

## Article

# In-Flight Calibration of GF-1/WFV Visible Channels Using Rayleigh Scattering

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**Abstract:** China is planning to launch more and more optical remote-sensing satellites with high spatial resolution and multistep gains. Field calibration, the current operational method of satellite in-flight radiometric calibration, still does not have enough capacity to meet these demands. Gaofen-1 (GF-1), as the first satellite of the Chinese High-resolution Earth Observation System, has been specially arranged to obtain 22 images over clean ocean areas using the Wide Field Viewing camera. Following this, Rayleigh scattering calibration was carried out for the visible channels with these images after the appropriate data processing steps. To guarantee a high calibration precision, uncertainty was analyzed in advance taking into account ozone, aerosol optical depth (AOD), seawater salinity, chlorophyll concentration, wind speed and solar zenith angle. AOD and wind speed were found to be the biggest error sources, which were also closely coupled to the solar zenith angle. Therefore, the best sample data for Rayleigh scattering calibration were selected at the following solar zenith angle of 19–22° and wind speed of 5–13 m/s to reduce the reflection contributed by the water surface. The total Rayleigh scattering calibration uncertainties of visible bands are 2.44% (blue), 3.86% (green), and 4.63% (red) respectively. Compared with the recent field calibration results, the errors are −1.69% (blue), 1.83% (green), and −0.79% (red). Therefore, the Rayleigh scattering calibration can become an operational in-flight calibration method for the high spatial resolution satellites.

**Keywords:** Rayleigh scattering; Gaofen-1 satellite (GF-1); absolute radiometric calibration; uncertainty analysis

## 1. Introduction

Gaofen-1 (GF-1), as the first satellite of the Chinese High-resolution Earth Observation System (CHEOS), was successfully launched 26 April 2013. This satellite provides data for land and resources surveys, disaster prevention and other fields. GF-1 has two high spatial resolution cameras and four Wide Field Viewing cameras (GF-1/WFV) [1]. The resolution of the WFV is 16 m, which includes four

spectral bands. The China Centre for Resource Satellite Data and Application (CRESDA) conducts field calibration experiments every year in the Dun Huang desert site [2].

The annual response changes of optical sensors cannot be ignored, especially at the bands of less than 500 nm [3]. The GF-1/WFV changed  $-1.41\%$  (blue),  $-4.72\%$  (green),  $-9.24\%$  (red), and  $8.78\%$  (near infrared), according to the calibration coefficients published on the website of CRESDA [4]. To guarantee the accuracy of quantitative application, in-flight calibration should be conducted at a high frequency [5]. GF-1 has no onboard calibration instruments, so vicarious and cross calibration methods should be developed [6,7]. Nowadays, the operational method taken by CRESDA for high spatial resolution satellites uses the Dun Huang desert site as a vicarious optical reference [7].

There exist various methods for in-flight radiometric calibration [8]. Slater et al. [9] calibrated the Thematic Mapper (TM) sensor on Landsat 4 using the White Sands site, which is uniform and flat in a large area. Biggar et al. [10] calibrated two sensors on the Earth Observing 1 (EO-1) satellite, performing many experiments at a variety of sites. Sun et al. [11] calibrated the Medium-Resolution Spectral Imager (MERSI) on the Chinese meteorological satellite FY-3A using the Dun Huang site annually.

With the increase in spatial resolution, optical sensors have multi-level gain and exposure time settings. Therefore, an optical sensor may have multiple calibration coefficients, with these multiple calibrations needing to be completed in a short time. In-flight calibration needs to use easily obtained natural scenes as a vicarious optical reference to improve the calibration efficiency.

Considering that the field calibration experiments need a considerable number of instruments, money and manpower in addition to not being applicable for history data, calibration methods that do not need field measurements were developed, such as cross [12], desert [13], moon [14], cloud [15] and Rayleigh scattering [16] calibration.

The Rayleigh scattering calibration method is developed for the in-flight vicarious calibration in the 443–670 nm wavelength bands using remote-sensing images over a clean ocean. Over the clean ocean area that is far away from the main land, the water and aerosol make only a small contribution to the radiance at the Top of Atmosphere (TOA) in the 443–670 nm wavelength bands, so the TOA signal is mainly contributed from the Rayleigh scattering of gas molecules. The sensor can be calibrated depending on the high precision of Rayleigh scattering radiative calculations.

The Rayleigh scattering calibration method, first proposed by Fraser and Kaufman [17], was used for the Visible Infrared Spin-Scan Radiometers (VISSR) onboard the Geostationary Operational Environmental Satellites (GOES5 and GOES6) with 2% error. Vermote et al. [18] calibrated the 450 nm and 550 nm bands for the Satellite Pour l'Observation de la Terre (SPOT) using Rayleigh scattering with 3% and 5% errors. Hagolle et al. [19] calibrated the Polarization and Directionality of the Earth's Reflectances instrument (POLDER), while Briottet et al. applied this method to vegetation [20]. Vermote and Kaufman calibrated the Advanced Very High-Resolution Radiometer (AVHRR) with an uncertainty of 2–3.5% [21]. Therefore, the Rayleigh scattering calibration has been applied to operational calibration of satellite sensors for its good accuracy.

The GF-1/WFV is calibrated using oceanic images and other supporting data. First, the necessity and research status of Rayleigh scattering calibration for high-resolution satellites are introduced. Second, the materials, principle and the procedure of the Rayleigh scattering calibration for the high spatial resolution GF-1/WFV are introduced. After that, the Rayleigh scattering calibration was performed and the primary results are shown. The rule of data selection is important to ensure high calibration accuracy when the Rayleigh scattering calibration method is applied to a sensor with high spatial resolution. Fourth, as the main part of this paper, the calibration uncertainty is discussed, considering each oceanic or atmospheric parameter using the radiative transfer model, which is named Second Simulation of the Satellite Signal in the Solar Spectrum (6S) [22]. Fifth, the final calibration results are calculated using the selected samples and total uncertainties are given. Finally, according to the uncertainty analysis, some data processing and sample selecting experiences are concluded to guarantee a high accuracy.

## 2. Materials and Methods

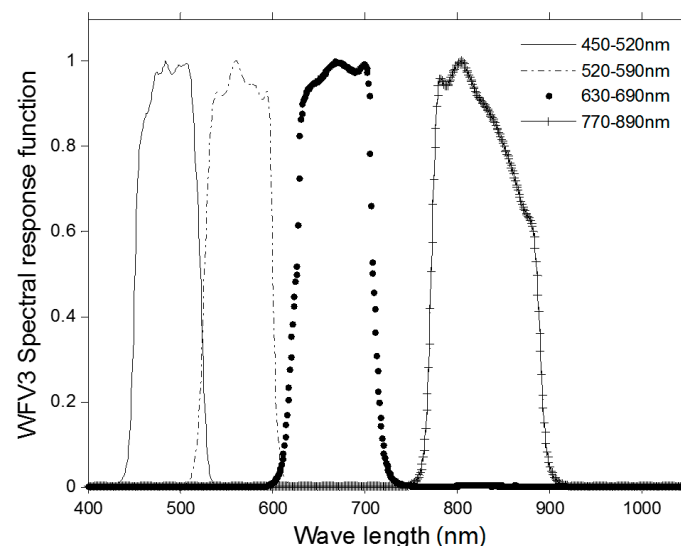
### 2.1. GF-1/WFV Sensor Characteristics and Imagery

GF-1/WFV is composed of four cameras in order to achieve both a large swath (800 km) and a high spatial resolution (16 m). The large swath brings a shorter time of global coverage, so the GF-1/WFV is widely used in many quantitative remote-sensing fields. Therefore, the radiometric calibration of GF-1/WFV has become more important. The Rayleigh scattering calibration method can improve the calibration time interval. The GF-1/WFV has a good response range and linearity in observing both dark targets (such as water) and bright targets (such as deserts), so the water objects could also be detected with a high sensitivity and signal-to-noise ratio. GF-1/WFV has the potential to use Rayleigh scattering signals for calibration over ocean.

From September to October 2015, the third camera WFV3 of GF-1/WFV was arranged to obtain a series of remote-sensing images over ocean areas to test the Rayleigh scattering calibration method. The other three WFVs were designed to be the same as WFV3, but there exist radiometric and geometric differences in reality. The specifications of a WFV are shown in Table 1. The spectral response functions are shown in Figure 1.

**Table 1.** Primary characteristics of the Wide Field Viewing cameras onboard Gaofen-1 satellite (GF-1/WFV).

Sensor	Bands (nm)	Central Wave Length (nm)	Spatial Resolution	Swath	View Zenith Angle	Revisit Period
WFV	450–520	484	16 m	800 km	WFV2 and WFV3: 0–24°	4 days
	520–590	560				
	630–690	665				
	770–890	800			WFV1 and WFV4: 24–40°	



**Figure 1.** Spectral response functions of the third Wide Field Viewing camera onboard Gaofen-1 satellite (GF-1/WFV3).

### 2.2. Atmospheric and Oceanic Parameters

The atmospheric and oceanic parameters are used to support the radiative transfer model (RTM) calculation as well as to estimate the calibration error. Before calibration calculation, the parameters measured simultaneously to the image should be prepared, including AOD, total ozone amount, seawater salinity, chlorophyll concentration and wind speed. The AOD is obtained from the Moderate

Resolution Imaging Spectroradiometer (MODIS) Level-2 AOD product (MOD04), which has a spatial resolution of  $10 \text{ km} \times 10 \text{ km}$  and an accuracy of  $\pm(5\% \text{AOD} + 0.03)$  [23]. The total ozone amount is obtained from the Total Ozone Mapping Spectrometer (TOMS) and Ozone Monitoring Instrument (OMI) daily products with a spatial resolution of  $1^\circ$  (latitude)  $\times 1.25^\circ$  (longitude) and an accuracy of  $\pm 2\%$  [24]. The seawater salinity is obtained from National Centers for Environmental Prediction (NCEP) Global Ocean Data Assimilation System (GODAS) with a spatial resolution of  $0.33^\circ \times 1^\circ$  and an accuracy of  $\pm 0.2\%$  [25]. The chlorophyll concentration is given by MODIS-Terra chlorophyll-a concentration data with spatial resolution of  $4.63 \text{ km} \times 4.63 \text{ km}$  with an accuracy better than 41% [26]. The wind speed is available as the European Center of Medium Weather Forecast (ECMWF) ERA-40 product with a spatial resolution of  $1^\circ \times 1^\circ$  and an accuracy of  $\pm 2 \text{ m/s}$  [27]. These data are selected to be synchronized with the satellite images for calibration.

### 2.3. Dun Huang Field Calibration Results

The calibration results released on the CRESDA website were used for validation. The Chinese optical satellite sensors are calibrated using measurements at the Dun Huang desert site. Usually in September and October every year, a comprehensive field experiment is conducted at the Dun Huang site. The surface reflectance can be measured using a spectroradiometer. The aerosol properties are measured using the French CIMEL CE318 automatic sun photometer synchronously with the satellite. The 6S RTM is used to calculate the TOA radiance to calibrate the sensor in the Dun Huang site calibration.

### 2.4. Principles and Procedures

Considering the minor contributions of aerosol scattering, water backscattering, whitecap diffuse reflection, surface specular reflection, gaseous absorption and the biggest contribution to Rayleigh scattering, the TOA radiance over ocean water can be calculated using an RTM [17]. Following this, the calibration coefficient can be given as a relationship between the Digital Number (DN) and the TOA radiance. The DN is the level 1A data of GF-1/WFV.

The radiance over a clean ocean received by a satellite sensor [28] is defined as

$$L = \left( L_{ray} + L_{aer} + (L_w + L_f) T_A + L_{oa} \right) \cdot T_g \quad (1)$$

where  $L_{ray}$  is the Rayleigh scattering contribution of gas molecules;  $L_{aer}$  is the scattering contribution of aerosol particles (when calculate  $L_{aer}$  using 6S, the aerosol-Rayleigh coupling is also included [22]);  $L_w$  and  $L_f$  represent the diffuse contributions of sea water and foam, respectively;  $T_A$  means the aerosol and Rayleigh transmission;  $L_{oa}$  represents the coupling contribution between ocean and atmosphere; and  $T_g$  is the gaseous transmission. When the atmosphere is clean, the  $L$  to which  $L_{ray}$  contributes as much as 90% [29] can be calculated accurately using a radiative transfer code.

The TOA radiance of band  $i$  can be written as

$$L_i = \frac{\int R_\lambda L_{i\lambda} d\lambda}{\int R_\lambda d\lambda} \quad (2)$$

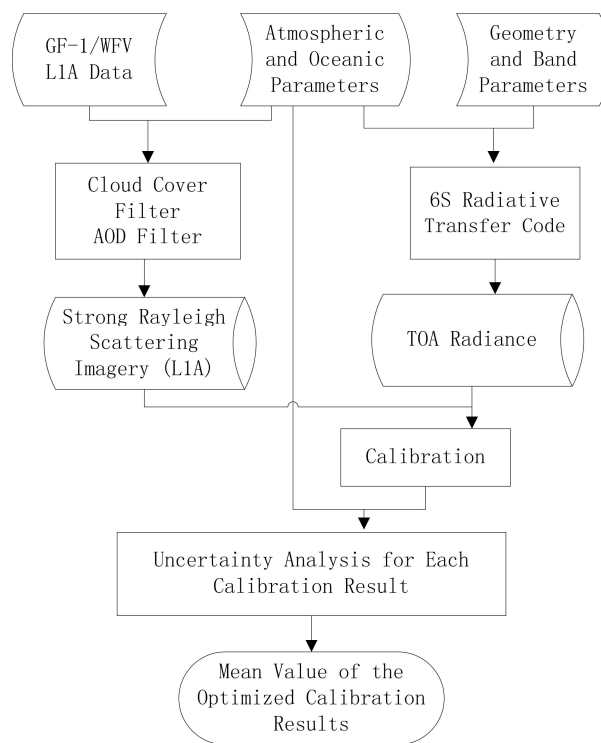
where  $R_\lambda$  is the spectral response of GF-1/WFV (Figure 1) and  $L_{i\lambda}$  is the TOA radiance wavelength  $\lambda$ . Vermote et al. [22] developed the 6S RTM, which improved the computational accuracy for aerosol and Rayleigh scattering compared with the successive order of scattering (SOS) algorithm [30]. The sea surface reflectance can be calculated as a function of the wind speed by the 6S model (details in Section 4.5.1). Therefore,  $L_i$  can be calculated by the 6S model using a dataset including spectral response, geometric angles, oceanic and atmospheric parameters, and wind speed.

The radiative response of an optical sensor describes the relationship between  $L_i$  and  $DN_i$ , which can be defined as

$$L_i = A_i \cdot DN_i + Offset_i \quad (3)$$

where  $DN_i$  is the digital number of band  $i$ , and the  $Offset_i$  is the band-specific additive rescaling bias. The GF-1/WFV level 1A data has been processed with a relative radiometric correction, including deduction of the offset induced by dark current and other factors [31]. In the Dun Huang calibration and the calibration of this paper, we consider  $Offset_i = 0$ . Therefore, the band-specific multiplicative rescaling gain  $A_i$  ( $\text{W} \cdot \text{m}^{-2} \cdot \text{sr}^{-1} \cdot \mu\text{m}^{-1}$ ) is the only absolute radiative calibration coefficient, which will be calculated by Equations (2) and (3).

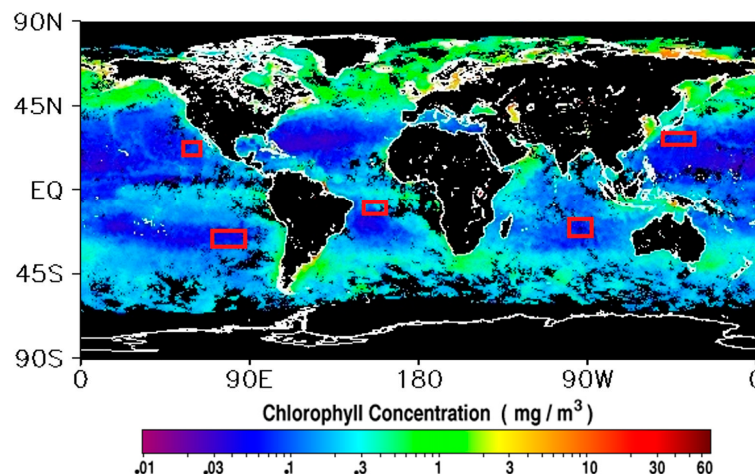
According to the Figure 2, the procedures of calibration are as follows:



**Figure 2.** The flow chart of the Rayleigh scattering calibration for the Wide Field Viewing cameras onboard Gaofen-1 satellite (GF-1/WFV).

(1) Obtain images over an ocean area for Rayleigh scattering calibration. As a land resources satellite, GF-1 has obtained seldom ocean images. For the Rayleigh scattering calibration, one of the four WFVs was specially arranged to obtain images over the ocean. The images for Rayleigh scattering calibration should follow the conditions [17] including: (a) case I waters for reducing the total reflection of seawater; (b) clean atmosphere for reducing aerosol scattering; and (c) large solar zenith angle for enhancing atmospheric molecules scattering. According to these conditions, we chose the case I water, which contains little chlorophyll in both spatial and temporal dimensions [32,33]. Figure 3 shows the September–October 2015 mean global distribution map of oceanic chlorophyll. GF-1/WFV obtained 31 images over the square areas shown in Figure 3 from September to October 2015.

(2) Obtain and process the atmospheric and oceanic parameters corresponding to each image. Strictly synchronized observations are difficult to obtain, so the nearest spatial and temporal observations are used instead. We assume that the related parameters have negligible changes within one hour, so the observations introduced in Section 2.2, such as AOD and wind speed, can be used in Rayleigh scattering calibration. The seawater salinity and chlorophyll concentration are mainly affected by precipitation, continental run-off and ocean current. These factors show slowly varying character, so daily data can be used for calibration. The influence of input parameters on the uncertainty analysis is discussed in the analysis part.



**Figure 3.** 2015 September–October global chlorophyll concentration (red squares show the target distribution for the Rayleigh scattering calibration, Unit:  $\text{mg}/\text{m}^3$ ).

(3) Select suitable images considering atmospheric and oceanic parameters. To ensure the atmosphere is “clean”, the selecting rules include: (a) the total cloud pixels accounted to no more than 15% in an image; (b) the cloud areas are not fragmented in an image (from visual inspection); (c) there exist no sun glint in an image (from visual inspection); and (d) the AOD should be smaller than 0.1 at the wavelength of 550 nm. Following this, 22 sample images were selected, which were distributed in five ocean areas (Table 2), including 11 in the Southeast Pacific, 6 in the Northwest Pacific, 2 in the Tropical Eastern Pacific, 2 in the South Indian Ocean and 1 in the Tropical South Atlantic.

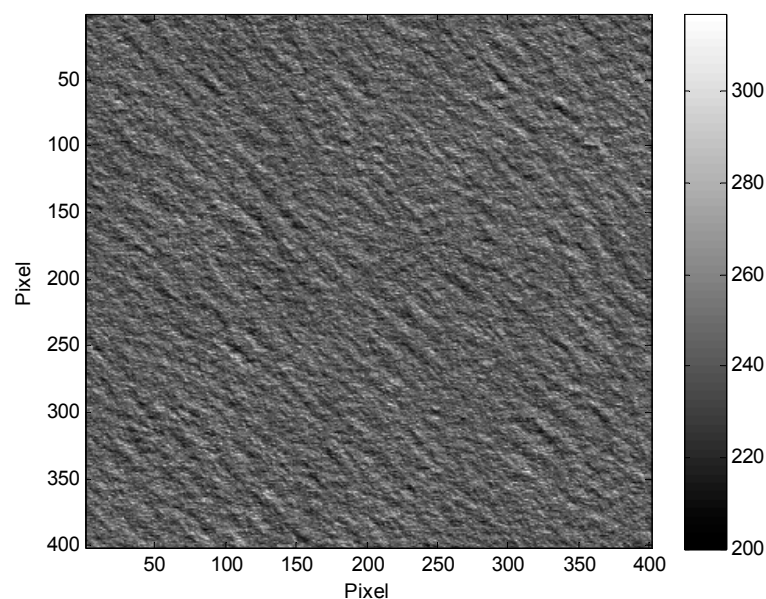
**Table 2.** The solar and viewing geometries as well as wind speed for the 22 samples.

Sample ID	Oceanic Area	Solar Zenith ( $^{\circ}$ )	Viewing Zenith ( $^{\circ}$ )	Relative Azimuth ( $^{\circ}$ )	Wind Speed (m/s)
1	Southeast Pacific	22.208	8.847	118.979	2.5
2		24.857	8.851	115.215	4
3		26.338	8.853	113.505	4
4		33.317	9.639	156.345	3
5		20.055	4.795	167.002	8
6		20.499	8.843	119.304	2.5
7		20.353	8.841	119.721	7.5
8		28.984	8.854	110.279	6
9		30.525	8.856	109.097	7.5
10		20.031	8.841	120.274	5
11		19.925	8.839	120.712	12.5
12	Northwest Pacific	39.394	10.734	81.76	8
13		34.3	16.388	138.434	5
14		34.901	8.787	113.513	4
15		35.935	15.816	105.528	7
16		37.831	8.78	112.186	7
17		25.549	8.798	122.76	6.1
18	Tropical Eastern Pacific	22.912	8.796	124.716	2.5
19	South Indian Ocean	20.09	8.799	120.712	2
20		33.998	10.816	68.908	9
21	Tropical South Atlantic	36.986	13.37	84.942	9
22		26.67	10.052	96.284	12.5

All 22 samples were used to complete the Rayleigh scattering calibration. In the discussion chapter of this paper, the optimal samples were selected with calibration uncertainty analysis.



(4) Process the cloud-free areas in the images. In the images from GF-1/WFV3, the bright and dark slopes of the sea waves are obviously different as shown in Figure 4. In order to ensure the accuracy and stability of Rayleigh scattering calibration, the DN values are averaged twice. Firstly, several cloud-free square areas (CFSA) in the image are selected and extracted to calculate the mean value (MCFSA). We tested different averaged scales of CFSA and at last, the side length is recommended to be 3–5 km because the averaged value in this scale is stable. Secondly, a number of CFSA in a limited viewing angle range (VAR, no more than  $1^\circ$ ) are selected to calculate each MCFSA. Thirdly, in the mean value sequence of MCFSAs, the single MCFSA that has a standard deviation bigger than 1% is removed. Finally, the rest mean values of MCFSAs are averaged again to get the final DN for calibration. Every parameter used for calibration is averaged to the same size as the calibration area (VAR) in the image.



**Figure 4.** Example of a slope distribution extracted from GF-1/WFV3 level 1A image. The size of one pixel is  $16\text{ m} \times 16\text{ m}$ . The Digital Number (DN) values are between 210 and 307.

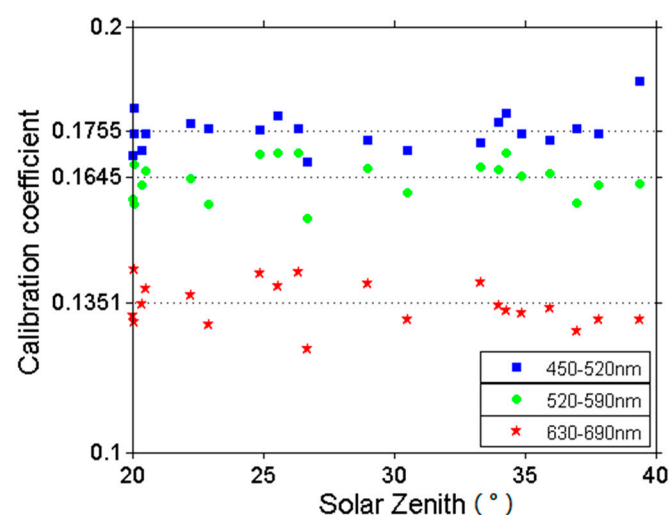
(5) Calculate the radiance and perform calibration. The 6S RTM is used to calculate TOA radiance with the inputs of solar and observing angles (center value in the VAR), month and day of imaging time, atmospheric model, aerosol model, spectral response function, atmospheric and oceanic parameters (mean value of each parameter in the VAR). The imaging time is from the metadata and the geometric conditions, including four angles, are from the geoinformation of GF-1/WFV image. The atmosphere model is selected from the built-in models, according to the geographical position and grabbing season. The aerosol model is set to be “Maritime”, which is also provided by the 6S model. The atmospheric and oceanic parameters are prepared from the measurements introduced in Section 2.2. The 6S can provide TOA radiance with a full input, following instructions in the main code. According to Equation (3), for each sample, the calibration coefficient can be calculated by comparing TOA radiance and the final DN mean value.

(6) Analyze the uncertainty and obtain the optimal calibration result. Jointly considering the calibration result and the error source, the optimal calibration coefficients are selected. At last, the optimal calibration result is the mean value of the optimal calibration coefficients.

### 3. Primary Results of All Samples

According to the principles and procedure of Rayleigh scattering calibration, the GF-1/WFV3 satellite sensor was calibrated using 22 oceanic images acquired in September and October in 2015. The

calibration results of the three visible bands at 450–520 nm (blue), 520–590 nm (green), and 630–690 nm (red) are shown in Figure 5. The mean values of 22 calibration coefficients (Table 3) are 0.1755 (blue), 0.1645 (green), and 0.1351 (red), respectively. Corresponding to the mean value, the largest deviation (defined and shown in Table 3) is 0.0117 (blue), 0.0117 (green), and 0.0126 (red) respectively, with the largest deviation ratios being in the range of 6.67–9.33%. For validation and assessment, the field calibration data are used to compare with the Rayleigh scattering calibration. The CRESDA conducted a field calibration campaign at the Dun Huang desert site in August 2015 for the land resources satellites. The calibration coefficients from this campaign are 0.1779 (blue), 0.1589 (green), and 0.1385 (red) [4]. Thus, the differences between Rayleigh scattering and field calibration are  $-0.0024$  ( $-1.35\%$ ),  $0.0056$  ( $3.52\%$ ), and  $-0.0034$  ( $-2.45\%$ ). The results of the two calibration methods are well consistent in the blue band, with bigger deviations occurring in the green and red bands. This may be caused by the weaker Rayleigh scattering of gas molecules and stronger aerosol scattering interference at the longer wavelength.



**Figure 5.** Absolute radiometric calibration coefficients ( $A_i$ ) of 22 samples computed by the Rayleigh scattering method versus solar zenith angle at wavelengths of 450–520 nm (blue), 520–590 nm (green), and 630–690 nm (red). The mean coefficients for the 3 bands are 0.1755, 0.1645, and 0.1351, respectively.

**Table 3.** The assessment for the Rayleigh scattering calibration of GF-1/WFV3 and the comparison with the Dun Huang field calibration.

Band	450–520 nm	520–590 nm	630–690 nm
Averaged calibration coefficients (ACC)	0.1755	0.1645	0.1351
Root mean square error	0.0041	0.0046	0.0052
Largest deviation (LD)	0.0117	0.0117	0.0126
Largest deviation ratio (LDR) *	6.67%	7.11%	9.33%
Calibration coefficients of the field calibration	0.1779	0.1589	0.1385
Relative deviation between Rayleigh scattering and field calibrations	$-1.35\%$	$3.52\%$	$-2.45\%$

\* LD is the biggest absolute difference of  $A_i$  with the mean  $A_i$ ; LDR = LD / ACC.

#### 4. Discussion of Uncertainty Analysis and Sample Selection Rules

The uncertainty of Rayleigh scattering calibration is dominated by the oceanic and atmospheric parameters. The ozone amount affects the gaseous absorption contribution, while the AOD affects the scattering contribution. Furthermore, the wind speed affects the whitecap and sun glint contributions, while the seawater salinity and chlorophyll concentration affect the water-leaving contributions.



The calibration uncertainties can be induced by the errors of seawater salinity, chlorophyll concentration, ozone amount, AOD, and wind speed. Through changing one of these parameters and keeping the rest unchanged, the radiance at TOA is recalculated utilizing the 6S model. Following this, the new calibration coefficient is recalculated. The uncertainty  $\sigma_i$  from each parameter of band  $i$  is

$$\sigma_i = \frac{A'_i - A_i}{A_i} \quad (4)$$

where  $A_i$  is the calibration coefficient from the observation input, and  $A'_i$  is the new calibration coefficient from the changed input.

The sea surface is typically a non-Lambertian object, so the uncertainty analysis of the Rayleigh scattering calibration focuses on the geometric conditions. The solar zenith angle is considered together with the analysis for the oceanic and the atmospheric parameters.

#### 4.1. Uncertainty from Seawater Salinity

The seawater salinity data is obtained from the Global Ocean Data Assimilation System (GODAS) with a  $\pm 0.2\%$  error [25]. Through adding the maximum error on the seawater salinity, the calibration coefficient is recalculated to analyze the calibration uncertainty. As shown in Table 4, the absolute maximum uncertainties in the 22 calibration samples are 0.002% (blue), 0.003% (green), and 0.006% (red), which are calculated by Equation (4). Consequently, the calibration uncertainty from seawater salinity is very small.

**Table 4.** Maximum and minimum value of  $\sigma_i$  induced by seawater salinity measurement errors in the 22 samples.

Uncertainty \ Band	Band	450–520 nm	520–590 nm	630–690 nm
Maximum		0.002%	0.003%	0.004%
Minimum		−0.002%	−0.003%	−0.006%

#### 4.2. Uncertainty from Ozone

The ozone amount is obtained from the TOMS and OMI daily product with a  $\pm 2\%$  error [24]. Through adding the maximum error on the ozone amount, the calibration coefficient is recalculated to analyze the calibration uncertainty. As shown in Table 5, the absolute maximum uncertainties in the 22 calibration samples are 0.03% (blue), 0.13% (green), and 0.08% (red), which are calculated by Equation (4). The green band has the biggest uncertainty because of the strong absorption of ozone at 500–650 nm. However, the ozone observation generally causes little uncertainty.

**Table 5.** Maximum and minimum value of  $\sigma_i$  induced by ozone measurement errors in the 22 samples.

Uncertainty \ Band	Band	450–520 nm	520–590 nm	630–690 nm
Maximum		0.03%	0.13%	0.08%
Minimum		−0.02%	−0.10%	−0.06%

#### 4.3. Uncertainty from Chlorophyll Concentration

The chlorophyll in seawater enhances the water-leaving reflectance, so its accuracy affects the calibration uncertainty. The MODIS chlorophyll-a observation is available only with an error reaching to 41% [26]. Through adding the error from the chlorophyll that amounts to a maximum of 41%, the calibration coefficient is recalculated to analyze calibration uncertainty. As shown in Table 6, the absolute maximum uncertainties in the 22 calibration samples are 0.70% (blue), 0.84% (green), and

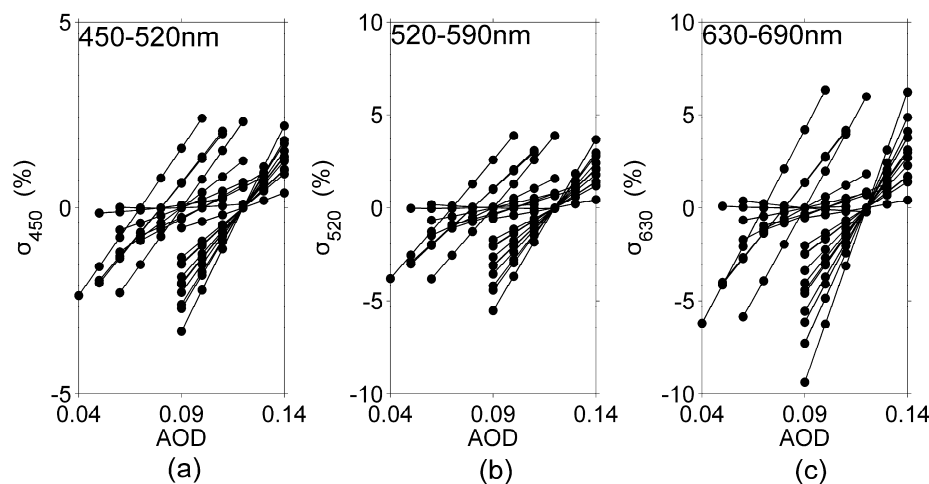
0.29% (red), which are calculated by Equation (4). An in-situ measurement of chlorophyll would improve the calibration accuracy.

**Table 6.** Maximum and minimum value of  $\sigma_i$  induced by chlorophyll concentration measurement errors in the 22 samples.

Uncertainty	Band	450–520 nm	520–590 nm	630–690 nm
Maximum		0.70%	0.58%	0.24%
Minimum		−0.61%	−0.84%	−0.29%

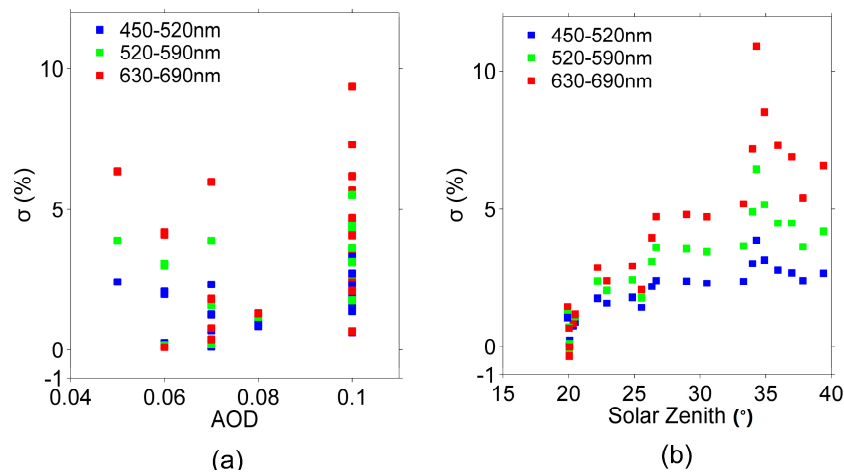
#### 4.4. Uncertainty from AOD

In order to reduce the uncertainty from the AOD, the calibration demands a clean atmospheric condition with AOD being no larger than 0.1 at the 550 nm wavelength. The AOD for calibration are taken from the MODIS product that has the minimum distances to the calibration image in both spatial and temporal dimensions. The error of the MODIS AOD product over the ocean is  $\pm(5\%AOD + 0.03)$  [23]. The “Ocean” aerosol model built-in the 6S RTM is chosen to calculate the radiance of the TOA. We added errors on the AOD input to the 6S RTM with a step width of 0.01 until the error reached  $\pm(5\%AOD + 0.03)$ , before we recalculated the calibration coefficients. As shown in Figure 6, the calibration uncertainty of every sample increases with a stable slope when the error of AOD is increasing, although these slopes vary widely.



**Figure 6.**  $\sigma_i$  (%), coefficient uncertainty with Aerosol Optical Depth (AOD) changes in steps of 0.01) of 22 examples varying with AOD at wavelengths of: (a) 450–520 nm; (b) 520–590 nm and (c) 630–690 nm (each line represents  $\sigma_i$  of one calibration sample with AOD changes).

With a further study, we found that the maximum calibration uncertainty has no stable relationship with the AOD (Figure 7a). When the AOD is added with an error of 0.035 (the maximum error of AODs of all the calibration samples), this results in the maximum calibration uncertainty. Jointly considering the principle of Rayleigh scattering calibration, a small AOD sample should be selected. In Figure 7b, the maximum calibration uncertainty increases with the Solar Zenith (SZ). The SZs of all the 22 samples are in 19–40°. When SZ is between 19 and 22°, the calibration uncertainty is at a minimum, which has values of 1.03% (blue), 1.28% (green), and 1.38% (red).



**Figure 7.**  $\sigma_i$  (%), coefficient uncertainty with Aerosol Optical Depth (AOD) changes  $\pm$  (5%AOD + 0.03)) of 22 examples versus: (a) AOD and (b) solar zenith angle (marked color represents different bands: 450–520 nm (blue), 520–570 nm (green), and 630–690 nm (red)).

#### 4.5. Uncertainty from Wind Speed

Wind speed changes rapidly in both dimensions of time and space, which brings uncertainty through affecting surface reflectance over the ocean. Firstly, in order to study the rules of each sample uncertainty, the sea surface reflectance was simulated using the 6S model, jointly considering wind speed ( $V$ ) and SZ. Secondly, we simulated the uncertainties induced by AOD and SZ in a wide range.

##### 4.5.1. Sample Uncertainty

Wind speed is taken from the ERA-40 dataset with an error of  $\pm 2$  m/s [27]. The uncertainty from the wind speed is evaluated by the Bidirectional Reflectance Distribution Function (BRDF). The total BRDF of the sea surface is calculated as follows [34]:

$$\rho_{os}(\theta_s, \theta_v, \varphi, \lambda) = \rho_{wc}(\lambda) + \{1 - W\} \cdot \rho_{gl}(\theta_s, \theta_v, \varphi, \lambda) + \{1 - \rho_{wc}(\lambda)\} \cdot \rho_{sw}(\theta_s, \theta_v, \varphi, \lambda) \quad (5)$$

$$W = 2.95 \times 10^{-6} V^{3.52} \quad (6)$$

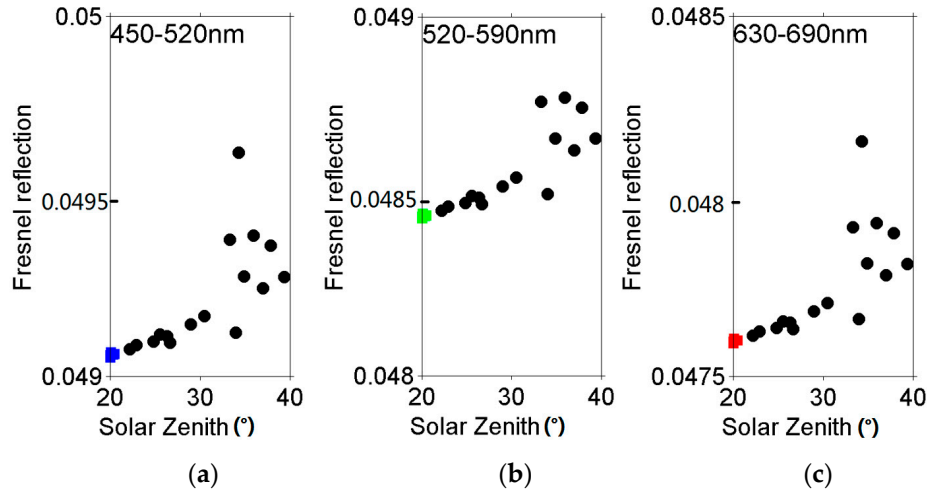
where  $\rho_{wc}$ ,  $\rho_{gl}$  and  $\rho_{sw}$  are the reflectance of whitecap, sun glint and water-leaving radiation, respectively;  $\lambda$  is the wavelength, so the total BRDF  $\rho_{os}$  gives a spectral reflectance with a certain  $\lambda$  range and fixed directions;  $V$  is the wind speed;  $\theta_s$  is the solar zenith;  $\theta_v$  is the observing zenith; and  $\varphi$  is the relative azimuth.  $\rho_{sw}$  is influenced by seawater salinity and chlorophyll concentration.  $W$  is the relative coverage ratio of the whitecaps, which is a function of  $V$  when the surface temperature is bigger than 14 °C [35]. For the coefficient uncertainty  $\sigma_i$  due to wind speed errors,  $\rho_{wc}$  and  $\rho_{sw}$  are negligible ( $10^{-3}$  in visible bands), compared to the biggest contributor  $\rho_{gl}$  ( $10^{-1}$ – $10^{-2}$ ).  $\rho_{gl}$  is calculated by the Cox and Munk model [36–38], which is a function of  $V$  and geometric conditions (in Equation (7)).

$$\rho_{gl}(\theta_s, \theta_v, \varphi_s, \varphi_v) = \frac{\pi P(Z'_x, Z'_y) R(n, \theta_s, \theta_v, \varphi_s, \varphi_v)}{4 \cos(\theta_s) \cos(\theta_v) \cos^4(\beta)} \quad (7)$$

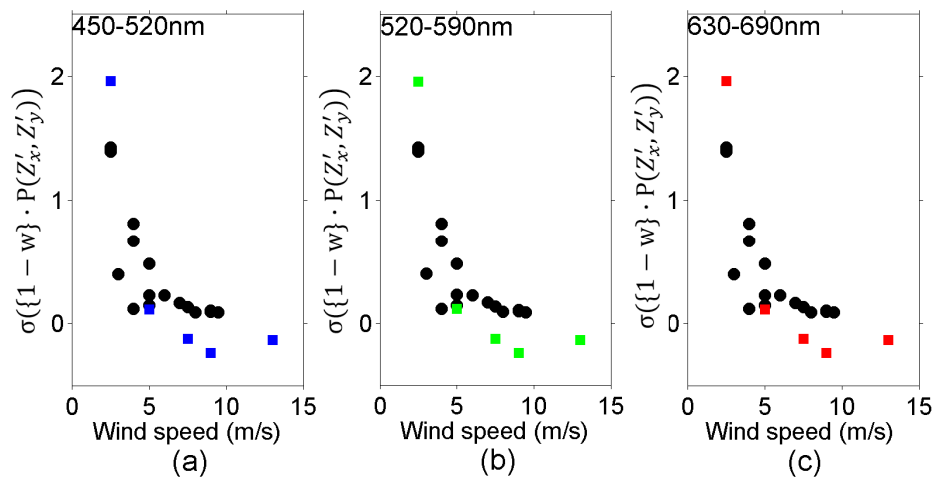
where  $P(Z'_x, Z'_y)$ , as a function of  $V$  [36–38], is the slope distribution expressed by a Gram-Charlier series;  $Z'_x$  and  $Z'_y$  are the two components of the surface slope;  $\beta$  is the tilt of the slope facet; and  $R(n, \theta_s, \theta_v, \varphi_s, \varphi_v)$  is Fresnel's reflection coefficient ( $n$  is the complex refractive index of sea water). Therefore, Equation (5) can be simplified to

$$\rho_{os}(\theta_s, \theta_v, \varphi, \lambda) \approx \{1 - W\} \cdot P(Z'_x, Z'_y) \cdot R(n, \theta_s, \theta_v, \varphi_s, \varphi_v) \cdot \frac{\pi}{4 \cos(\theta_s) \cos(\theta_v) \cos^4(\beta)} \quad (8)$$

where  $R(n, \theta_S, \theta_V, \varphi_S, \varphi_V)$  is related to geometric conditions. Figure 8 shows that Fresnel's reflection coefficient increases with SZ.  $\{1 - W\} \cdot P(Z'_x, Z'_y)$ , as the only term dominated by  $V$  changes in Equation (8), can be used to show the calibration uncertainty induced by  $V$ . With the  $\pm 2$  m/s error of  $V$ , Figure 9 shows that  $\{1 - W\} \cdot P(Z'_x, Z'_y)$  introduces little error, when  $V$  is larger than 5 m/s for all the samples.

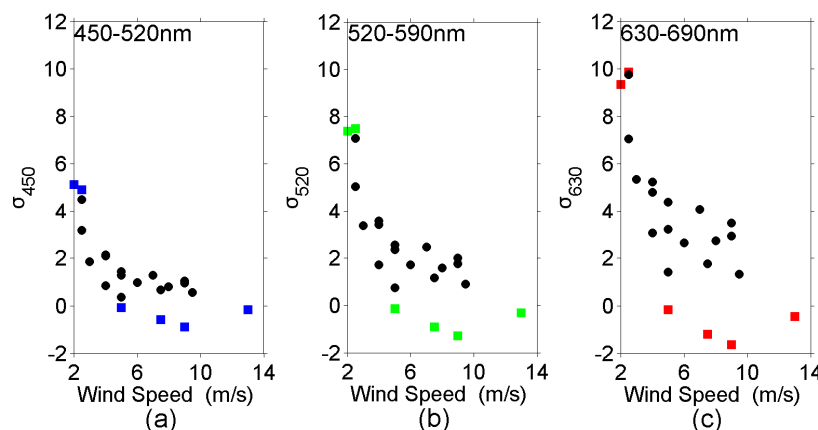


**Figure 8.** Fresnel reflection coefficients of 22 samples: (a) 450–520 nm; (b) 520–590 nm; (c) 630–690 nm (colored squares represent samples with solar zenith angles between 19 and 22°).



**Figure 9.**  $\{1 - W\} \cdot P(Z'_x, Z'_y)$  variations with a wind speed that changed  $\pm 2$  m/s in the 22 samples: (a) 450–520 nm, (b) 520–590 nm, and (c) 630–690 nm (colored squares represent samples with solar zenith angles between 19 and 22°).

According to the uncertainty analysis of AOD, the samples with a SZ of 19–22° (Sample ID 5, 6, 7, 10, 11 and 19) have small uncertainties. Meanwhile, as shown in Figures 7 and 8, when the  $V$  is larger than 5 m/s, the smaller SZ value samples have smaller variations in their Fresnel's reflection coefficients and  $\{1 - W\} \cdot P(Z'_x, Z'_y)$  induced by the  $V$  errors (colored squares in Figure 9, except for Samples 6 and 9 which both have a  $V$  of less than 5 m/s). Considering a  $V$  of 5–13 m/s, Samples 5, 7, 10, and 11 are selected. Figure 10 shows that the  $\sigma_i$  of these 4 samples reaches its minimum among 22 samples. Therefore, selecting the best samples should consider two conditions of SZ (19–22°) and  $V$  (5–13 m/s), based on the uncertainty analysis of 22 calibration samples.



**Figure 10.**  $\sigma_i$  (calibration coefficients uncertainty) with a wind speed that changed  $\pm 2$  m/s in the 22 samples: (a) 450–520 nm; (b) 520–590 nm and (c) 630–690 nm (colored squares represent samples with solar zenith angles between 19 and 22°).

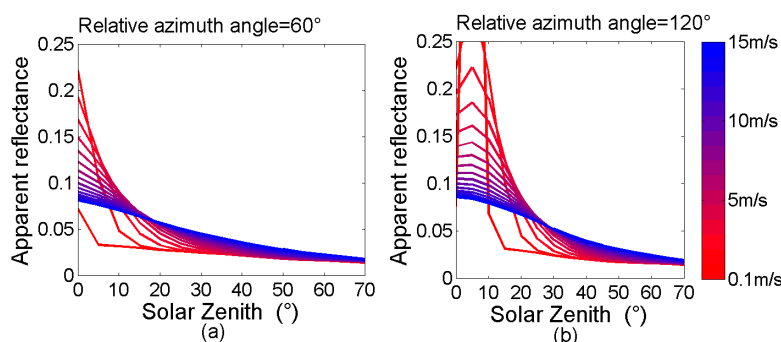
#### 4.5.2. Simulation Analysis

The analysis in Section 4.5.1 is based on 22 samples, with an SZ and  $V$  range of 19–40° and 2–13 m/s, respectively. An analysis with more general conditions is conducted using the 6S RTM (input parameters in Table 7) in order to study the influence of wind speed and geometric conditions. Under certain wind and geometric conditions, the TOA reflectance is insensitive to the wind speed error and presents stable calibration results.

**Table 7.** 6S input parameters for the simulation of wind speed and solar zenith angles.

	Solar Zenith Angle	View Zenith Angle	Relative Azimuth Angle	Wind Speed
Range	0–70°	10°	0–180°	0–15 m/s
Step	5°		60°	1 m/s

Figure 11 shows that there exist two conditions in which the calibration uncertainty is small. (1) When the SZ is large ( $>60^\circ$ ), the Rayleigh scattering of gas molecules contributes more to the TOA signals and the  $V$  error brings little uncertainty. This situation is consistent with Vermote et al. [22] and Dilligeard et al. [28]. (2) However, when the SZ is small ( $<30^\circ$ ), a larger  $V$  leads to little uncertainty. In this paper, the SZs of all the samples are not large enough so that the larger  $V$  is good for Rayleigh scattering calibration. Therefore, Figure 10 (small SZ and big  $V$ ) is consistent with the second condition in Figure 11.



**Figure 11.** TOA (Top Of the atmosphere) reflectance (630–690 nm) varying with solar zenith angle and wind speed, calculated at a relative azimuth angle of: (a) 60° and (b) 120°. The view zenith angle is 10° in both cases (a sun glint reflectance greater than 0.25 is not shown).



## 5. Final Results and Total Uncertainty

According to the uncertainty analysis, the calibration coefficient uncertainty induced by seawater salinity, chlorophyll concentration and ozone is not a large value. The uncertainty of the AOD should be considered, but its selection rule is not clear when AOD is not larger than 0.1. For selecting from the 22 samples, the sun zenith angle and wind speed was set at a range of 19–22° and 5–13 m/s, respectively. Under these conditions, 4 samples (ID: 5, 7, 10, and 11) were selected to give the mean calibration values of 0.1749 (blue), 0.1618 (green), and 0.1374 (red), while the largest deviation ratios were 3.89% (blue), 3.59% (green), and 4.8% (red). The calibration coefficients for each sample are listed in Table 8. The relative deviations between the final results of Rayleigh scattering calibration and the Dun Huang field calibration are −1.69% (blue), 1.83% (green), and 0.79% (red). Compared with the primary result (−1.35% (blue), 3.52% (green), and −2.45% (red)), the blue band has a slightly larger deviation, while the green and red bands results become much more stable.

**Table 8.**  $A_i$  (Calibration coefficients) of Samples 5, 7, 10, and 11.

Sample ID	Band			
		450–520 nm	520–590 nm	630–690 nm
5		0.1817	0.1656	0.1425
7		0.1713	0.1589	0.1335
10		0.1701	0.156	0.1308
11		0.1764	0.1668	0.1429
Mean (ACC)		0.1749	0.1618	0.1374
Largest deviation (LD)		0.0068	0.0058	0.0066
Largest deviation ratio (LDR)		3.89%	3.59%	4.8%

The averaged calibration coefficient uncertainties of the selected four samples are listed in Table 9. The uncertainty for each factor (ozone, AOD, sea water salinity, chlorophyll concentration and wind speed) is lower than that of 22 samples. The total uncertainty is the square root of the sum of the squares of the uncertainties from all five factors. The total uncertainties in Table 9 are 2.44% (blue), 3.86% (green), and 4.63% (red). As the Dun Huang field calibration accuracy is 5.3% [31], the Rayleigh scattering method becomes comparably even more accurate after sample selection.

**Table 9.**  $\sigma_i$  (Calibration coefficient uncertainties for Samples 5, 7, 10, and 11) from atmosphere and marine parameters (total uncertainty is the square root of the sum of the squares of the uncertainties from all five factors).

Factor	Band			
		$\sigma_i$ 450–520 nm	$\sigma_i$ 520–590 nm	$\sigma_i$ 630–690 nm
Ozone		0.39%	1.56%	0.61%
AOD		0.87%	0.98%	1.03%
Salinity		0.03%	0.04%	0.06%
Chlorophyll		0.17%	0.3%	0.06%
Wind speed		2.24%	3.39%	4.37%
Total uncertainty		2.44%	3.86%	4.63%

Moreover, the systematic calibration uncertainty induced by the RTM should be considered. The choice of the RTM may introduce an uncertainty less than 1%, although the errors are largely counterbalanced [39]. In this paper, we only considered the main uncertainty induced by measurements. Therefore, the absolute calibration uncertainty should be a little bigger than the total uncertainties in Table 9.

## 6. Conclusions

The absolute radiometric calibration is the key factor to determine the accuracy of the optical remote-sensing product. The WFOV payload onboard the GF-1 satellite was calibrated with the Rayleigh scattering method for its visible bands. Based on 31 ocean images of GF-1/WFOV, 22 targets were selected for the Rayleigh scattering calibration. Through calibration coefficients uncertainty analysis of related atmosphere and marine factors, the samples were selected with strict rules. According to the two fixed conditions of sun zenith angle ( $19\text{--}22^\circ$ ) and wind speed ( $5\text{--}13\text{ m/s}$ ), four samples were selected. The mean values for the calibration coefficients are 0.1749 (blue), 0.1618 (green), and 0.1374 (red), with a maximum uncertainty of less than 4.63% (caused by measurements only). This shows comparable accuracy with the Dun Huang field calibration.

In order to get more accurate results, the image processing and sample selection rules are as follows:

### (1) Target Scale

Due to the characteristics of light and dark surfaces distinguished in a high spatial resolution sensor image, the averaging method for reducing the scale is used to eliminate the influence of the wave surface. The spatial resolution of a WFOV image is 16 m and  $3\text{--}5\text{ km}$  is suitable for scale selection, which ensures accuracy of the surface radiation calculation.

### (2) Quality Recommendations

The coefficient uncertainty induced by wind speed shows the most obvious variations among all atmospheric and oceanic factors. Considering the simulation results ( $\pm 2\text{ m/s}$  for wind speed error), the selection rule for optimal targets concluded from TOA reflectance simulations needs specifying a combination of geometry conditions and wind speed. Therefore, the sample selection standard for quality control could be concluded as follows.

Samples with a large sun zenith angle ( $>60^\circ$ ) represent accurate calibration results, while simulation analysis indicates that more samples can be selected with a small sun zenith angle and fast wind speed. According to the simulation and sample uncertainty analysis, the sun zenith angle of  $19\text{--}22^\circ$  and a wind speed of  $5\text{--}13\text{ m/s}$  were chosen for the selection rule. The selected samples performed with a small coefficient of uncertainty of 2.44% (blue), 3.86% (green), and 4.63% (red).

Although the sample number in this calibration campaign is limited and not in accordance with the original selection rule of a large sun zenith angle, after using the TOA reflectance simulation with varying sun zenith angle and wind speed, the sample selecting rules were specified and the calibration accuracy was improved by using the above sample selection rule, becoming slightly better than the field calibration. Furthermore, high spatial resolution satellites could be arranged to acquire ocean images in a short time (possibly one day) with multi-level gains and exposure time settings. Following this, the multiple calibrations will be finished quickly and easily. As for the calibration of the near-infrared bands of GF-1/WFOV, sun glint and cloud calibration [19] can be implemented by transferring the Rayleigh calibration result of the visible bands. As some in-situ measurements of the atmospheric and oceanic parameters can be obtained from some stable sites, the Rayleigh scattering calibration could be an effective method for the operational in-flight calibration of high-resolution terrestrial satellites.

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