



# Article Gas-Phase Infrared Action Spectroscopy of CH<sub>2</sub>Cl<sup>+</sup> and CH<sub>3</sub>ClH<sup>+</sup>: Likely Protagonists in Chlorine Astrochemistry

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Abstract: Two fundamental halocarbon ions,  $CH_2Cl^+$  and  $CH_3ClH^+$ , were studied in the gas phase using the FELion 22-pole ion trap apparatus and the Free Electron Laser for Infrared eXperiments (FELIX) at Radboud University, Nijmegen (the Netherlands). The vibrational bands of a total of four isotopologs,  $CH_2^{35,37}Cl^+$  and  $CH_3^{35,37}ClH^+$ , were observed in selected wavenumber regions between 500 and 2900 cm<sup>-1</sup> and then spectroscopically assigned based on the results of anharmonic force field calculations performed at the CCSD(T) level of theory. As the infrared photodissociation spectroscopy scheme employed probes singly Ne-tagged weakly bound complexes, complementary quantum-chemical calculations of selected species were also performed. The impact of tagging on the vibrational spectra of  $CH_2Cl^+$  and  $CH_3ClH^+$  is found to be virtually negligible for most bands; for  $CH_3ClH^+$ –Ne, the observations suggest a proton-bound structural arrangement. The experimental band positions as well as the best estimate rotational molecular parameters given in this work provide a solid basis for future spectroscopic studies at high spectral resolutions.

Keywords: chlorine chemistry; astrochemistry, vibrational spectroscopy; ion trap; action spectroscopy

#### 1. Introduction

Out of the more than 300 inter- and circumstellar molecules detected to date, more than 25% contain heavy elements from the second row of the periodic table. With nearly fifty species, molecules containing silicon and sulfur are particularly abundant (an up-to-date list of astronomically detected molecules with complementary information is maintained at the Cologne Database for Molecular Spectroscopy (CDMS) [1,2] accessible online at https://cdms.astro.uni-koeln.de/ (accessed on 12 December 2023)).

In contrast, the number of chlorine-bearing molecules is much smaller. Only seven such species have been observed so far: HCl [3]; AlCl, NaCl, and KCl [4]; HCl<sup>+</sup> [5]; H<sub>2</sub>Cl<sup>+</sup> [6]; and CH<sub>3</sub>Cl [7]. The latter, pentatomic methyl chloride, or CH<sub>3</sub>Cl, is the most complex chlorine-bearing species presently known in space and, thus, an interesting target for studies of chlorine astrochemistry. In this context, it is assumed that CH<sub>3</sub>Cl may be formed through reactions on grain surfaces but also possibly in the gas phase through ion–molecule reactions, for example, in a scenario suggested by Garrod (cf., Ref. [8]).

$$CH_3^+ + HCl \rightarrow CH_2Cl^+ + H_2$$
 (1)

$$CH_2Cl^+ + H_2 \rightarrow CH_3ClH^+ + h\nu$$
 (2)

$$CH_3ClH^+ + e^- \rightarrow CH_3Cl + H$$
 (3)

Here, in the first step, the chloromethyl ion  $CH_2Cl^+$  is formed from a reaction of the methyl cation  $CH_3^+$  and HCl; in the second, the halonium ion  $CH_3ClH^+$  is produced via



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**Copyright:** © 2024 by the authors. Licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (https:// creativecommons.org/licenses/by/ 4.0/). radiative association between  $CH_2Cl^+$  and molecular hydrogen, finally yielding  $CH_3Cl$  through dissociate recombination with a free electron.

While the CH<sub>3</sub><sup>+</sup>, HCl, H<sub>2</sub>, and CH<sub>3</sub>Cl species are all known to be present in astronomical environments (the non-polar  $CH_3^+$  molecular ion being detected only very recently through observations with the James Webb Space Telescope [9]), our knowledge of the  $CH_2Cl^+$  and  $CH_3ClH^+$  molecular ions also participating in the above chain of reactions is still highly fragmentary. From the viewpoint of laboratory experiments, only little spectroscopic information has been collected previously for the chloromethyl cation. Photoelectron spectroscopy of the CH<sub>2</sub>Cl radical has provided a vibrational wavenumber  $\nu = 1040(30) \text{ cm}^{-1}$  of the Cl–C-stretching mode of CH<sub>2</sub>Cl<sup>+</sup> [10,11]. A somewhat more comprehensive picture of the vibrational spectrum has been obtained through infrared spectroscopy of  $CH_2Cl^+$  trapped in solid Ar matrices [12]. In these studies, four out of the six vibrational fundamentals were observed for the most abundant isotopic species, CH<sub>2</sub><sup>35</sup>Cl<sup>+</sup>. Using precursors enriched in <sup>13</sup>C or deuterium permitted the observation of four vibrational fundamentals of  ${}^{13}CH_2{}^{35}Cl^+$  or five of  $CD_2{}^{35}Cl^+$ , respectively. Owing to a significant isotopic shift, the  $C^{-37}$ Cl stretching mode could also be observed for all three isotopic species. Direct infrared spectroscopic studies of  $CH_2Cl^+$  in the gas phase have not been reported yet. Hexatomic protonated methyl chloride, CH<sub>3</sub>ClH<sup>+</sup>, has been known from mass spectrometry and studies of reaction kinetics [13,14]; however, it has never been spectroscopically studied.

In the present work, the vibrational spectra of the <sup>35</sup>Cl and <sup>37</sup>Cl isotopic species of  $CH_2Cl^+$  and  $CH_3ClH^+$  were observed in the gas phase, marking the first direct gas-phase infrared study of  $CH_2Cl^+$  and the first ever spectroscopic detection of  $CH_3ClH^+$ . The characterization of both ions was accomplished in an ion trap using a messenger action spectroscopy scheme, i.e., employing infrared photodissociation (IRPD) of their weakly bound complexes with neon,  $CH_2Cl^+$ -Ne, and  $CH_3ClH^+$ -Ne. Spectroscopic analysis was supported based on complementary high-level quantum-chemical calculations performed at the CCSD(T) level of theory. A detailed account of the experimental and theoretical work as well as of the analysis will be given in the following.

#### 2. Results and Discussion

#### 2.1. Structures of $CH_2Cl^+$ , $CH_3ClH^+$ , and Their Weakly Bound Complexes with Ne

Equilibrium molecular structures of  $CH_2Cl^+$  and  $CH_3ClH^+$  (and their isoelectronic sulfur-bearing counterparts, cf. Section 4) were calculated at the CCSD(T) level of theory using different basis sets. Full sets of internal coordinates are given in Appendices A and B. Best-estimate equilibrium structures were calculated using the cc-pwCVQZ basis set and all electrons in the correlation treatment, providing a level of theory previously shown to be of high accuracy [15]. The fc-CCSD(T)/cc-pV(T + d)Z level was employed for the calculation of harmonic and anharmonic force fields of the bare ions and used for the interpretation of their vibrational spectra. The impact of Ne-tagging on the molecular structures and harmonic force fields was estimated at the fc-CCSD(T)/aug-cc-pV(T + d)Z level of theory. To determine the energetically favorable locations of the Ne atom with respect to the  $CH_2Cl^+$  and  $CH_3ClH^+$  ions, potential energy surfaces (PESs) were calculated using grids of appropriate size and sampling (see also Ref. [16]). Briefly, in these calculations, the fc-CCSD(T)/aug-cc-pV(T + d)Z equilibrium structures of CH<sub>2</sub>Cl<sup>+</sup> and CH<sub>3</sub>ClH<sup>+</sup> were kept fixed while varying the position of the Ne atom on the grid over an area of a few tens of  $Å^2$ and at a spacing of 0.25 Å. At each of the grid points, single-point energy calculations were performed. PESs were calculated for the two mirror planes of  $CH_2Cl^+$  and for the single mirror plane of  $CH_3ClH^+$ . The resulting potential energy maps of  $CH_2Cl^+$  are shown in Figure 1.



**Figure 1.** fc-CCSD(T)/aug-cc-pV(T + d)Z potential energy maps of the CH<sub>2</sub>Cl<sup>+</sup>–Ne weakly bound complex calculated in the two mirror planes of CH<sub>2</sub>Cl<sup>+</sup> (see text for details). Atom color code: hydrogen (gray), carbon (black), chlorine (green). Contours show the potential energy of the complex as a function of the Ne atom position relative to the ion and cover the interval [0.1, 1.0] kcal/mol in steps of 0.1 kcal/mol above the global minimum.

Out of the four unique local minima seen, the shallow ones corresponding to linear C–Cl–Ne and Cl–C–Ne cluster variants were not considered further. The global minimum arrangements of each plot, i.e., one H-bound in-plane variant ( $CH_2Cl^+$  Ne-ip, Figure 1, left map) and the out-of-plane variant ( $CH_2Cl^+$  Ne-oop, right), were finally structurally refined in fully relaxed calculations. The corresponding structures are shown in Figure 2, and the full sets of structural parameters are given in Appendix A. As can be seen, the molecular structure of  $CH_2Cl^+$  hardly changes upon Ne-tagging at either position, and as a consequence, also the harmonic vibrational wavenumbers of  $CH_2Cl^+$  collected in Table 1 show very little dependence on tagging. The Ne-ip and Ne-oop species are calculated to be almost isoenergetic, with the latter configuration being slightly more stable by about 0.1 kcal/mol. Ne bond dissociation energies under the consideration of an estimate of zero-point vibrational effects amount to 0.6 kcal/mol (Ne-ip) and 0.7 kcal/mol (Ne-oop), respectively.



**Figure 2.** Bond lengths of CH<sub>2</sub>Cl<sup>+</sup> and its complexes with neon (in Å; fc-CCSD(T)/aug-cc-pV(T + d)Z values of all three species given in regular font and fc-CCSD(T)/cc-pV(T + d)Z values of the bare species given in italics). Bare CH<sub>2</sub>Cl<sup>+</sup> is of  $C_{2v}$  symmetry, whereas the symmetry of the Ne-complexes is lowered to  $C_s$ . Full sets of structural parameters are given in Appendix A. For further details, see text.

Mode	Bare Ion	Ne–ip	Ne-oop
$\omega_1$ CH sym stretching	3102 ( <i>a</i> <sub>1</sub> , 18)	3106 (a', 17)	3105 ( <i>a</i> ', 16)
$\omega_2$ CH <sub>2</sub> scissoring	1478 ( <i>a</i> <sub>1</sub> , 0.3)	1476 (a', 1)	1478 (a', 0.2)
$\omega_3$ Cl–C stretching	1059 ( <i>a</i> <sub>1</sub> , 80)	1059 (a', 78)	1058 (a', 82)
$\omega_4$ oop bending	1137 ( <i>b</i> <sub>1</sub> , 22)	1138 ( <i>a</i> ", 21)	1136 ( <i>a</i> ′, 20)
$\omega_5$ CH asym stretching	3239 ( <i>b</i> <sub>2</sub> , 51)	3243 (a', 59)	3243 ( <i>a</i> ", 50)
$\omega_6 \mathrm{CH}_2$ wagging	$1047 (b_2, 0.02)$	1048 (a', 0.3)	1047 ( <i>a</i> ", 0.003)
$\omega_7^{2,3}$		93 (a', 21)	131 (a'', 1)
$\omega_8^{2,4}$		70 (a'', 0.2)	89 (a', 22)
$\omega_9^{2,5}$		25(a', 1)	34(a', 4)

**Table 1.** Harmonic vibrational wavenumbers of  $CH_2^{35}Cl^+$  and  $CH_2^{35}Cl^+$ –Ne calculated at the CCSD(T)/aug-cc-pV(T + *d*)Z level of theory (in cm<sup>-1</sup>)<sup>1</sup>.

<sup>1</sup> Mode symmetries and infrared band intensities (km/mol) are given in parentheses. Mode indices of the Netagged variants are borrowed from bare CH<sub>2</sub>Cl<sup>+</sup> for the sake of comparability. <sup>2</sup> Extra-low-frequency vibrational modes introduced through Ne-tag, arbitrary mode index. <sup>3</sup> H–Ne stretching (Ne-ip species) and CH<sub>2</sub> twisting (Ne-oop species). <sup>4</sup> C–H–Ne stretching oop bending (Ne-ip species) and C–Ne stretching (Ne-oop species). <sup>5</sup> C–H–Ne ip bending (Ne-ip species) and Cl–C–Ne bending (Ne-oop species).

For CH<sub>3</sub>ClH<sup>+</sup>, by analogy with other rare gas-tagged protonated species, one would expect the Ne atom to be bound to the proton in a near-linear Cl-H····Ne fashion (e.g., Refs. [17–19]). Indeed, from the potential energy map obtained for the molecular mirror plane (Figure 3), this arrangement is identified as a deep, putatively global minimum. Another minimum is identified on the opposite side of the molecule and a third shallow one where Ne is arranged in a near-linear Cl-C-Ne fashion. Fully relaxed structures of the first two configurations ("p-bound" and "CH<sub>3</sub>/Cl-bound") are shown in Figure 4. According to those calculations, the p-bound variant is more stable by 0.4 kcal/mol. The third minimum as well as potential arrangements where the Ne atom would be located outside of the symmetry plane of CH<sub>3</sub>ClH<sup>+</sup> have not been studied further. The results of harmonic force field calculations to study the influence of Ne-tagging are presented in Table 2. As can be seen, the influence of tagging is virtually negligible for the majority, but not all, of the vibrational fundamentals (see further discussion in Section 2.3). Under the consideration of zero-point vibrational effects, the Ne bond dissociation energies are 1.0 kcal/mol and 0.7 kcal/mol for the p-bound and CH<sub>3</sub>/Cl-bound species, respectively.



**Figure 3.** fc-CCSD(T)/aug-cc-pV(T + d)Z potential energy map of the CH<sub>3</sub>ClH<sup>+</sup>–Ne weakly bound complex calculated in the mirror plane of CH<sub>3</sub>ClH<sup>+</sup> (see text for details). Atom color code: hydrogen (gray), carbon (black), chlorine (green). Contours show the potential energy of the complex as a function



of the Ne atom position relative to the ion and cover the interval [0.1, 1.0] kcal/mol in steps of 0.1 kcal/mol above the global minimum.

**Figure 4.** Bond lengths of CH<sub>3</sub>ClH<sup>+</sup> and the p-bound and CH<sub>3</sub>-bound complexes with neon (in Å; fc-CCSD(T)/aug-cc-pV(T + d)Z values of all three species given in regular font and fc-CCSD(T)/cc-pV(T + d)Z values of the bare species given in italics). All species exhibit  $C_s$  symmetry. Full sets of structural parameters are given in Appendix A. For further details, see text.

**Table 2.** Harmonic vibrational wavenumbers of  $CH_3^{35}ClH^+$  and two  $CH_3^{35}ClH^+$ –Ne variants calculated at the CCSD(T)/aug-cc-pV(T + *d*)Z level of theory (in cm<sup>-1</sup>)<sup>1</sup>.

Mode	Bare Ion	p-bound	CH <sub>3</sub> /Cl-bound
$\omega_1$ C–H asym stretching	3220 ( <i>a</i> ', 18)	3220 ( <i>a</i> ′, 17)	3224 (a', 20)
$\omega_2$ C–H sym stretching	3087(a', 5)	3087 ( <i>a</i> ′, 5)	3089 ( <i>a</i> ', 5)
$\omega_3$ Cl–H stretching	2823 (a', 176)	2811 ( <i>a</i> ′, 307)	2824 (a', 174)
$\omega_4$ CH <sub>3</sub> asym bending	1472 (a', 14)	1472 (a', 14)	1471 (a', 13)
$\omega_5 \text{ CH}_3 \text{ umbrella}$	1386 (a', 11)	1386 (a', 11)	1385 (a', 10)
$\omega_6$ CH <sub>3</sub> ip rocking	1118 (a', 3)	1120 ( <i>a</i> ′, 3)	1118 (a', 3)
$\omega_7$ C-Cl-H bending	771 ( <i>a</i> ′, 13)	784 ( <i>a</i> ′, 11)	773 (a', 14)
$\omega_8$ Cl–C stretching	565 (a', 51)	570 (a', 50)	566 (a', 50)
$\omega_9$ C–H asym stretching	3238 ( <i>a</i> ", 15)	3237 ( <i>a</i> <sup>''</sup> , 15)	3238 ( <i>a</i> <sup>''</sup> , 15)
$\omega_{10}$ CH <sub>3</sub> asym bending	1468 ( <i>a</i> ", 14)	1469 ( <i>a</i> ", 14)	1470 ( <i>a</i> <sup>''</sup> , 12)
$\omega_{11}$ CH <sub>3</sub> oop rocking	1040 ( <i>a</i> ", 0.5)	1040 (a'', 0.5)	1040 ( <i>a</i> ", 0.4)
$\omega_{12}$ torsion	202 (a'', 39)	244 ( <i>a</i> <sup>''</sup> , 43)	211 (a'', 42)
$\omega_{13}^{2,3}$		90 $(a', 7)$	69 (a', 8)
$\omega_{14}^{2,4}$		72(a'',7)	46(a', 0.4)
$\omega_{15}^{2,5}$		20 ( <i>a</i> ′, 7)	40 ( <i>a</i> ", 10)

<sup>1</sup> Mode symmetries and infrared band intensities (km/mol) are given in parentheses. <sup>2</sup> Extra-low-frequency vibrational modes introduced through Ne-tag, arbitrary mode index. <sup>3</sup> H–Ne stretching. <sup>4</sup> Ne–H–Cl oop bending (p-bound species) and Ne–H–Cl ip bending (CH<sub>3</sub>/Cl-bound species). <sup>5</sup> Ne–H–Cl ip bending (p-bound species) and Ne–H–Cl oop bending (CH<sub>3</sub>/Cl-bound species).

# 2.2. IRPD Spectrum of CH<sub>2</sub><sup>35</sup>Cl<sup>+</sup>-Ne

The m/z = 69 filtered IRPD spectrum from 900 to  $1600 \text{ cm}^{-1}$  is shown in Figure 5 (top red trace). From Table 3, it can be seen that four fundamental modes of the CH<sub>2</sub><sup>35</sup>Cl<sup>+</sup> species are calculated in this wavenumber range: the  $v_6$  mode at  $1038 \text{ cm}^{-1}$ ,  $v_3$  at  $1046 \text{ cm}^{-1}$ ,  $v_4$  at  $1134 \text{ cm}^{-1}$ , and  $v_2$  at  $1451 \text{ cm}^{-1}$ , only two of which have sizable calculated infrared intensities:  $v_3$  (Cl–C stretching) and  $v_4$  (out-of-plane (oop) bending). While infrared intensities are not strictly related to the intensities in the action spectroscopy scheme employed here, usually good qualitative agreement is observed. Clearly,  $v_3$  and  $v_4$  are detected at  $1048 \text{ cm}^{-1}$  and  $1129 \text{ cm}^{-1}$ , respectively, which is in very good agreement with the anharmonic CCSD(T)/cc-pV(T + d)Z values (1046 and  $1134 \text{ cm}^{-1}$ , as shown in Table 3) and also show proper intensity behavior. Both  $v_3$  and  $v_4$  show weak satellite band progressions with approximate spacings of  $20 \text{ cm}^{-1}$  to their blue sides (see inset in Figure 5), which is thought to originate from combination modes with the participation of the Ne-tag (most likely involving the  $\omega_9$  fundamental (Table 1); see,

e.g., Refs. [20,21]). The CH<sub>2</sub> scissoring mode  $v_2$  is calculated to be very weak and would be located in a region affected by somewhat higher noise and some baseline instabilities. The CH<sub>2</sub> wagging mode  $v_6$  is calculated to have almost no infrared intensity and is either too weak to be detected or overlapped with the much stronger  $v_3$  band. Apart from that, no other clear spectroscopic feature is observed. As concluded from harmonic fc-CCSD(T)/aug-cc-pV(T + *d*)Z calculations of bare CH<sub>2</sub>Cl<sup>+</sup> and its singly tagged Ne-clusters (Table 1), tagging is expected to have a negligible influence on the IRPD spectrum of this ion.



**Figure 5.**  $CH_2Cl^+$ –Ne and  $CH_3ClH^+$ –Ne infrared photodissociation (IRPD) spectra as obtained here. The traces indicate m/z = 69 corresponding to the  $CH_2^{35}Cl^+$ –Ne species (top spectrum in red),

m/z = 71 (isobaric mixture of CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup>–Ne and CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup>–Ne, blue center trace) and m/z = 73 (CH<sub>3</sub><sup>37</sup>ClH<sup>+</sup>–Ne, green bottom trace). For comparison, the vibrational fundamentals from VPT2 calculations (fc-CCSD(T)/cc-pV(T + d)Z) of the bare ions are shown as inverted sticks. For clarity, the intensities of selected calculated and experimental bands have been multiplied by the factors indicated. For the CH<sub>3</sub>ClH<sup>+</sup>–Ne species, dashed sticks indicate blueshifts of the  $v_8$  and  $v_7$  bands and a redshift of the  $v_3$  band, which is indicative of Ne being bound to the proton in a near-linear Cl-H···Ne fashion. The bands labeled with an asterisk are not vibrational fundamentals. Insets highlight blueshifted satellites, presumably due to combination modes involving the Ne tag. For further details, see text.

**Table 3.** Harmonic and anharmonic fc-CCSD(T)/cc-pV(T + *d*)Z vibrational wavenumbers (in cm<sup>-1</sup>) and infrared band intensities (int, km/mol), experimental wavenumbers (exp, cm<sup>-1</sup>), full width at half maximum of observed bands (FWHM, cm<sup>-1</sup>), and difference between observed and calculated anharmonic wavenumbers  $\delta$  (cm<sup>-1</sup>) of CH<sub>2</sub>Cl<sup>+</sup>.

Mada	$CH_2$ <sup>35</sup> $Cl^+$					
Mode –	Harm	Anharm	Int	Exp <sup>1</sup>	FWHM	δ
$\nu_1(a_1)$ CH sym stretching	3105.2	2993.3	14	пс		
$v_2(a_1)$ CH <sub>2</sub> scissoring	1491.1	1451.1	0.7			
$v_3(a_1)$ Cl–C stretching	1061.8	1045.6	81	1048	13	2
$v_4(b_1)$ oop bending	1151.9	1134.1	27	1129	17	-5
$v_5(b_2)$ CH asym stretching	3243.9	3101.1	47	пс	• • •	
$\nu_6(b_2)$ CH <sub>2</sub> wagging	1054.8	1038.2	< 0.01			•••
N. 1.	$CH_2^{37}Cl^+$					
Mode –	Harm	Anharm	Int	Exp <sup>1</sup>	FWHM	δ
$\nu_1(a_1)$ CH sym stretching	3105.2	2993.3	14	пс		
$v_2(a_1)$ CH <sub>2</sub> scissoring	1491.0	1451.0	0.7			
$v_3(a_1)$ Cl–C stretching	1053.6	1037.6	82	1037	15	-1
$v_4(b_1)$ oop bending	1151.8	1134.0	27	1125	13	-9
$v_5(b_2)$ CH asym stretching	3243.9	3101.1	47	пс	• • •	
$\nu_6(b_2)$ CH <sub>2</sub> wagging	1054.2	1037.6	< 0.01		•••	•••

<sup>1</sup> Wavenumber errors amount to 0.5% of the band position; *nc*: not covered.

# 2.3. IRPD Spectra of CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup>-Ne and CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup>-Ne

Because  $CH_2^{37}Cl^+$ –Ne and  $CH_3^{35}ClH^+$ –Ne are isobaric species (m/z = 71), both ions were present in the ion trap, and hence, the corresponding spectrum comprises the spectra of both ions superimposed. The center trace in Figure 5 shows the final spectrum obtained in the wavenumber regions from 500 to  $1530 \text{ cm}^{-1}$  and from 2625 to  $2740 \text{ cm}^{-1}$ . The two regions feature no less than seven vibrational bands, all but one of which can be assigned in a straightforward fashion based on the CCSD(T)/cc-pV(T + d)Z force field calculations. The two vibrational bands located at 1037 cm<sup>-1</sup> and 1125 cm<sup>-1</sup> are assigned as the strong  $v_3$ and  $\nu_4$  vibrational fundamentals of CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup> (cf. Table 3). By comparison with the results from the CCSD(T)/cc-pV(T + d)Z force field calculations collected in Table 4, five bands are assigned as to originate from CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup>, marking the first spectroscopic detection of this molecular ion.  $CH_3ClH^+$  has a total of 3N - 6 = 12 vibrational fundamentals, four of which are lying outside of the spectral regions covered here, namely, the three C-H stretching modes  $v_1$  (calculated at 3085 cm<sup>-1</sup>),  $v_2$  (2979 cm<sup>-1</sup>), and  $v_9$  (3101 cm<sup>-1</sup>), as well as the low-energy torsional mode  $v_{12}$  (177 cm<sup>-1</sup>). Most other fundamentals are clearly identified, i.e., the Cl–C stretching mode  $v_8$  at 551 cm<sup>-1</sup>, the C–Cl–H bending mode  $v_7$  at 766 cm<sup>-1</sup>, the CH<sub>3</sub> in-plane rocking mode  $v_6$  at 1083 cm<sup>-1</sup>, and the CH<sub>3</sub> umbrella mode  $v_5$  at 1345 cm<sup>-1</sup>. According to the calculations, the strong feature at  $1420 \text{ cm}^{-1}$  is from an overlap of the energetically nearly degenerate CH<sub>3</sub> asymmetric bending modes  $\nu_4(a')$  and  $\nu_{10}(a'')$ . The vibrational band with the largest calculated infrared intensity (Table 4) and indeed the strongest feature in the spectrum covered here is the Cl–H stretching mode  $\nu_3$  detected at 2677 cm<sup>-1</sup>. The weak out-of-plane rocking mode  $v_{11}$  is overlapping with the much

stronger  $v_3$  band of CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup>. Overall, the agreement between the IRPD spectrum and the anharmonic CCSD(T)/cc-pV(T + *d*)Z force field calculation is excellent. The harmonic force field calculations of the bare ions and the singly tagged variants in Table 2 suggest that the influences of Ne should be subtle for most bands but noticeable for some. For the bands covered here, significant shifts might be expected for the  $v_3$  (redshift),  $v_7$  (blueshift), and, possibly,  $v_8$  bands (blueshift) if Ne was connected to the site of protonation. Indeed, the experimental findings are qualitatively consistent with this geometrical arrangement, as the largest differences  $\delta$  between experimental and calculated wavenumbers of all bands observed are found for these three modes that also agree with the predicted sign of the shift. Lastly, as seen from the inset in Figure 5 (blue center trace), similarly to the  $v_3$  and  $v_4$  bands in CH<sub>2</sub>Cl<sup>+</sup>, the  $v_8$  and  $v_7$  bands (possibly also  $v_3$ ) show weak satellite bands about 15 cm<sup>-1</sup> blueshifted from the fundamentals, presumably combination modes involving low-energy fundamentals introduced through Ne-tagging (most likely the  $\omega_{15}$  mode, as shown in Table 2).

**Table 4.** Harmonic and anharmonic fc-CCSD(T)/cc-pV(T + *d*)Z vibrational wavenumbers (in cm<sup>-1</sup>) and infrared band intensities (int, km/mol), experimental wavenumbers (exp, cm<sup>-1</sup>), full width at half maximum of observed bands (FWHM, cm<sup>-1</sup>), and difference between observed and calculated wavenumbers  $\delta$  (cm<sup>-1</sup>) of CH<sub>3</sub>ClH<sup>+</sup>.

M. 1.	CH <sub>3</sub> <sup>35</sup> ClH <sup>+</sup>					
Mode –	Harm	Anharm	Int	Exp <sup>1</sup>	FWHM	δ
$\nu_1(a')$ C–H asym stretching	3225.4	3084.9	16	пс		
$v_2(a')$ C–H sym stretching	3091.6	2978.8	5	пс		
$v_3(a')$ Cl–H stretching	2829.5	2713.5	173	2677	15	-37
$v_4(a')$ CH <sub>3</sub> asym bending	1465.1	1420.6	15	1415	10	-6
$v_5(a')$ CH <sub>3</sub> umbrella	1378.2	1348.1	10	1345	8	-3
$v_6(a')$ CH <sub>3</sub> ip rocking	1113.6	1080.8	3	1083	9	2
$v_7(a')$ C-Cl-H bending	768.4	757.1	14	766	9	9
$v_8(a')$ Cl–C stretching	569.8	544.1	55	551	11	7
$v_9(a'')$ C–H asym stretching	3242.9	3100.6	14	пс		
$v_{10}(a'')$ CH <sub>3</sub> asym bending	1462.5	1419.6	15	1415	10	-5
$v_{11}(a'')$ CH <sub>3</sub> oop rocking	1033.2	1020.6	1	blended		
$v_{12}(a'')$ torsion	205.5	177.4	45	пс		
	CH <sub>3</sub> <sup>37</sup> ClH <sup>+</sup>					
Mode –	Harm	Anharm	Int	Exp <sup>1</sup>	FWHM	δ
$\nu_1(a')$ C–H asym stretching	3225.4	3084.9	16	пс		
$v_2(a')$ C–H sym stretching	3091.6	2978.8	5	пс		
$v_3(a')$ Cl–H stretching	2827.3	2711.5	173	2678	20	-34
$v_4(a')$ CH <sub>3</sub> asym bending	1465.1	1420.5	15	1416	13	-5
$v_5(a')$ CH <sub>3</sub> umbrella	1377.9	1347.9	10	1344	11	-4
$v_6(a')$ CH <sub>3</sub> ip rocking	1112.7	1079.9	3	1082	11	2
$v_7(a')$ C-Cl-H bending	768.1	756.8	14	765	9	8
$v_8(a')$ Cl–C stretching	565.3	539.9	56	547	11	7
$v_9(a'')$ C–H asym stretching	3242.9	3100.6	14	пс		
				1416	10	4
$\nu_{10}(a'')$ CH <sub>3</sub> asym bending	1462.5	1419.6	15	1416	13	-4
$v_{10}(a'')$ CH <sub>3</sub> asym bending $v_{11}(a'')$ CH <sub>3</sub> oop rocking	1462.5 1032.8	1419.6 1020.2	15 1	1416 1023	13 12	$^{-4}_{3}$

<sup>1</sup> Wavenumber errors amount to 0.5% of the band position; *nc*: not covered.

2.4. IRPD Spectrum of CH<sub>3</sub><sup>37</sup>ClH<sup>+</sup>-Ne

The IRPD spectrum of the m/z = 73 species covering the 500 to  $1630 \text{ cm}^{-1}$  and 2505 to 2880 cm<sup>-1</sup> regions is shown in Figure 5 (green bottom trace). Clearly, the strong vibrational band quintett comprising the  $v_8$ ,  $v_7$ ,  $v_5$ ,  $v_{10}/v_4$ , and  $v_3$  bands nicely matches the pattern observed also for the CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup>–Ne isotopic species at m/z = 71. Likewise, the  $v_8$  and  $v_7$  bands of CH<sub>3</sub><sup>37</sup>ClH<sup>+</sup>–Ne show weaker yet clearly discernible shoulders to their blue sides. IRPD scans of the strong  $v_8$  band reveal that depletion is nearly 100%, suggesting that the

m/z = 73 spectrum should provide a clean view of the CH<sub>3</sub><sup>37</sup>ClH<sup>+</sup>–Ne species and not be affected by isobaric contamination. As a consequence, in addition to the weak CH<sub>3</sub> ip rocking mode  $\nu_6$  band at 1082 cm<sup>-1</sup>, now also the very weak CH<sub>3</sub> out-of-plane rocking mode  $\nu_{11}$  can be identified at 1023 cm<sup>-1</sup>. In addition to these vibrational fundamentals, the spectrum shows a spectroscopic feature at 1461 cm<sup>-1</sup> that is also seen as a somewhat less prominent signal in the m/z = 71 spectrum (see bands marked with asterisks in Figure 5). From the anharmonic force field calculations, this feature is not compatible with a binary but, possibly, a ternary combination mode comprising the three energetically lowest fundamental modes,  $\nu_7 + \nu_8 + \nu_{12}$ . However, from the present investigation, a definitive assignment is not feasible.

#### 3. Materials and Methods

#### 3.1. Quantum-Chemical Calculations

Theoretical molecular structures of both  $CH_2Cl^+$  ( $C_{2\nu}$ ) and  $CH_3ClH^+$  ( $C_s$ ) have previously been provided on several occasions (e.g., Refs. [12,22-25]). In the present study, complementary quantum-chemical calculations of  $CH_2Cl^+$  and  $CH_3ClH^+$  (as well as their isostructural and isoelectronic sulfur analogs, H<sub>2</sub>CS and CH<sub>3</sub>SH) were performed at the coupled cluster single and double (CCSD) level augmented by a perturbative treatment of triple excitations (CCSD(T)) [26], together with correlation-consistent (augmented) polarized valence and (augmented) polarized weighted core-valence basis sets, i.e., cc-pVXZ [27], aug-ccpVXZ [27–29], and cc-pwCVXZ [27,30] (with X = T, Q). For basis sets denoted as cc-pV(X + d)Zor aug-cc-pV(X + d)Z, an additional tight d function [31] was added to the chlorine (and sulfur) atom only, while for all other elements, cc-pV(X)Z or aug-cc-pV(X)Z was used, respectively. Equilibrium geometries were calculated using analytic gradient techniques [32], while harmonic frequencies were computed using analytic second-derivative techniques [33,34]. For anharmonic computations, second-order vibrational perturbation theory (VPT2) [35] was employed, and an additional numerical differentiation of analytic second derivatives was applied to obtain the third and fourth derivatives required for the application of VPT2 [34,36]. The frozen core approximation is indicated throughout with "fc", and "ae" indicates that all electrons were used in the correlation treatment. All calculations were carried out using the CFOUR program package [37,38].

#### 3.2. Experiment

Experimental infrared action spectroscopic characterization of CH<sub>2</sub>Cl<sup>+</sup> and CH<sub>3</sub>ClH<sup>+</sup> was performed in selected wavenumber ranges from 500 to 2900 cm<sup>-1</sup> using a cryogenic 22-pole ion trap apparatus FELion connected to the Free Electron Laser for Infrared eXperiments (FELIX) [39], located at Radboud University (Nijmegen, the Netherlands). The FELion apparatus and its central part, the 22-pole ion trap, have been described elsewhere in detail [40,41], and the experimental conditions applied were similar to those used in other recent studies of molecular ions [16,17,42–44]. CH<sub>2</sub><sup>35,37</sup>Cl<sup>+</sup> (m/z = 49,51) and CH<sub>3</sub><sup>35,37</sup>ClH<sup>+</sup> (m/z = 51,53) were produced in an ion source under the same experimental conditions from commercially available methyl chloride, CH<sub>3</sub>Cl, using electron impact ionization at an electron energy of about 30 eV. Because of the high natural abundance of <sup>37</sup>Cl (24.24%), the corresponding heavy isotopologs CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup> and CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup> were easily accessible. It should be stressed that CH<sub>2</sub><sup>37</sup>Cl<sup>+</sup> and CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup> are isobaric species (m/z = 51); hence, the corresponding spectra were expected to appear superimposed in the same scan.

After extracting the ion cloud from the source and selection in the first quadrupole mass analyzer, the m/z of choice was introduced into the 22-pole ion trap, where it was cooled via a cold pulse of a 3:1 mixture of He:Ne kept at a nominal temperature of 8 K and at a high number density, leading to efficient tagging with Ne atoms. Per filling cycle of the ion trap, typically, several thousand cation–Ne complexes (CH<sub>2</sub><sup>35,37</sup>Cl<sup>+</sup>–Ne at m/z = 69/71 or CH<sub>3</sub><sup>35,37</sup>ClH<sup>+</sup>–Ne at m/z = 71,73) can be formed via three-body collisions (Figure A1 in Appendix C). After a selected storage time of typically 1.6 or 2.6 s, the trap content was

extracted from the trap, mass-filtered in a second quadrupole mass analyzer, and counted using a very sensitive Daly-type detector.

In the infrared photodissociation (IRPD) action spectroscopy scheme employed herein, the number of singly tagged  $CH_2Cl^+$ -Ne cluster ions (similar for  $CH_3ClH^+$ -Ne) was monitored while the FELIX (FEL-2) IR radiation traversing the ion trap was tuned in wavenumber. FELIX was operated in a pulsed mode (10 Hz), with typical pulse energies of a few up to a few tens of mJ (measured at the exit of the ion trap) and a Fourier-limited full width at half maximum (FWHM) bandwidth on the order of 0.7%. When a vibrational mode of the cluster and the radiation source are coincident in wavenumber, dissociation of the cluster occurs, resulting in depletion in the ion–Ne counts. Finally, several power-normalized spectra were co-added to obtain the final spectrum. Details on the normalization procedure have been given in Ref. [41].

Since the weakly bound, rare gas Ne generally only slightly perturbs the structure of the ion, an IRPD spectrum is highly representative of that of the bare ion (e.g., Refs. [16,17,21,44]). Band parameters such as positions and FWHM were obtained by fitting (multi-component) Gaussian profiles to the experimental spectra. Typical wavenumber uncertainties of 0.5% are mainly dependent on calibration uncertainties due to the grating spectrum analyzer.

#### 4. Conclusions

The present study reports on the first infrared gas-phase study of  $CH_2Cl^+$  and the first spectroscopic detection and characterization of CH<sub>3</sub>ClH<sup>+</sup>, two fundamental chlorinebearing ions possibly linked to the astrochemical gas-phase production of methyl chloride, CH<sub>3</sub>Cl. Now that the IR spectra of  $CH_2Cl^+$  and  $CH_3ClH^+$  have been detected at low spectral resolution, corresponding studies at high resolution are imperative, in particular, to permit radio astronomical searches. Following the very recent technical developments in action spectroscopy of molecular ions, high-resolution infrared spectroscopy of both species may be performed using leak-out spectroscopy [45], a new method that does not require any rare gas tagging and has already been shown, on several occasions, to yield spectra of superb signal-to-noise ratios [46–49]. The method may even be combined with millimeterwave radiation in a double-resonance fashion to derive accurate information about the pure rotational spectrum [46,47,50], a prerequisite for radio astronomy. Using the high-level CCSD(T) structural and force field calculations performed here, very good predictions of the ground state rotational and centrifugal distortion constants can be derived. These estimates may even be further improved using a simple scaling procedure, if experimental data of isostructural/isoelectronic species are available (see, e.g., Refs. [16,50,51]). As the Cl<sup>+</sup> ion is isoelectronic with neutral atomic S, thioformaldehyde, H<sub>2</sub>CS, and methyl mercaptan, CH<sub>3</sub>SH, may potentially be used as calibrators, the pure rotational spectra of which are very well known from previous microwave and millimeter-wave studies (see, e.g., Refs. [52,53] for recent reports). Calculated and experimental rotational constants of H<sub>2</sub>CS and CH<sub>3</sub>SH (structural parameters are given in Appendix B) as well as calculated constants of  $CH_2Cl^+$ and  $CH_3ClH^+$  are collected in Table 5. For the sulfur species, the agreement between the calculated and experimental ground state rotational constants is already very good. Consequently, the scaling factors, i.e., the ratios  $A_{0,exp}/A_{0,calc}$ ,  $B_{0,exp}/B_{0,calc}$ , and  $C_{0,exp}/C_{0,calc}$ are almost but not quite unity. Multiplying those factors with the calculated ground state rotational constants of CH<sub>2</sub>Cl<sup>+</sup> and CH<sub>3</sub>ClH<sup>+</sup> finally yields the scaled best-estimate rotational constants in the last column of Table 5. Both species are very polar with centerof-mass frame equilibrium dipole moment components  $\mu_a = 3.02 \text{ D}$  of CH<sub>2</sub>Cl<sup>+</sup> as well as  $\mu_a = 1.22 \text{ D}$  and  $\mu_b = 1.42 \text{ D}$  of CH<sub>3</sub>ClH<sup>+</sup>.

From a chemical viewpoint, extending the spectroscopy of  $CH_2X^+$  and  $CH_3XH^+$  to halogens other than chlorine also seems appealing. The situation for X=Br, for example, appears similar to the one with X=Cl prior to the present study. To date, spectroscopic studies of  $CH_2Br^+$  have been restricted to infrared matrix isolation [12,54]. Protonated methyl bromide,  $CH_3BrH^+$ , seems not to have been spectroscopically studied to date.

Demonster	H <sub>2</sub>	CH2 <sup>35</sup> Cl+		
rarameter	Calc.	Exp. [52]	Calc.	Scaled
A <sub>e</sub>	295,043.772		278,232.178	
B <sub>e</sub>	17,770.906		17,929.887	
Ce	16,761.346		16 <i>,</i> 844.397	
$\Delta A_0$	1928.402		2249.656	
$\Delta B_0$	76.366		91.765	
$\Delta C_0$	109.739		125.604	
$A_0$	293,115.370	291,613.339	275,982.522	274,568.285
<i>B</i> <sub>0</sub>	17,694.540	17,698.994	17,838.122	17,842.613
$C_0$	16,651.607	16,652.498	16,718.793	16,719.687
Demonstra	СН	CH <sub>3</sub> SH		
Parameter	Calc.	<b>Exp.</b> [53] <sup>2</sup>	Calc.	Scaled
A <sub>e</sub>	103,863.953		103,144.040	
B <sub>e</sub>	13,029.847		12,472.858	
Ce	12,493.723		12,016.948	
$\Delta A_0$	1054.186		1069.132	
$\Delta B_0$	131.546		207.375	
$\Delta C_0$	117.267		184.225	
$A_0$	102,809.767	102,767.147	102,074.907	102,032.592
<i>B</i> <sub>0</sub>	12,898.301	12,951.372	12,265.483	12,315.950
$C_0$	12,376,457	12.388.033	11.832.724	11.843.792

Table 5. Rotational parameters of H<sub>2</sub>CS, CH<sub>2</sub><sup>35</sup>Cl<sup>+</sup>, CH<sub>3</sub>SH, and CH<sub>3</sub><sup>35</sup>ClH<sup>+</sup> (in MHz)<sup>1</sup>.

<sup>1</sup> Equilibrium rotational constants and zero-point vibrational contributions calculated at the ae-CCSD(T)/ccpwCVQZ and fc-CCSD(T)/cc-pV(T + d)Z levels of theory, respectively. <sup>2</sup> For the sake of comparability, the rotational constants from Ref. [53] given in the rho-axis system were converted to the principal axis system. Differences are virtually negligible.

**Author Contributions:** Conceptualization, S.T., P.C.S., O.A. and S.S.; software, S.T. and L.B.; formal analysis, S.T., T.D. and L.B.; investigation, S.T., K.S., T.D., P.C.S., and S.B.; writing—original draft preparation, S.T.; writing—review and editing, S.T., K.S., T.D., P.C.S., L.B., O.A., S.B. and S.S.; visualization, S.T. and L.B.; project administration, S.T., K.S. and S.B.; funding acquisition, S.T., O.A. and S.S. All authors have read and agreed to the published version of the manuscript.

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Conflicts of Interest: The authors declare no conflicts of interest.

# Appendix A. Molecular Structures of $\rm CH_2Cl^+$ and $\rm CH_3ClH^+$ and Their Weakly Bound Clusters with Neon

```
Appendix A.1. CH<sub>2</sub>Cl<sup>+</sup>
CL
C 1 r1
H 2 r2 1 a1
H 2 r2 1 a1 3 d180
       cc-pV(T+d)Z / aug-cc-pV(T+d)Z / cc-pwCVQZ
                  / 1.59662
                                  / 1.58812
r1
     =
         1.59619
                                   / 1.08620
                  / 1.08860
r2
   =
         1.08829
    = 119.09851 / 119.03255
                                    / 119.07897
a1
d180 = 180.00000 / 180.00000
                                   / 180.00000
Appendix A.2. CH_2Cl^+ Ne-ip
CL
C 1 r1
H 2 r2 1 a1
H 2 r3 1 a2 3 d180
NE 4 r4 2 a3 3 d180
      aug-cc-pV(T+d)Z
r1
     =
          1.59672
r2
          1.08853
   =
    = 119.01368
a1
r3
    =
          1.08827
   = 118.82770
a2
d180 = 180.00000
r4 =
          2.39056
a3 = 133.32909
Appendix A.3. CH_2Cl^+ Ne-oop
CL
C 1 r1
NE 2 rNE 1 aNE
H 2 r2 3 a1 1 d1
H 2 r2 3 a1 1 md1
      aug-cc-pV(T+d)Z
r1
     =
          1.59691
rNE =
          2.79264
aNE = 108.25118
        1.08829
r2 =
    = 81.35078
a1
    = 117.85276
d1
md1 = -117.85276
Appendix A.4. CH<sub>3</sub>ClH<sup>+</sup>
Η
C 1 r1
CL 2 r2 1 a1
H 3 r3 2 a2 1 d180
H 2 r4 3 a3 1 d1
```

H 2 r4 3 a3 1 md1 cc-pV(T+d)Z / aug-cc-pV(T+d)Z / cc-pwCVQZ r1 1.08652 / 1.08729 / 1.08435 = r2 1.85970 / 1.86174 / 1.84696 = = 101.97978 / 101.91486 / 102.10778 a1 r3 1.29569 / 1.29687 / 1.29436 = a2 = 99.43628 / 99.36597 / 99.64973 d180 = 180.00000 / 180.00000 / 180.00000 / 1.08443 / r4 = 1.08514 1.08215 = 105.70626 / 105.53383 / 105.84821 a3 = 118.70156 / 118.73149 / 118.68039 d1 md1 = -118.70156 / -118.73149 / -118.68039 Appendix A.5. CH<sub>3</sub>ClH<sup>+</sup>–Ne, p-bound Н C 1 r1 CL 2 r2 1 a1 H 3 r3 2 a2 1 d180 H 2 r4 3 a3 1 d1 H 2 r4 3 a3 1 md1 X 4 rd 3 a90 2 d180 NE 4 r5 7 a4 3 d180 r1 = 1.08720 r2 = 1.85987 a1 = 102.05144 r3 = 1.29827 a2 = 99.19249 d180 = 180.00000 r4 = 1.08515 a3 = 105.57033 118.75594 d1 = md1 = -118.75594 rd 1.00000 = a90 = 90.00000 2.07322 r5 = a4 = 94.38835 Appendix A.6. CH<sub>3</sub>ClH<sup>+</sup>-Ne, CH<sub>3</sub>/Cl-bound Н C 1 r1 CL 2 r2 1 a1 H 3 r3 2 a2 1 d180 H 2 r4 3 a3 1 d1 H 2 r4 3 a3 1 md1 X 1 rd 2 a90 3 d0 NE 1 r5 7 a4 2 d180 r1 = 1.08692 r2 = 1.86155 a1 = 101.81875 r3 = 1.29674 a2 = 99.27846

d180	=	180.00000
r4	=	1.08514
a3	=	105.52239
d1	=	118.75328
md1	=	-118.75328
rd	=	1.00000
a90	=	90.00000
d0	=	0.00000
r5	=	2.57934
a4	=	40,24330

### Appendix B. Molecular Structures of H<sub>2</sub>CS and CH<sub>3</sub>SH

```
Appendix B.1. H<sub>2</sub>CS
S
C 1 r1
H 2 r2 1 a1
H 2 r2 1 a1 3 d180
            cc-pV(T+d)Z / cc-pwCVQZ
               1.61826 / 1.60890
r1
     =
r2
    =
                1.08766 /
                          1.08531
              121.92777 / 121.85524
    =
a1
d180 =
              180.00000 / 180.00000
Appendix B.2. CH<sub>3</sub>SH
Η
C 1 r1
S 2 r2 1 a1
H 3 r3 2 a2 1 d180
H 2 r4 3 a3 1 d1
H 2 r4 3 a3 1 md1
            cc-pV(T+d)Z / cc-pwCVQZ
               1.08915 /
r1
     =
                             1.08669
                1.82085 /
r2
    =
                             1.81020
              106.24841 / 106.25194
a1
     =
r3 =
               1.33803 /
                           1.33502
a2
   =
              96.64050 /
                          96.88414
             180.00000 / 180.00000
d180 =
               1.08804 /
r4
    =
                            1.08554
            111.27155 / 111.25848
a3 =
d1
    =
             118.26139 / 118.25357
           -118.26139 / -118.25357
md1 =
```



## Appendix C. Mass Spectrum of CH<sub>2</sub>Cl<sup>+</sup> and CH<sub>2</sub>Cl<sup>+</sup>–Ne as Extracted from the Ion Trap

**Figure A1.** Trap content extracted after initial mass selection for m/z = 49, trapping and cooling in the 8K He/Ne-bath of the 22-pole ion trap featuring some 35,000 CH<sub>2</sub>Cl<sup>+</sup> ions and almost 7000 weakly bound clusters with <sup>20</sup>Ne available for IRPD. Black trace and left ordinate: linear plot of counts; orange dotted trace and right ordinate: logarithmic plot of counts

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