptive filter; affine projection sign algorithm; sparse system; impulsive noise; convergence rate; filter performance

1. Introduction

The adaptive filtering theory has been widely applied in several domains such as system identification, channel estimation, as well as noise and echo cancellation [1–6]. As shown in Figure 1, the main purpose of an adaptive filter is to accomplish the precise estimation of filter coefficients to minimize error signals with the same input signals. Representative adaptive filtering algorithms include the least-mean-squares (LMS) algorithm and normalized LMS algorithm, which have a low computational complexity and can be easily implemented. In addition, the affine projection algorithm (APA) [7–9] has been developed to enhance the convergence performance associated with correlated input signals. However, because LMS-type and APA-type algorithms are based on the \mathcal{L}_2 -norm optimization of error signals, their performance deteriorates in the presence of system output noise, which includes impulsive noise.

To ensure satisfactory filter performance even in the presence of impulsive noise, several adaptive filtering algorithms derived from \mathcal{L}_1 -norm optimization have recently been proposed [10–13]. A representative algorithm that is robust against impulsive noises is the affine projection sign algorithm (APSA) [11]. However, the APSA is not particularly useful in sparse system identification as network echo cancellation or underwater acoustic channel estimation, in which the impulse response is primarily composed of near-zero coefficients and only a few large coefficients. In this context, many adaptive filters have been proposed to optimize filter performance by considering the system sparsity [12–17].



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Copyright: © 2021 by the authors. Licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (https:// creativecommons.org/licenses/by/ 4.0/). Among such frameworks, adaptive filtering algorithms based on the concept of an \mathcal{L}_0 -norm constraint have been considered to enhance the convergence rate of sparse systems [13,14]. In accordance with the \mathcal{L}_0 -norm constraint, a \mathcal{L}_p -norm-like concept has been introduced to enhance the filter performance of LMS-type algorithms [18–20] and it has been demonstrated that the \mathcal{L}_p -norm-like concept can contribute to the performance enhancement of adaptive filters in sparse systems. However, due to the inherent characteristics of LMS-type algorithms, their performance remains unsatisfactory in cases involving impulsive noise.

Considering these aspects, this paper reports on an improved APSA for sparse system identification. The algorithm is based on a novel cost function derived from \mathcal{L}_1 -norm optimization including the \mathcal{L}_p -norm-like constraint. Unlike the original APSA, the proposed cost function uses the a priori error instead of the a posteriori error. Consequently, the proposed APSA does not employ the approximation of the a posteriori error with the a priori error while formulating the equation to update the filter coefficient vectors. The performance of the proposed APSA in a general sparse system was evaluated and compared with those of the APSA [11], real-coefficient proportionate APSA (RP-APSA) [12], real-coefficient improved proportionate APSA (RIP-APSA)[12] and \mathcal{L}_0 -APSA [13].

This paper is organized as follows. In the following Section 2, the original APSA is described. In Section 3, the proposed \mathcal{L}_p -norm-like APSA is explained in detail. In Section 4, we present the simulation results to verify the performance of the proposed APSA. Finally, Section 5 gives the conclusions of this paper.



Figure 1. Structure of an adaptive filter in the presence of impulsive noise *n*_{*i*}.

2. Original APSA

The data vector \mathbf{d}_i from an unknown target system is defined as

$$d_i = \mathbf{u}_i^T \mathbf{w}_o + v_i,\tag{1}$$

where \mathbf{w}_o is the n-dimensional column vector to be estimated, v_i indicates the measurement noise with a variance σ_v^2 , and the input vector $\mathbf{u}_i = [u_i \ u_{i-1} \ \cdots \ u_{i-n+1}]^T$. The output error vector is denoted by \mathbf{e}_i , where the desired output data vector is denoted by \mathbf{d}_i , the data matrix is denoted by \mathbf{U}_i , and $\hat{\mathbf{w}}_i$ is the estimate of \mathbf{w}_o at iteration *i*, as follows:

$$\mathbf{e}_i = \mathbf{d}_i - \mathbf{U}_i^T \widehat{\mathbf{w}}_i,\tag{2}$$

$$\mathbf{d}_i = \begin{bmatrix} d_i & d_{i-1} & \cdots & d_{i-M+1} \end{bmatrix}^T. \tag{3}$$

$$\mathbf{U}_i = [\mathbf{u}_i \ \mathbf{u}_{i-1} \ \cdots \ \mathbf{u}_{i-M+1}], \tag{4}$$

$$\widehat{\mathbf{w}}_i = [\widehat{w}_i(0), \dots, \widehat{w}_i(n-1)]^T.$$
(5)

The original APSA [11] is derived by minimizing the \mathcal{L}_1 -norm of the a posteriori error vector with a constraint on the filter coefficient vectors, as shown below:

$$\min_{\widehat{\mathbf{w}}_{i+1}} ||\mathbf{d}_i - \mathbf{U}_i^T \widehat{\mathbf{w}}_{i+1}||_1 \text{ subject to } ||\widehat{\mathbf{w}}_{i+1} - \widehat{\mathbf{w}}_i||_2^2 \le \mu^2.$$
(6)

where μ^2 ensures that the filter coefficient vectors do not abruptly change. Using the gradient descent method, the filter coefficient vector of the original APSA is recursively defined using the following updating Equation [11]:

$$\widehat{\mathbf{w}}_{i+1} = \widehat{\mathbf{w}}_i + \mu \frac{\mathbf{U}_i \operatorname{sgn}(\mathbf{e}_i)}{\sqrt{\operatorname{sgn}(\mathbf{e}_i^T) \mathbf{U}_i^T \mathbf{U}_i \operatorname{sgn}(\mathbf{e}_i)}},\tag{7}$$

where μ represents the step size, $sgn(\cdot)$ is the sign function and $sgn(\mathbf{e}_i) = [sgn(e_i), \dots, sgn(e_i - M + 1)]^T$.

3. Proposed \mathcal{L}_p -Norm-like APSA

The proposed affine projection sign algorithm is newly formulated by minimizing the \mathcal{L}_1 -norm of the *apriori* error vector with the \mathcal{L}_p -norm-like constraint [18] as follows:

$$\min_{\widehat{\mathbf{w}}_i} \quad ||\mathbf{d}_i - \mathbf{U}_i^T \widehat{\mathbf{w}}_i||_1 + \gamma ||\widehat{\mathbf{w}}_i||_p^p, \tag{8}$$

where $|| \cdot ||_p^p$ denotes the \mathcal{L}_p -norm-like entity that exerts zero attraction, p denotes the order of the norm, and γ is used to adjust the effect of the \mathcal{L}_p -norm-like constraint. The proposed cost function can be derived from (8) as

$$J(\widehat{\mathbf{w}}_i) = ||\mathbf{e}_i||_1 + \gamma ||\widehat{\mathbf{w}}_i||_p^p.$$
(9)

The derivative of the proposed cost function (9) with respect to the filter coefficient vector $\hat{\mathbf{w}}_i$ is as follows:

$$\nabla_{\widehat{\mathbf{w}}_{i}} J(\widehat{\mathbf{w}}_{i}) = \frac{\partial J(\widehat{\mathbf{w}}_{i})}{\partial \widehat{\mathbf{w}}_{i}} = -\mathbf{U}_{i} \operatorname{sgn}(\mathbf{e}_{i}) + \gamma \frac{\partial ||\widehat{\mathbf{w}}_{i}||_{p}^{p}}{\partial \widehat{\mathbf{w}}_{i}}$$
$$\triangleq -\mathbf{U}_{i} \operatorname{sgn}(\mathbf{e}_{i}) + \gamma \mathbf{f}(\widehat{\mathbf{w}}_{i}), \tag{10}$$

where $\mathbf{f}(\widehat{\mathbf{w}}_i) \triangleq [f(\widehat{w}_i(0)), \dots, f(\widehat{w}_i(n-1))]^T$. The f -porm-like entity is widely defined.

The \mathcal{L}_p -norm-like entity is widely defined as

$$||\widehat{\mathbf{w}}_{i}||_{p}^{p} = \sum_{k=0}^{n-1} |\widehat{w}_{i}(k)|^{p}, \qquad 0 \le p \le 1.$$
(11)

The derivative of (11) with respect to the filter coefficient vector can be expressed in a component-wise manner as

$$f(\widehat{w}_{i}(k)) = \frac{\partial ||\widehat{\mathbf{w}}_{i}||_{p}^{p}}{\partial \widehat{w}_{i}(k)}$$
$$= \frac{p \operatorname{sgn}(\widehat{w}_{i}(k))}{|\widehat{w}_{i}(k)|^{1-p} + \epsilon}, \quad \forall 0 \le k < n,$$
(12)

where ϵ is an extremely small positive parameter that is introduced to avoid division by zero.

Moreover, the normalized gradient descent method is used to modify the updating equation of the filter coefficient vectors derived from the proposed cost function:

$$\widehat{\mathbf{w}}_{i+1} = \widehat{\mathbf{w}}_i - \mu \frac{\nabla_{\widehat{\mathbf{w}}_i} J(\widehat{\mathbf{w}}_i)}{||\nabla_{\widehat{\mathbf{w}}_i} J(\widehat{\mathbf{w}}_i)||_2} = \widehat{\mathbf{w}}_i + \mu \frac{(\mathbf{U}_i \operatorname{sgn}(\mathbf{e}_i) - \gamma \mathbf{f}(\widehat{\mathbf{w}}_i))}{\sqrt{(\mathbf{U}_i \operatorname{sgn}(\mathbf{e}_i) - \gamma \mathbf{f}(\widehat{\mathbf{w}}_i))^T (\mathbf{U}_i \operatorname{sgn}(\mathbf{e}_i) - \gamma \mathbf{f}(\widehat{\mathbf{w}}_i))}},$$
(13)

where μ is a step-size parameter of the proposed APSA. As can be seen in (7) and (13), the original APSA and proposed APSA have a symmetric relationship to maintain their robustness against impulsive noise. Moreover, owing to the \mathcal{L}_p -norm-like constraint, the proposed APSA has the zero-attraction property for increasing the convergence rate in the sparse system.

4. Simulation Results

The filter performance of the proposed APSA was evaluated through computer simulations involving system-identification scenarios. An unknown target system was randomly generated with 128 taps (n = 128) and the adaptive filter and unknown target system were considered to have the same number of taps. Moreover, we set 96 near-zero filter coefficients among the 128 taps to establish a general sparse system. The general sparse system has the property of sparsity as shown in Figure 2.



Figure 2. Characteristics of a general sparse system (*n* = 128).

Each adaptive filter was tested for a projection order M = 4. In our simulations, three types of input signal—white input, autoregressive (AR) and autoregressive–moving-average (ARMA)—were used, and correlated input signals for AR and ARMA models were generated by filtering white Gaussian noise through the following systems as

$$G_1(z) = \frac{1}{1 - 0.7z^{-1}},\tag{14}$$

$$G_2(z) = \frac{1 + 0.6z^{-1}}{1 + z^{-1} + 0.21z^{-2}}.$$
(15)

The signal-to-noise ratio (SNR) was set to 30 dB for adding the measurement noise at the output signal y_i . The SNR is defined by $y_i = \mathbf{u}_i^T \mathbf{w}_o$ as follows:

$$SNR \triangleq 10 \log_{10} \left(\frac{E[y_i^2]}{E[v_i^2]} \right).$$
(16)

The impulsive noise n_i was generated as $n_i = k_i A_i$ where k_i represents a Bernoulli process with a probability of success $P[k_i = 1] = Pr$, and A_i represents zero-mean Gaussian noise with the power $\sigma_A^2 = 1000\sigma_y^2$ [11,13]. *Pr*, which denotes the probability of the occurrence of impulsive noise, was set to 0.001. The normalized mean squared deviation (NMSD) was defined as

$$\text{NMSD} \triangleq 10 \log_{10} \left(\frac{E[\widetilde{\mathbf{w}}^{T}(k)\widetilde{\mathbf{w}}(k)]}{\mathbf{w}_{o}^{T}\mathbf{w}_{o}} \right), \tag{17}$$

where the filter-coefficient error vector $\tilde{\mathbf{w}}_i \triangleq \mathbf{w}_o - \hat{\mathbf{w}}_i$. The results were obtained via ensemble averaging over 100 trials.

4.1. System Identification for Sparse System in Presence of Impulsive Noises

Figure 3 shows the NMSD learning curves for the original APSA [11], RP-APSA [12], RIP-APSA [12], \mathcal{L}_0 -APSA [13] and proposed APSA with white input signals. Figures 4 and 5 also indicate the NMSD learning curves for the original APSA, RP-APSA, RIP-APSA, \mathcal{L}_0 -APSA and the proposed APSA with correlated input signals $G_1(z)$ and $G_2(z)$, respectively. According to Figures 3–5, the proposed \mathcal{L}_p -norm-like APSA has a faster convergence rate than the existing algorithms when the representative white, AR and ARMA input signals were used. Because the convergence rate and steady-state estimation errors have a trade-off relationship, we focus on the comparison in aspects of the convergence rate. To ensure a fair comparison of the convergence rate of the proposed APSA and other algorithms, the algorithm parameters were selected such that the steady-state errors for all the algorithms were identical. Specifically, the parameters were set as follows: APSA ($\mu = 0.005$); RP-APSA ($\mu = 0.007$, p = 0.1); RIP-APSA ($\mu = 0.009$, $\alpha = 0$); \mathcal{L}_0 -APSA ($\mu = 0.022$, $\beta = 20$, $\gamma = 0.003$); and the proposed APSA $(\mu = 0.017$ for the white input/ $\mu = 0.02$ for $G_1(z)$, $G_2(z)$ and speech input, $\gamma = 0.03$, p = 0.1, $\epsilon = 0.01$). To set almost the same convergence rate for all algorithms, the step sizes of the proposed algorithm for the white input are slightly slight difference compared to those of the AR and ARMA cases. The p parameter of RP-APSA and the α parameter of RIP-APSA were set with reference to an existing study [12] for fair comparison. As shown in Figures 1 and 2, the proposed APSA exhibited a higher convergence rate than those of the other algorithms. In addition, the proposed APSA maintains filter performance in terms of the convergence rate even though the system change suddenly occurs—as shown in Figures 6–8.



Figure 3. NMSD learning curves for the white input in a sparse system with impulsive noises (Pr = 0.01).



Figure 4. NMSD learning curves for the correlated input generated using $G_1(z)$ in a sparse system with impulsive noises (Pr = 0.01).



Figure 5. NMSD learning curves for the correlated input generated using $G_2(z)$ in a sparse system with impulsive noises (Pr = 0.01).



Figure 6. NMSD learning curves for the white input in a sparse system with impulsive noises (Pr = 0.01). The system suddenly changes ($\mathbf{w}_o \rightarrow -\mathbf{w}_o$) at iteration 3.8 × 10³).



Figure 7. NMSD learning curves for the correlated input generated using $G_1(z)$ in a sparse system with impulsive noises (Pr = 0.01). The system suddenly changes ($\mathbf{w}_o \rightarrow -\mathbf{w}_o$) at iteration 3.8 × 10³.





4.2. Speech Input Test Including a Double-Talk Situation

The proposed APSA was subjected to an experiment using a speech input signal to ensure the filter performance in practical scenarios, as can be seen in Figure 9. Since speech input signals are real human speech data, this simulation result for speech input increases the reliability of the proposed APSA for the practical use. As can be seen in Figure 10, we can find that the proposed APSA can accomplish a faster convergence rate and smaller steady-state estimation errors compared to the other algorithms. The proposed algorithm was also tested in a double-talk situation as shown in Figure 11. The far-end input signal and near-end input signal were speech signals where the power of the near-end input signal was two times greater than that of the far-end input signal. The near-end input signal was added between iterations 5.2×10^3 and 6.2×10^3 . Figure 11 shows that the proposed APSA delivered better performance than the other algorithms in terms of the convergence rate and the steady-state estimation error. Even after the double-talk occurrence, we found that the proposed APSA consistently has smaller steady-state estimation errors.



Figure 9. Speech input signal.



Figure 10. NMSD learning curves for the speech input.



Figure 11. NMSD learning curves for the double-talk situation.

4.3. Practical Considerations for the p Parameter

Through Equation (19), we can analyze that when the p parameter is close to 0, the proposed \mathcal{L}_p -norm-like APSA is the same as the \mathcal{L}_0 -norm APSA. Moreover, the better filter performance of the proposed APSA compared to the \mathcal{L}_0 -norm APSA is demonstrated via Figures 3–8 and Figures 10 and 11. On the other hand, if the p parameter is close to 1 as can be seen in Equation (20), the proposed \mathcal{L}_p -norm-like APSA is similar to the \mathcal{L}_1 -norm APSA that does not have the advantages for the sparse system due to the characteristic of the \mathcal{L}_1 norm. Therefore, it is meaningful to choose the specific p value between 0 and 1 to improve the filter performance of the proposed \mathcal{L}_p -norm-like concept. We find that

the specific *p* value can maximize the improvement of our proposed algorithm through the parameter tuning method, as shown in Figures 12–14. As shown in Figures 12–14, the specific *p* value consistently verifies the improved filter performance regardless of the white, ar and arma input signals. When the *p* parameter is set to 0.1, the proposed APSA has the best filter performance, achieving a fast convergence rate and small steady-state estimation errors. Even though the parameter tuning method cannot precisely provide optimal value, p = 0.1 gives the best performance in our simulation scenarios. Since the parameter *p* value has a restriction between 0 and 1, it is easy to find the proper *p* value according to the variation of *p* values to adequately guarantee the improved filter performance in each scenario.

It is important to decide the *p* parameter in the proposed APSA because the effect of the *p* value is dominant in adjusting the filter performance in terms of convergence rate and steady-state estimation errors. The \mathcal{L}_p -norm-like definition can be represented by

$$\|\widehat{\mathbf{w}}_{i}\|_{p}^{p} = \sum_{k=0}^{n-1} |\widehat{w}_{i}(k)|^{p}, \qquad 0 \le p \le 1.$$
(18)

From Equation (18), we can derive the \mathcal{L}_0 -norm term and the \mathcal{L}_1 -norm term by adjusting the *p* value as follows:

$$\lim_{p \to 0} ||\widehat{\mathbf{w}}_i||_p^p = ||\widehat{\mathbf{w}}_i||_0 = \#\{k | \widehat{w}_i(k) \neq 0\},\tag{19}$$

$$\lim_{p \to 1} ||\widehat{\mathbf{w}}_i||_p^p = ||\widehat{\mathbf{w}}_i||_1 = \sum_{k=0}^{n-1} |\widehat{w}_i(k)|.$$
(20)



Figure 12. NMSD learning curves with several values of *p* to decide the *p* value to ensure the best filter performance for the white input in a sparse system with impulsive noises (Pr = 0.01).



Figure 13. NMSD learning curves with several values of *p* to decide the *p* value to ensure the best filter performance for the correlated input generated using $G_1(z)$ in a sparse system with impulsive noises (Pr = 0.01).



Figure 14. NMSD learning curves with several values of *p* to decide the *p* value to ensure the best filter performance for the correlated input generated using $G_2(z)$ in a sparse system with impulsive noises (Pr = 0.01).

5. Conclusions

An improved affine projection sign algorithm for sparse system identification was developed using a \mathcal{L}_p -norm-like constraint. The \mathcal{L}_p -norm-like constraint could accelerate the convergence of the near-zero coefficients of a general sparse system. Moreover, a novel cost function including the \mathcal{L}_p -norm-like constraint was used for the proposed APSA. Consequently, the convergence rate of the proposed APSA was higher than those of the original APSA, RP-APSA, RIP-APSA and \mathcal{L}_0 -APSA, as demonstrated by simulation results.

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