



Article The Influence of Instrumental Line Shape Degradation on Gas Retrievals and Observation of Greenhouse Gases in Maoming, China

Dandan Liu^{1,*}, Yinbo Huang², Zhensong Cao², Xingji Lu² and Xiangyuan Liu¹

- ¹ College of Electrical and Optoelectronic Engineering, West Anhui University, Luan 237012, China; baobaoldd@126.com
- ² Key Laboratory of Atmospheric Optics, Anhui Institute of Optics and Fine Mechanics, Chinese Academy of Sciences, Hefei 230031, China; ybhuang@aiofm.ac.cn (Y.H.); zscao@aiofm.ac.cn (Z.C.); lxj168@mail.ustc.edu.cn (X.L.)
- * Correspondence: dandan86@mail.ustc.edu.cn; Tel.: +86-18860465860

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Copyright: © 2021 by the authors. Licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (https:// creativecommons.org/licenses/by/ 4.0/). Abstract: The instrument line shape (ILS), as a very important parameter, has a significant influence on the inversion of trace gas concentration. Different levels of ILS degradation for H₂O, CO₂, CH₄. and CO gases were investigated, and the influence of ILS on the inversion of column-averaged dry air mole fractions (DMFs) was assessed. Our results indicate that the averages of XH₂O, XCH₄, and XCO with modulation efficiency (ME) amplitude values have a positive correlation, the correlation coefficients are 0.9925, 0.9968, and 0.9981 respectively, whereas the relationship between the average of XCO₂ and ME is a negative correlation with 0.986 correlation coefficient. For a typical ILS degradation, a decrease of 5% in the modulation efficiency amplitude value results in the average of XCO₂ changing by 0.744%, XCH₄ and XH₂O are less sensitive species, with average values of -0.206% and -0.464%, whereas XCO shows the strongest intraday variability with an average value of -0.238%. However, with a decrease of 2% in the phase error (PE) value, the average of XCO changed by -0.150%, XCO₂ and XH₂O almost coincided with the same average value of -0.141%, whereas XCH₄ was the least sensitive species with an average value of -0.133%. At the same time, we measured the ILS for EM27/SUN spectrometers-the mean values of modulation efficiency amplitudes and phase errors were 0.9611 and 0.00593. Compared with standard values, the modulation efficiency amplitudes and the phase error deviations were 2.450% and 0.433%. During the observation period, the daily average of XCO₂ ranged from 415.09 to 421.78 ppm. XCH₄ ranged from 1.96 to 2.02 ppm with a mean of 1.982 ppm, and the daily average of XCO ranged from 0.118 to 0.157 ppm with a mean of 0.137 ppm. For the relationship between XCO₂ and XCH₄, the linear regression line shows a good correlation with the correlation coefficient $R^2 \ge 0.5$. Especially, for the correlation coefficient $R^2 = 0.82$ on 8 October, our studies found a weak correlation in the variation of CO₂ and CO during the observations. The correlation coefficient $R^2 \ge 0.5$ was only found on 30 September and 3 October. The trajectories dram at a height of 10 km give a hint of trace gas transport from the bay of western India, Bengal, and the Arabian Sea, whereas for the trajectories dram boundary layer height, trace gases were transported from southwest and east of China. These results provide a theoretical basis to understand the time and space distribution and the changes of greenhouse gas in the atmosphere as well as providing a theoretical basis for calculations of atmospheric radiation transmission.

Keywords: Fourier transform infrared spectroscopy; greenhouse gases; ILS; DMFs; HYSPLIT

1. Introduction

The continuing increase of atmospheric greenhouse gas is the main driver of global warming, particularly carbon dioxide (CO_2), methane (CH_4), and carbon monoxide (CO). In recent years, many countries have put great efforts into the observation of concentrations

of greenhouse gases, including total column concentration, column-averaged dry air mole fractions, and flux of greenhouse gases with global coverage. For example, the Greenhouse Gases Observing Satellite (GOSAT) launched by Japan and the Orbiting Carbon Obserbvatory-2 (OCO-2) satellite launched by the USA are tasked with long-term observations of global greenhouse gas concentrations [1-4]. The data derived from satelliteborne sensing have provided useful information but results still need to be validated. The ground-based Fourier transform spectrometer can accurately measure greenhouse gas concentrations by recording direct solar spectra for the validation of satellite-borne results. There are two well-known international networks based on ground-based solar FTIR instruments: the Total Carbon Column Observing Network and the Network for the Detection of Atmospheric Composition Change [5,6]. Both TCCON and NDACC networks have more than 20 sites around the word, and are dedicated to retrieval of total column concentration of CO2, CH4, CO, and other gases by ground-based high resolution Fourier transform infrared (FTIR) spectrometers (the type 125HR from Bruker). For example, Messerschmidit et al. (2011) claimed that the accuracy and precision of a column-averaged dry air mole fraction of CO_2 was better than 0.25% [7]. However, due to the type 125HR spectrometers applied for two networks not being portable, it is difficult for observation of sources and sinks on a regional scale. A portable low resolution FTIR spectrometer EM27/SUN has the advantages of lightweight, excellent stability, and lowcost use for the observation of greenhouse gases by recording direct solar spectra. The long-term performance of the EM27/SUN spectrometer with respect to a reference high resolution instrument was investigated. Namely, Gisi et al. (2012) found that the total column XCO₂ by EM27/SUN using solar absorption spectroscopy was in agreement with TCCON results with $(0.12 \pm 0.08)\%$ being achieved [8]. Hase et al. (2015), and Frey et al. (2015) combined several EM27/SUN spectrometers to estimate urban carbon dioxide and CH_4 emissions [9,10].

With the high precision requirements of remote sensing applications, atmospheric radiation transfer calculation, instrument model, and inversion algorithm are the main factors restricting atmospheric trace gas inversion. In the instrument model, the instrument line shape (ILS) is the most important parameter, which determines the degree of modulation of the instrument on the incident radiation, that is, the spectral resolution of the instrument and the spectral line energy distribution. The ILS can be divided into two parts. The first part describes the modulation loss caused by the inherent self-apodization of the instrument. It can be used to calculate the optical thickness OPD of the spectrometer and the field of view FOV. The second part is caused by the uncalibrated and optical phase difference of the instrument, characterized by modulation efficiency (ME) amplitude and phase error (PE). The modulation efficiency amplitude characterizes the width of ILS, and the phase error quantifies the degree of ILS symmetry. A standard procedure to derive the ILS of the TCCON and NADCC spectrometer are gas cell measurements, using a low-pressure calibration gas cell to diagnose a misalignment of the spectrometer. For example, the ILS of the TCCON spectrometer is a HCL gas cell. In contrast, a successful alignment scheme for low-resolution spectrometers was proposed, the ILS of a EM27/SUN spectrometer by measuring the H_2O spectrum in laboratory air to calibrate the ILS of the instrument. As light source a collimated standard 50 W halogen light bulb was used. Recently, researchers studied the effect of the ILS on gas inversion. Such as, the ILS modulation efficiency of the EM27/SUN spectrometer before and after the campaign with 0.24%, corresponding to a change of only 0.04% for the XCO₂ [8]. An increase of 1% in the modulation efficiency results in an XCO₂ increase of 0.15%, whereas XCH₄ increases 0.1% [8,11]. In the study of Frey 2018, the modulation efficiency at maximum optical path difference ranges between 0.9835 and 0.9896, with a mean value of 0.9862 and a standard deviation of 0.0015. The phase error was close to zero for the whole time series with a mean value of 0.0019 ± 0.0018 by studying ILS of thirty EM27/SUN.

In the first part of this study, we quantitatively calculated the impact of ILS degradation on XH₂O, XCO₂, XCH₄, and XCO and analyzed the average gas concentration changes

with modulation efficiency amplitude and phase error. We performed lab-air observations of water gas signatures for the determination of instrumental line shape characteristics and obtained the values of ME and PE. Moreover, the change trend of greenhouse gas concentration in the Maoming area was analyzed, while at the same time we analyzed the correlation between XCH₄, XCO, and XCO₂.

2. Measurement Sit and Instruments

Measurements of the trace gases were made using the EM27/SUN spectrometer for XH_2O , XCO_2 , XCH_4 , and XCO. The instruments were located at Bohe Marine Meteorological Observatory (21.453° N, 111.315° E, 0.02 km above sea level), adjacent to the South China Sea, located in the south Binhai New area of Maoming city (Figure 1 left). The coastline is very straight and has no terrain blocking, which is very conducive to the observation of various marine meteorological elements. We installed the instrument, consisting of the EM27/SUN spectrometer and solar tracker. A detailed description of the spectrometer can be found in Gisi et al. (2012), in the following we give only a short overview.



Figure 1. Observing site (Bohe Marine Meteorological Observatory, **left**). FTIR spectrometer (EM27/SUN, **right**).

To reach high stability with regard to thermal influences and vibrations, the EM27/SUN features a RockSolidTM pendulum interferometer with two cube corner mirrors and a CaF₂ beam splitter. The instrument achieves 1.8 cm optical path difference (OPD) with a maximum spectral resolution of 0.5 cm⁻¹, as shown in Figure 1 (right). Measurements are recorded with an InGaAs detector operated at ambient temperature. Due to spectral coverage from 5000 to 11,000 cm⁻¹, the spectral bandwidth enables the detection of O₂, H₂O, CO₂, CH₄, and CO. The detector signal is DC coupled and thereby supports the correction of variable atmospheric transmission [12].

3. Data Processing and Methods

This section provides a concise and precise description of the experimental results, their interpretation, as well as the experimental conclusions that can be drawn.

Atmospheric CO₂, CH₄, and CO are measured with a ground-based Fourier transform spectrometer that records the near infrared spectrum. We recorded double-sided interferograms with 0.5 cm^{-1} resolution. With 10 scans, one measurement takes about 58 s. In order to ensure data quality, a pre-processing is performed. To suppress the negative sidelobes, sometimes a numerical apodisation is applied. This is especially important for low resolution instruments. On the downside an apodisation decreases the spectral resolution and amplifies the correlation between measured values from different spectral positions. A good compromise between acceptable resolution, degradation, and suppressing the negative sidelobes is the Norton–Beer medium function [13]. Furthermore, a DC correction is performed. In addition, a quality filter discards interferograms with intensity fluctuations above 10% and intensities below 10% of the maximal modulation amplitude. To ensure high-precision greenhouse gas concentration, not only the altitude, latitude, and longitude, a priori temperature, humidity, pressure, and other parameters need to be considered, but also the real-time meteorological parameters of the site, such as surface temperature, surface pressure, and the instrumental characteristics of the measurement device (ILS), etc.

Precise knowledge of instrumental line shape (ILS) is of utmost importance to gain correct information from measurement signals. Instrumental line shape is the Fourier transform of the weighting applied to the interferogram. It consists of two parts and affects the accuracy of the final inversion results. Due to inherent self-apodization of the spectrometer, which is present also in an ideal instrument, one part describes the modulation loss. This contribution can easily be calculated utilizing the OPD and FOV of the spectrometer. The other part of the ILS results from misalignments and optical aberrations of the spectrometer and can be described by a modulation efficiency amplitude (ME) and a phase error (PE) as a function of the OPD. The theoretical ideal ILS, is a convolution of sinc and rectangular functions, representing the finite length of the interferogram and the finite circular field of view of the spectrometer [14], defined as:

$$Sinc(\sigma, L) = 2L \frac{\sin(2\pi\sigma L)}{2\pi\sigma L}$$
(1)

$$\operatorname{Rect}(\sigma, \sigma_0, \theta) = \begin{cases} \frac{2}{\sigma_0 \theta^2} & -0.5\sigma_0 \theta^2 \le \sigma \le 0\\ 0 & otherwise \end{cases}$$
(2)

$$ILS(\sigma, \sigma_0, L, \theta) = Sinc(\sigma, L) \times \operatorname{Rect}(\sigma, \sigma_0, \theta)$$
(3)

Here, σ is the wavenumber, σ_0 is the central wavenumber, *L* is the optical path difference, and θ is the angular radius of the circular internal FOV of the spectrometer. The ILS of a real spectrometer is equivalent to complex modulation efficiency in the interferogram, many problems lead to smooth variation of the complex modulation efficiency as for example, misalignments and optical aberrations of the spectrometer. The phase corrected interferogram generated by a spectral line is of the form [15]:

$$IFG(x) \sim Mod_amp(x).\cos(2\pi\sigma - Mod_phas(x))$$
(4)

where IFG(x) is the interferogram, $Mod_amp(x)$ is the modulation efficiency amplitude (ME), and $Mod_phas(x)$ is the phase error (PE). The spectral resolution is defined as:

$$\Delta v = 0.9/MOPD \tag{5}$$

MOPD means the maximum optical path difference.

In this work, we analyzed spectra utilizing the PROFFIT retrieval fitting algorithm [10], which is in wide use and has been validated for retrieving dry-air mole fractions (DMF) of trace gases [8,16–18]. The PROFFIT is a non-linear least-squares fitting algorithm. The atmospheric forward model is used to calculate synthetic spectra. We fitted the atmospheric spectra by scaling of a priori trace gas profiles with low resolution of the EM27/SUN as well as a priori vertical profiles for temperature, pressure, and water gas from the National Centers for Environmental Prediction (NCEP). Then an inverse method compares the synthetic spectra with the measured spectra. Typically, an inversion calculation starts with a forward model $\vec{F}(\vec{x})$, includes the instrumental characteristics of the measurement device and underlying physics which relates the measured quantities and the target variable. The approximation of the physics of the measurement process:

$$\vec{y} = \vec{F}(\vec{x}) + \vec{\varepsilon}_y \tag{6}$$

Here, \vec{y} represents the measurement with the corresponding measurement error \vec{e}_y , \vec{x} is state vector, include pressure, temperature, or the volume mixing ratio. The assumption of discrete values for these parameters, the linearization of the forward model:

$$\vec{F}(\vec{x}) = \vec{F}(\vec{x}_0) + \frac{\partial \vec{F}}{\partial \vec{x}}(\vec{x} - \vec{x}_0) + \vec{\varepsilon}_y$$

= $\vec{F}(\vec{x}_0) + \mathbf{K}(\vec{x} - \vec{x}_0) + \vec{\varepsilon}_y$ (7)

Here, \vec{x}_0 is the reference state, **K** is the Jacobian matrix. The fitting residual is defined as follows:

$$|\Delta y|^2 S_y^{-1} = |y_{meas} - y|^2 S_y^{-1} = (y_{meas} - y)^T S_y^{-1} (y_{meas} - y)$$
(8)

where y_{meas} is the measured spectrum, y is the synthetic spectrum, S_y^{-1} is the covariance matrix of inversion state parameters. The volume mixing ratio of O₂ in the atmospheric altitude range up to 100 km is nearly constant at approximately 0.2095. The calculation of the ratio of the target gas and the column amount of O₂ retrieved from the same spectrum to remove the effects of surface pressure variation, the column-averaged dry air mole fraction (DMF) is defined as (9), using the column abundance of O₂ as a reference to reduce the systematic errors:

$$X_{gas} = 0.2095 \times \frac{\text{Column}_{gas}}{\text{Column}_{O_2}} \tag{9}$$

where column $_{gas}$ and column O_2 are the column abundance of the retrieval gas and O_2 respectively. Details of the retrieval method are given in Inverse Methods for Atmospheric Sounding: Theory and Practice [18]. The spectra windows for retrieval of the column-averaged dry air mole fraction (DMF) O_2 , CH_4 , CO, and O_2 are listed in Table 1. To make the measurements comparable to WMO scale, in post processing in our work, the calibration factors applied for XCO_2 , XCH_4 , and XCO were 0.9869, 0.9898, and 0.925, respectively [11]. Figure 2 shows a typical measured atmospheric spectrum, the mode spectrum, and residuals.

Table 1. Spectral windows for retrieval of the column of CO₂, CH₄, and CO.

Gas	Spectral Windows [cm ⁻¹]	Interfering Molecule
CO ₂	6173.0-6390.0	H ₂ O, HDO, CH ₄
CH_4	5897.0-6145.0	H ₂ O
CO	4208.7-4318.8	CH_4 , H_2O , HDO
O ₂	7765.0-8005.0	H_2O , HF, CO_2



Figure 2. Typical measured atmospheric spectrum, mode spectrum, and residuals.

4. Results Measurements and Discussion

The ME amplitude is connected to the width of the ILS, while the PE quantifies the degree of ILS asymmetry. The modulation efficiency amplitude (ME) is unity and the phase error (PE) is zero with an ideal ILS. However, if the spectrometer is not calibrated, the ME amplitude and PE would deviate from unity and zero. These parameters have to be calculated from laboratory measurement.

4.1. Influence of Instrumental Line Shape on Greenhouse Gas Inversion

In order to quantitatively analyze the impact of ILS on inversion, we analyzed the concentration changes of XH₂O, XCO₂, XCH₄ and XCO, when the ME deviate from $\pm 5\%$, $\pm 10\%$, $\pm 15\%$, $\pm 20\%$ compared to the ideal value unity, and the PE deviate from $\pm 2\%$, $\pm 4\%$, $\pm 6\%$, $\pm 8\%$, $\pm 10\%$ compared to the ideal value zero. We took the retrievals with an ideal ILS as the reference. The difference is defined as:

$$D\% = \frac{X - X_{ref}}{X_{ref}} \times 100 \tag{10}$$

Taking the standard instrument as reference, the influence of the modulation efficiency amplitudes and phase errors on greenhouse gas inversion are shown in Figure 3. In general, the mean of XH_2O , XCH_4 , and XCO with ME values has a positive correlation. The correlation coefficients are 0.9925, 0.9968, and 0.9981 respectively. However, the relationship between the mean of XCO_2 and ME is the opposite, the mean of XCO_2 decreases as the ME loss increases. It has a negative correlation with ME loss value, the correlation coefficient is 0.986. This is opposed to prior studies reporting an increase of XCO₂ and decrease of XCH_4 for an increase of the modulation efficiency [19]. The reason may be that the total amount of the gas column depends on the airmass while the interference molecules in the inversion window of each gas molecule are different. This is in agreement with the findings from Frey, M dissertation (2018), who reported that XCO₂ decreases with increasing ILS, whereas XCH₄ increases [11]. The Xgas (XH₂O, XCO₂, XCH₄ and XCO)/ME loss slopes are 271.3, -59.33, 0.192, and 0.0069. The time series of relative difference of these species in terms of the total column are displayed in Figure 3c. A decrease of 5% in the ME value results in the highest amount of XCO_2 with an average value of 0.744%. While XCH_4 and XH_2O are the less sensitive species, with average values of -0.206% and -0.464%. Among all these species, XCO shows the strongest intraday variability with an average value of -0.238%, ranging from -0.088% to -0.831%. The absorption of CO with the selected spectral window is much weaker than the other gases because of the superimposition from the nearby strong interfering lines of CH₄ and H₂O.

The means of XH_2O , XCO_2 , XCH_4 and XCO with different PE loss have a similar tendency, a negative correlation with PE loss (Figure 3 left). The correlation coefficients are higher than 0.999. The Xgas (XH₂O, XCO₂, XCH₄ and XCO)/PE loss slopes are -4451.08, -295.22, -1.35, and -0.107. The time series of the relative difference of these species in terms of total column shows a similar intraday variability (Figure 4d). The highest XCO has an average value of -0.150% with a decrease of 2% in the PE value, while XCO₂ and XH_2O almost coincide with a same average value of -0.141%. Among all these species, XCH₄ are the least sensitive species with an average value of -0.133%. Figure 4a,b shows ΔXH_2O , ΔXCO_2 , ΔXCH_4 , and ΔXCO with ME and PE increasing or decreasing by the same value. The values of ΔXCO_2 are different when ME increases and decreases by the same value. Similarly, the values of ΔXCH_4 are different when ME increases and decreases by the same value. The reason for the asymmetry may be that the interference molecules in the retrieval band of each gas molecule are different, and the modulation degree of the instrument on the interferogram is also different, while ΔXH_2O and ΔXCO are little different when ME increases and decreases by the same value. However, the values of ΔXH_2O , ΔXCO_2 , ΔXCH_4 , and ΔXCO are almost the same when PE increases and decreases by the same value.



Figure 3. Influence of the modulation efficiency amplitudes and phase errors on XH₂O, XCO₂, XCH₄, and XCO with the reference standard instrument.

Table 2 shows the difference of XH₂O, XCO₂, XCH₄, and XCO with \pm 1% in modulation efficiency and \pm 0.01 in phase error. For a typical ILS degradation of 1% in modulation efficiency, the column-averaged dry air mole fraction of H₂O, CO₂, CH₄, and CO changed by 0.0492, 0.1151, 0.1042, 0.0523%, respectively, while an increase of 0.01in phase error leads to a decrease of 0.697% on XH₂O, 0.689% on XCO₂, 0.652% on XCH₄, and 0.737% on XCO.

Table 2. Difference of XH₂O, XCO₂, XCH₄, and XCO with $\pm 1\%$ in modulation efficiency and ± 0.01 in phase error.

	ΔXH_2O	ΔXCO_2	$\Delta X C H_4$	ΔΧCΟ	
$\Delta ME = 1\%$	0.0492%	0.1151%	0.1042%	0.0523%	
$\Delta ME = -1\%$	0.0368%	0.1707%	0.0863%	0.0452%	
$\Delta PE = 0.01$	0.697%	0.689%	0.652%	0.737%	
$\Delta PE = -0.01$	0.715%	0.729%	0.689%	0.775%	



Figure 4. The e difference of XH₂O, XCO₂, XCH₄, and XCO with ME loss (**a**) and PE loss (**b**) with standard instrument, time series of relative difference of XH₂O, XCO₂, XCH₄, and XCO due to a 5% decrease of the ME values (**c**) and 2‰ rad PE values (**d**).

4.2. Instrumental Line Shape Monitoring

For the EM27/SUN spectrometer, the standard procedure to derive the ILS are laboratory measurements. Several meters of lab air measurements using a collimated standard 50 W halogen light bulb as source (Figure 5 left) and a stabilized digital laboratory DC power (11 V) supply were used and the water vapor lines were evaluated in the spectral region between 7000 and 7400 cm⁻¹. As the water column inside the spectrometer could not be neglected, the instrument was vented. The ILS retrievals are performed using LIN-FIT 14.5 [15]. As the ILS characteristics were close to nominal, we used the two parameters ILS model.

Due to the heat of the lamp affecting a non-negligible section of the open path, the distance between instrument and lamp should not be chosen as too small. Furthermore, the image of the lamp on the field stop is evenly illuminated and exceeds the diameter of the field stop. The resulting ILS values are presented in Tables 3 and 4, the modulation efficiency (ME) amplitudes and phase errors (PE) are shown in Figure 5 (right). Due to the venting, the mixing ratio of H₂O inside the spectrometer is the same as outside, thus the plan of the simple analysis assumes a uniform path between lamp and detector. The modulation efficiency (ME) amplitudes and phase errors are relatively close at distances of 401, 519, and 605 cm. The mean value of the modulation efficiency (ME) amplitudes and phase errors (PE) are 0.9611and 0.00593. Compared with standard values (ME = 0.9855, PE = 0.0016), the modulation efficiency (ME) amplitude and the phase error (PE) deviations are 2.450% and 0.433%. The ILS results show the alignment and stability of the instrument over the whole period. For the trace gas retrieval, we used the mean value of the measurement.

(a) 0.98 **Aodulation efficiency** 0.96 0.94 0.92 0.90 600 400 450 0.0070 (b) 0.0065 error 0.0060 0.0055 0.0050 400 450 500 550 600 detector-lamp(cm)



Table 3. ME amplitudes and ph	ase errors (12 September 2018).
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Distance/cm	ME	PE
401	0.9572	0.00569
519	0.9658	0.00613
605	0.9602	0.00598
mean	0.9611	0.00593
Standard deviation	0.00437	$2.23681 imes 10^{-4}$

Table 4. ME amplitudes and phase errors (12 April 2019).

Distance/cm	ME	РЕ
396	0.9591	0.00504
496	0.9622	0.0048
596	0.963	0.00508
696	0.9657	0.00571
mean	0.9625	0.00516
Standard deviation	0.00272	$3.89081 imes 10^{-4}$

4.3. Variation of XCO₂, XCH₄ and XCO

The direct absorption spectra were collected under clear-sky weather conditions from 30th September to 8 October 2018. The time series of the column-averaged dry air mole fraction of CO_2 , CH_4 and CO were retrieved. Because of instrument mechanical failure or adverse weather conditions, the data were not continuous, and all datasets were calculated by Equation (9). The WMO calibration factors applied for XCO_2 , XCH_4 , and XCO were 0.9869, 0.9898, and 0.925, respectively.

The diurnal variations of XCO₂, XCH₄, and XCO are illustrated in Figure 6. The time series of XCO₂, XCH₄, and XCO show a significant diurnal variation. The XCO₂ and XCH₄ have similar daily variation, these reached a maximum at 15:00 p.m., then dropped until sunset. The similar daily variation indicates that XCO₂ and XCH₄ have a good correlation, detailed analysis in Section 4.4. The daily variation of XCO slowly varies, it reaches a minimum at 16:00 p.m., then climbs until sunset. Time series of daily averages of XCO₂, XCH₄ and XCO are plotted in Figure 7. The daily average of XCO₂ ranged from 415.09 to 421.78 ppm during the campaign. XCH₄ ranged from 1.96 to 2.02 ppm with a mean of 1.982 ppm, showing higher variation than XCO₂. The concentration of XCH₄ shows a positive correlation with the temperature. This might be due to the warmer weather increasing the activity of methanogens [20], resulting in higher atmospheric XCH₄. The daily average of XCO shows a similar variation tendency to XCH₄. The highest XCO



concentration was observed at the beginning of the measurements. The daily average of XCO ranges from 0.118 to 0.157 ppm with a mean of 0.137 ppm.

Figure 6. The four plots show the diurnal variation of XCO₂, XCH₄, and XCO from 30 September 2018 to 8 October 2018.



Figure 7. Time series of the daily averages of XCO_2 (a), XCH_4 (b) and XCO (c) observed with the EM27/SUN spectrometer.

4.4. The Correlation between XCH₄, XCO, and XCO₂

The sources of the CO_2 emissions are fossil fuel use and biological respiration. Photosynthesis is the sink of CO_2 . The main sources of CH_4 are biogenic and artificial. The biogenic source comes from landfills and wetlands, oceans and forest, and artificial sources include fossil fuel burning, waste treatment, and geological sources [21]. In the paper of Denman, CH_4 produced by gas and oil production, industry, landfills, and waste treatment accounts for 15% to 40% of global anthropogenic CH_4 emissions. In addition, bacteria decompose organic carbon, converting it to CO_2 and CH_4 . CO comes from incomplete combustion, the main sources are biomass burning and fossil fuel [22]. Our site is located on the coastline on the south side of the Liantou peninsula. The coastline is northeast-southwest. It is adjacent to the South China Sea and the north is a sparsely populated hilly area. Due to air–sea exchange, the photosynthesis of algae and the breathing of animals, the oceans are the sources and sinks of CO_2 and CH_4 .

A correlation study was carried out between XCH₄, XCO, and XCO₂ for the entire study period. The diurnal variations of XCH₄, XCO, and XCO₂ are highly correlated (Figure 8). Fang et al. (2015) suggest that a correlation coefficient value higher than 0.50 indicates a similar source of CO₂ and CH₄ [23]. The correlations were determined by linear regression of the data shown in Figure 7. Our study also reveals a strong positive correlation observed between XCO₂ and XCH₄. For the relationship between XCO₂ and XCH₄, the linear regression line shows a good correlation with the correlation coefficient R² \ge 0.5 (except on 2 and 3 October)—especially, the correlation coefficient R² = 0.82 on 8 October. The strong correlation between CO₂ and CH₄ indicates that atmospheric CO₂ and CH₄ are generated from common sources.

CO is a product of inefficient combustion that has often been used as a tracer of CO_2 from combustion [24]. The correlation slope of XCO to XCO_2 provides a characteristic signature of source regions and source type [25]. In the study of Wunch et al. (2009), the slope of the correlations of XCO to XCO₂ was 11 ± 2 ppb ppm⁻¹ in the South Coast Air Basin around Los Angeles [26]. Wang wei et al. (2017) calculated the correlation slope of CO to CO₂ at the Hefei site as 5.66 ppb ppm^{-1} on 25 October 2014 [27]. However, our studies found a weak correlation in the variation of CO₂ and CO (Figure 8b) during the observations. The correlation coefficient $R^2 \ge 0.5$ was only on 30th September and 3 October; the weak correlation between CO₂ and CO shows that there is a low influence of combustion emissions on CO_2 . The correlation slopes of CO to CO_2 were 1.94 ppb ppm⁻¹ and 3.69 ppb ppm⁻¹ on 30th September and 3 October, respectively. The correlation coefficient was much smaller on the other days, suggesting that CO₂ is dominated by the biosphere. In particular, the overall respiration component from a densely populated urban area may be significant relative to combustion because respiratory CO₂ emissions by urban residents are collocated with urban combustion sources. Our studies found a strong correlation in the variation of CO_2 and CH_4 (Figure 8a) during the observations. The correlation coefficient of 6 days was greater than 0.5 with a maximum of 0.77. It shows that CO_2 and CH_4 have the same source.

The HYSPLIT (the Hybrid Single-Particle Lagrangian Integrated Trajectory) model developed by the NOAA (the National Oceanic and Atmospheric Administration), based on the characteristics of the Lagrangian trajectory, can simulate the track of airflow and clearly indicates the source of the flow. It has been applied to studies on weather and climate. The HYSPLIT model was used to analyze the trajectories of air masses [28]. The calculated trajectories are helpful for resolving the evolution of airflow along the transport pathway.

The concentration of trace gases can be modulated not only by local emission but also by long transport from other regions. During the observation period, the diurnal variations of XCH₄ and XCO₂ were highly correlated. We computed HYSPLIT backward trajectories for the Maoming area. The trajectories used the GDAS model data on a 0.5 degree latitude longitude spatial resolution. The levels of 1 km, 5 km, and 10 km were taken as the initial height. The time interval was 6 h for output. Figure 9 shows backward air trajectories produced by HYSPLIT between 30 September and 8 October 2018. Figure 9a shows the 72 h backward movement trajectories on 3 October. The airflows in Figure 8a are mainly affected by long-range transport from India, Yunnan, and Jiangxi Province. Trajectories for 5 km and 10 km heights indicate movement of air masses from western India and the Yunnan province of China, whereas the trajectories dram at a height of 1 km give hint of trace gas transport from the Jiangxi province. Figure 9b shows the 72 h backward movement trajectories on 6 October. Trajectories at 5 km and 10 km heights indicate movement of air masses from the Bay of Bengal, whereas the trajectories dram at a height of 1 km give a hint of trace gas transport from Jiangsu province. Figure 9c shows the 72 h backward movement trajectories on 9 October. The trajectories dram at a height of 10 km gives a hint of trace gas transport from the Arabian Sea, whereas trajectories for 1 km and 5 km heights indicate movement of air masses from Myanmar and Jieyang City. From 30 September to 6 October, the upper atmosphere was mainly transported by external sources and during the whole observation period, the lower atmosphere was mainly affected by local sources. This shows that the high correlation between CO₂ and CH₄ in Figure 8 is due to the transportation from external sources.



Figure 8. The correlation between XCH_4 and XCO_2 (**a**), the correlation between XCO and XCO_2 (**b**) from 30 September 2018 to 8 October 2018.



Figure 9. Cluster analysis of 72 h back trajectories of Maoming from 30 September–8 October, 2018, (**a**) 30 September–2 October, (**b**) 2–5 October, (**c**) 6–9 October.

5. Conclusions

In this paper, a solar observatory was deployed at Maoming China to collect nearinfrared solar spectra. We assessed the influence of instrumental line shape degradation on the retrievals of the greenhouse gases. The study concluded that the influence of instrumental line shape degradation can be expressed as the modulation efficiency amplitude influence and the phase error influence. The modulation efficiency amplitude influence is the most important compared to the phase error influence. The means of XH_2O , XCH_4 , and XCO with ME values have a positive correlation. The correlation coefficients are 0.9925, 0.9968, and 0.9981 respectively. However, the relationship between the mean of XCO_2 and ME is opposite, with a correlation coefficients of 0.986. For a typical ILS degradation of 1% in modulation efficiency, the column-averaged dry air mole fraction of H_2O , CO_2 , CH_4 , and CO changed by 0.0492, 0.1151, 0.1042, 0.0523%, respectively, while an increase of 0.01 in phase error led to a decrease of 0.697% on XH_2O , 0.689% on XCO_2 , 0.652% on XCH_4 and 0.737% on XCO.

The column-averaged dry air mole fraction of CO₂, CH₄ and CO were successfully retrieved from low-resolution ground-based FTS (EM27/SUN) measurements. The daily average of XCO₂ ranged from 415.09 to 421.78 ppm during the campaign. XCH₄ ranged from 1.96 to 2.02 ppm with a mean of 1.982 ppm, showing higher variation than XCO₂. The daily average of XCO ranged from 0.118 to 0.157 ppm with a mean of 0.137 ppm. At the same time, we analyzed CH₄ and CO correlation with CO₂. Our study revealed a strong positive correlation observed between XCO₂ and XCH₄ with the correlation coefficient $R^2 \ge 0.5$. The strong correlation between CO₂ and CH₄ indicates that atmospheric CO₂ and CH₄ are generated from common sources. However, there was a weak correlation between CO and CO₂ and the correlation slopes of CO to CO₂ were 1.94 ppb ppm⁻¹ and 3.69 ppb ppm⁻¹ on 30 September and 3 October, respectively; the CO₂ is dominated by the biosphere. The results of backward movement trajectories, indicate that the airflows are mainly affected by long-range transport from the Arabian Sea, India, Myanmar, Yunnan and Jiangxi Province. In the near future, more long-term in situ measurements are needed.

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