Multidisciplinary Scientific Journal

## Article

# The Random Gas of Hard Spheres 

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Received: 19 April 2019; Accepted: 28 May 2019; Published: 31 May 2019


#### Abstract

The inconsistency between the time-reversible Liouville equation and time-irreversible Boltzmann equation has been pointed out by Loschmidt. To avoid Loschmidt's objection, here we propose a new dynamical system to model the motion of atoms of gas, with their interactions triggered by a random point process. Despite being random, this model can approximate the collision dynamics of rigid spheres via adjustable parameters. We compute the exact statistical steady state of the system, and determine the form of its marginal distributions for a large number of spheres. We find that the Kullback-Leibler entropy (a generalization of the conventional Boltzmann entropy) of the full system of random gas spheres is a non-increasing function of time. Unlike the conventional hard sphere model, the proposed random gas system results in a variant of the Enskog equation, which is known to be a more accurate model of dense gas than the Boltzmann equation. We examine the hydrodynamic limit of the derived Enskog equation for spheres of constant mass density, and find that the corresponding Enskog-Euler and Enskog-Navier-Stokes equations acquire additional effects in both the advective and viscous terms.


Keywords: hard sphere gas; Enskog equation; Boltzmann equation; Loschmidt's paradox

## 1. Introduction

It is known that the atoms in an electrostatically neutral monatomic gas interact via the Lennard-Jones potential [1]. At short range, the Lennard-Jones potential is inversely proportional to the 12th power of the distance between the centers of two interacting atoms. Due to this high order singularity, the mean field approximation [2], which is often used to close the Bogoliubov-Born-Green-Kirkwood-Yvon (BBGKY) hierarchy [3-5] for the long-range (that is, electrostatic or gravitational) potentials, cannot be applied to the molecular collisions, since the relevant spatial integrals diverge as the distance between the atoms approaches zero. To circumvent this issue, in molecular kinetic theory, the Lennard-Jones potential interaction was replaced with the hard sphere collision model [6-11]. According to this model, the gas molecules are presumed to be rigid impenetrable spheres, which interact instantaneously and elastically according to the mechanics of the collision of two hypothetical billiard balls, which have nonzero mass, but do not possess the moment of inertia.

In the conventional setting of the hard sphere gas dynamics, one starts with the Liouville equation $[7-9,11]$ for the probability density function of the complete system of all interacting spheres. The Liouville equation is a homogeneous transport equation, whose characteristics are the straight lines of free flight of the spheres. To describe the collisions, the Liouville equation is endowed with deflection conditions at the collision surface (the set of points in the coordinate space where a pair of spheres is separated by their diameter). These deflection conditions restrict the solutions of the Liouville equation to those for which the probability of a pair of spheres entering the collision surface equals the probability of a pair of spheres simultaneously exiting the collision surface along the corresponding direction of deflection. Assuming that the Liouville equation is already solved with the solution being
known, one then computes the corresponding BBGKY hierarchy [3-5]. The lowest-order identity in the BBGKY hierarchy is then converted to the Boltzmann equation [6-13] via an approximation known as the Boltzmann hierarchy, and the subsequent factorization of the joint probability density of two spheres into the product of single-sphere densities.

Upon the examination of the standard closure of hard sphere dynamics [8,9,11], we observe that the derivation of the Boltzmann equation from the leading order BBGKY identity violates the assumptions under which the BBGKY identity itself was derived. The observed contradictions are similar to those pointed out by Loschmidt [14]. However, the Boltzmann equation is known in practice to yield an accurate approximation to the observable gas dynamics. Thus, we propose that the Boltzmann equation instead originates, in a consistent manner, from a different model of hard sphere dynamics, which, on the one hand, does not violate Loschmidt's observation, but, on the other hand, is on a par with the hard sphere model at approximating real molecular interactions.

To avoid Loschmidt's objection [14] and derive the Boltzmann equation in a consistent fashion, we propose a new random process to model the underlying hard sphere dynamics, where the changes in the velocities of the spheres still obey the mechanics of the rigid sphere collision, but the "collisions" themselves are triggered by a point process. This random dynamical system possesses the infinitesimal generator, so that the corresponding forward Kolmogorov equation for its probability density is readily available via the integration by parts. The proposed random hard sphere process can approximate the deterministic collision dynamics by increasing the intensity of the triggering point process via a parameter. We compute some of the steady states of the proposed process, and find that, while the conventional Boltzmann entropy can both increase and decrease, depending on a solution, the Kullback-Leibler entropy [15] between a solution and the steady state is a non-increasing function of time. We also examine the structure of marginal distributions of the steady state in the limit of infinitely many spheres.

Then, we compute the forward equation for a single-sphere marginal distribution, under the assumption that the corresponding multisphere probability density is invariant under an arbitrary reordering of spheres. The closure that we use is, however, different from conventional-rather than using a direct factorization of the probability state (used, for example, in the conventional transition from the Boltzmann hierarchy to the Boltzmann equation), we take into account the structure of the previously computed steady state of the full multisphere system. We then find that, in the limit as the intensity of the point process increases to infinity, a variant of the Enskog equation [16-21] emerges, which is a more accurate model of a dense gas [22] than the conventional Boltzmann equation. We then examine the hydrodynamic limit of the Enskog equation for spheres of constant mass density, as it appears to be physically plausible for atoms of the noble gases [23]. In this limit, we find that the resulting Enskog-Euler and Enskog-Navier-Stokes equations acquire additional non-vanishing terms, which are not present in the conventional gas dynamics equations originating from the Boltzmann equation $[7-10,13]$. These additional effects disappear in the dilute gas approximation, which yields the usual Boltzmann, Euler, and Navier-Stokes equations, respectively.

## 2. The Hard Sphere Collision Model and the Boltzmann Equation

For the sake of clarity of the exposition, we start by citing the standard derivation of the Boltzmann equation from the collision mechanics of hard spheres [7-9,11]. For simplicity, we consider only two identical hard spheres, each of diameter $\sigma$. We denote by $x$ and $v$ the coordinate of the center and velocity of the first sphere, respectively, and by $\boldsymbol{y}$ and $\boldsymbol{w}$ the corresponding coordinate of the center and velocity of the second sphere. In the absence of contact, the spheres maintain their constant velocities $v$ and $w$, while their respective coordinates $x$ and $y$ are given via

$$
\begin{equation*}
\frac{\mathrm{d} x}{\mathrm{~d} t}=v, \quad \frac{\mathrm{~d} y}{\mathrm{~d} t}=w . \tag{1}
\end{equation*}
$$

Whenever the distance $\|x-y\|$ between the centers of the two spheres equals their diameter $\sigma$, their velocities are changed, at that instance of time, to

$$
\begin{equation*}
v^{\prime}=v+(w-v) \cdot(x-y) \frac{x-y}{\|x-y\|^{2}}, \quad w^{\prime}=w+(v-w) \cdot(x-y) \frac{x-y}{\|x-y\|^{2}} \tag{2}
\end{equation*}
$$

where $v^{\prime}$ and $w^{\prime}$ are the new values of velocities. Such a transformation preserves the momentum and kinetic energy of the system of the two spheres:

$$
\begin{equation*}
v^{\prime}+w^{\prime}=v+w, \quad\left\|v^{\prime}\right\|^{2}+\left\|w^{\prime}\right\|^{2}=\|v\|^{2}+\|w\|^{2} \tag{3}
\end{equation*}
$$

Here, and below, we assume that the total momentum of the system (the sum of velocities of the spheres) is zero without loss of generality, as otherwise the momentum can be set to zero via a suitable Galilean shift of the reference frame.

The transformation above in Equation (2) is fully symmetric; indeed, subtracting the first relation from the second, we obtain

$$
\begin{equation*}
v^{\prime}-w^{\prime}=v-w+2(w-v) \cdot(x-y) \frac{x-y}{\|x-y\|^{2}} \tag{4}
\end{equation*}
$$

Scalar-multiplying by $(\boldsymbol{x}-\boldsymbol{y}) /\|x-y\|$ on both sides, we further obtain

$$
\begin{equation*}
\left(v^{\prime}-w^{\prime}\right) \cdot \frac{x-y}{\|x-y\|}=-(v-w) \cdot \frac{x-y}{\|x-y\|} \tag{5}
\end{equation*}
$$

Substituting the above expression into Equation (2), we arrive at

$$
\begin{equation*}
v=v^{\prime}+\left(w^{\prime}-v^{\prime}\right) \cdot(x-y) \frac{x-y}{\|x-y\|^{2}}, \quad w=w^{\prime}+\left(v^{\prime}-w^{\prime}\right) \cdot(x-y) \frac{x-y}{\|x-y\|^{2}} \tag{6}
\end{equation*}
$$

It is also easy to see that the Jacobian of the change of variables $(v, w) \rightarrow\left(v^{\prime}, w^{\prime}\right)$ is unity; indeed, observe that, for $(\boldsymbol{x}-\boldsymbol{y})$ taken as a fixed parameter,

$$
\begin{equation*}
\operatorname{det}\left(\frac{\partial\left(v^{\prime}, w^{\prime}\right)}{\partial(v, w)}\right)=-1 \tag{7}
\end{equation*}
$$

which can be verified via the rank-one update lemma for determinants.

### 2.1. The Liouville Problem for Two Spheres

Let $F(t, x, y, v, w)$ denote the probability density function of the system of two spheres. Then, $F$ satisfies the following conditions:

- If the distance between the spheres is less than their diameter (that is, the spheres are overlapped), then the density is set to zero:

$$
\begin{equation*}
F(x, y, v, w)=0, \quad\|x-y\|<\sigma \tag{8}
\end{equation*}
$$

This condition ascertains that the spheres are rigid and may not overlap.

- If the distance between the spheres is greater than their diameter (that is, the spheres are separated), then $F$ obeys the Liouville equation [7-9]:

$$
\begin{equation*}
\frac{\partial F}{\partial t}+v \cdot \frac{\partial F}{\partial x}+w \cdot \frac{\partial F}{\partial y}=0, \quad\|x-y\|>\sigma \tag{9}
\end{equation*}
$$

- The Liouville equation (Equation (9)) is solved in the open set $\|x-y\|>\sigma$-therefore, a boundary condition is needed at the collision surface $\|x-y\|=\sigma$. This boundary condition is given via

$$
\begin{equation*}
F\left(x, y, v^{\prime}, w^{\prime}\right)=F(x, y, v, w), \quad\|x-y\|=\sigma \tag{10}
\end{equation*}
$$

where $v^{\prime}$ and $w^{\prime}$ are given in Equation (2) as functions of $x, y, v$ and $w$.

- Observe that in the open set $\|x-y\|<\sigma, F=0$ also solves the Liouville equation in Equation (9)—albeit with zero initial and boundary conditions. This leads to a possible discontinuity of $F$ at the collision surface $\|x-y\|=\sigma$, which is taken into account below.
- Assuming that there is a solution of Equation (9) in both open regions $\|x-y\|<\sigma$ and $\|\boldsymbol{x}-\boldsymbol{y}\|>\sigma$ (with boundary conditions given via zero and Equation (10), respectively), on the collision surface $\|x-y\|=\sigma$ itself we assign $F$ to be equal to the "outer" boundary condition in Equation (10). This definition of $F$ at the collision surface means that $F$ is continuous as $\|x-y\| \downarrow \sigma$, and possibly discontinuous as $\|\boldsymbol{x}-\boldsymbol{y}\| \uparrow \sigma$.

In what follows, we assume for convenience that the spatial part of the domain has a finite volume, but no boundary effects other than those in Equation (10). For example, one can assume that the space of coordinates is periodic (that is, if a sphere leaves the coordinate "box" through a wall, it immediately re-enters from the opposite wall).

According to Equation (10), the probability that a pair of spheres enters the collision surface at the point $(x, y)$ and time $t$ with velocities $(v, w)$ is the same as that of a pair of spheres exiting at the same point and time, but with the velocities $\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$. The condition in Equation (10) preserves the normalization of $F$ separately in the region $\|x-y\|>\sigma$.

Note that the formulation of the problem in Equations (8)-(10) is not fully equivalent to the actual dynamics in Equations (1) and (2). For the formulation in Equations (8)-(10) to correspond precisely to Equations (1) and (2), one needs the characteristic curves of the Liouville equation in Equation (9) to be the trajectories of the system in Equations (1) and (2). However, this is not the case in Equations (8)-(10)-the characteristic curves are straight lines given via constant parameters $v$ and $w$, which pierce the collision surface $\|\boldsymbol{x}-\boldsymbol{y}\|=\sigma$. Instead, the "collisions" in Equation (8)-(10) are implemented via reassigning the values of $F$ between different characteristics according to Equation (10).

### 2.2. The BBGKY Identity for Two Spheres

Here, we follow Cercignani et al. [9] and derive the BBGKY hierarchy (which, for two spheres, consists of a single identity) and, subsequently, the Boltzmann equation. In what follows, we take advantage of the fact that $F=0$ inside the overlapped region $\|\boldsymbol{x}-\boldsymbol{y}\|<\sigma$ also satisfies the Liouville equation in Equation (9) with zero initial and boundary conditions.

Let us assume that a solution $F$ is computed for the Liouville problem in Equations (8)-(10) for both overlapped and non-overlapped regions, so that the relation in Equation (9) is no longer an equation, but rather an identity. In what follows, we assume that $F$ is symmetric under the permutations of two spheres-that is, $F(t, x, y, v, w)=F(t, y, x, w, v)$. Our goal is to manipulate the Liouville identity in Equation (9) so as to obtain an appropriate identity for the marginal distribution of a single sphere

$$
\begin{equation*}
f(t, x, v)=\int F(t, x, y, v, w) \mathrm{d} y \mathrm{~d} w \tag{11}
\end{equation*}
$$

First, we integrate the identity in Equation (9) in $\mathrm{d} y \mathrm{~d} w$ over the whole space, which includes both overlapped (that is, $\|x-y\|<\sigma$ ) and non-overlapped (that is, $\|x-y\|>\sigma$ ) regions:

$$
\begin{equation*}
\int\left(\frac{\partial F}{\partial t}+v \cdot \frac{\partial F}{\partial x}+w \cdot \frac{\partial F}{\partial y}\right) \mathrm{d} y \mathrm{~d} w=0 \tag{12}
\end{equation*}
$$

Next, we exchange the order of differentiation and integration above. For the $t$-derivative term, the exchange is done in a direct manner, since the $\mathrm{d} \boldsymbol{y} \mathrm{d} \boldsymbol{w}$-integration is unrelated to $t$ :

$$
\begin{equation*}
\int \frac{\partial F}{\partial t} \mathrm{~d} y \mathrm{~d} w=\frac{\partial}{\partial t} \int F \mathrm{~d} y \mathrm{~d} w=\frac{\partial f}{\partial t} \tag{13}
\end{equation*}
$$

However, due to the discontinuity in $F$ at the collision surface $\|x-y\|=\sigma$, one cannot directly swap the spatial derivatives and the spatial integration. For the term with $y$-derivative, we use the Gauss theorem:

$$
\begin{align*}
\int w & \frac{\partial F}{\partial y} \mathrm{~d} y \mathrm{~d} w=\int \mathrm{d} w\left(\int_{\|x-y\|<\sigma} \frac{\partial}{\partial y} \cdot(F w) \mathrm{d} y+\int_{\|x-y\|>\sigma} \frac{\partial}{\partial y} \cdot(F w) \mathrm{d} y\right)= \\
& =\int \mathrm{d} \boldsymbol{w} \int_{\|x-y\|=\sigma}\left(F_{<\sigma}-F_{>\sigma}\right) \boldsymbol{w} \cdot \boldsymbol{n}_{\boldsymbol{x}}(\boldsymbol{y}) \mathrm{d} S_{x}(\boldsymbol{y})=-\int \mathrm{d} \boldsymbol{w} \int_{\|x-y\|=\sigma} F w \cdot \boldsymbol{n}_{x}(\boldsymbol{y}) \mathrm{d} S_{x}(\boldsymbol{y}) . \tag{14}
\end{align*}
$$

Above, $\mathrm{d} S_{x}(\boldsymbol{y})$ is the surface area element at the point $y$ of the sphere of radius $\sigma$ centered at $x$, the unit normal vector $n_{x}(y)$, which originates at $x$ in the direction of $y$, is given via

$$
\begin{equation*}
n_{x}(y)=\frac{y-x}{\|y-x\|} \tag{15}
\end{equation*}
$$

In addition, $F_{<\sigma}$ denotes the boundary value for $\|\boldsymbol{x}-\boldsymbol{y}\|<\sigma$ (which is zero), $F_{>\sigma}$ denotes the boundary value for $\|\boldsymbol{x}-\boldsymbol{y}\|>\sigma$, given via Equation (10), and we replace $F_{>\sigma}$ with $F$ in the last identity, as it is agreed above that $F$ at the collision surface $\|x-y\|=\sigma$ is assigned the boundary value given via Equation (10).

For the term with the $x$-derivative, we need to use the generalized Leibniz rule. First, we split

$$
\begin{equation*}
\int v \cdot \frac{\partial F}{\partial x} \mathrm{~d} y \mathrm{~d} w=\int \mathrm{d} w\left(\int_{\|x-y\|>\sigma} v \cdot \frac{\partial F}{\partial x} \mathrm{~d} y+\int_{\|x-y\|<\sigma} v \cdot \frac{\partial F}{\partial x} \mathrm{~d} y\right) \tag{16}
\end{equation*}
$$

Then, for the first term in the right-hand side we write

$$
\begin{align*}
& \int_{\|x-y\|>\sigma} v \cdot \frac{\partial}{\partial x} F(x, y) \mathrm{d} y=\int_{\|z\|>\sigma} v \cdot\left(\frac{\partial}{\partial x}-\frac{\partial}{\partial z}\right) F(x, x+z) \mathrm{d} z=v \cdot \frac{\partial}{\partial x} \int_{\|z\|>\sigma} F(x, x+z) \mathrm{d} z- \\
& -\int_{\|z\|>\sigma} \frac{\partial}{\partial z} \cdot(F(x, x+z) v) \mathrm{d} z=v \cdot \frac{\partial}{\partial x} \int_{\|x-y\|>\sigma} F(x, y) \mathrm{d} y+\int_{\|x-y\|=\sigma} F_{>\sigma} v \cdot n_{x}(y) \mathrm{d} S_{x}(y) . \tag{17}
\end{align*}
$$

For the second term, we repeat the calculations in a similar way:

$$
\begin{equation*}
\int_{\|x-y\|<\sigma} v \cdot \frac{\partial}{\partial x} F(x, y) \mathrm{d} y=v \cdot \frac{\partial}{\partial x} \int_{\|x-y\|<\sigma} F(x, y) \mathrm{d} y-\int_{\|x-y\|=\sigma} F_{<\sigma} v \cdot n_{x}(y) \mathrm{d} S_{x}(y) \tag{18}
\end{equation*}
$$

Adding the two terms, recalling that $F_{<\sigma}$ is zero, $F_{>\sigma}$ is $F$ itself on $\|\boldsymbol{x}-\boldsymbol{y}\|=\sigma$, and using Equation (11), we arrive at the identity

$$
\begin{equation*}
\int v \cdot \frac{\partial F}{\partial x} \mathrm{~d} y \mathrm{~d} w=v \cdot \frac{\partial f}{\partial x}+\int \mathrm{d} w \int_{\|x-y\|=\sigma} F v \cdot n_{x}(y) \mathrm{d} S_{x}(y) . \tag{19}
\end{equation*}
$$

Combining the above expression with Equations (13) and (14), and substituting into Equation (12), we arrive at the identity

$$
\begin{equation*}
\frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=\int \mathrm{d} w \int_{\|x-y\|=\sigma} F(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}_{x}(\boldsymbol{y}) \mathrm{d} S_{x}(\boldsymbol{y}) \tag{20}
\end{equation*}
$$

For further convenience, we can rename the dummy variables of integration so that the surface integration is computed over a unit sphere (that is, over $\mathrm{d} \boldsymbol{n}$ ). Namely, since $\boldsymbol{y}$ follows the surface of the sphere of radius $\sigma$ centered at $x$, we denote

$$
\begin{equation*}
y=x+\sigma n, \quad \mathrm{~d} S_{x}(\boldsymbol{y})=\sigma^{2} \mathrm{~d} n \tag{21}
\end{equation*}
$$

In the above variables, the identity in Equation (2) becomes

$$
\begin{equation*}
v^{\prime}=v+((w-v) \cdot n) n, \quad w^{\prime}=w+((v-w) \cdot n) n \tag{22}
\end{equation*}
$$

where we note that there is no dependence on $\sigma$. The integral over the collision surface, written in the new variables, yields the following identity:

$$
\begin{equation*}
\frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=\sigma^{2} \int F(t, x, x+\sigma n, v, w)(w-v) \cdot n \mathrm{~d} n \mathrm{~d} w \tag{23}
\end{equation*}
$$

To transform the surface integral in the right-hand side above into the Boltzmann collision integral, one further needs to rearrange the surface integration with the help of Equation (10). Consider the distance $\|\boldsymbol{y}(t)-\boldsymbol{x}(t)\|$ between the spheres as a function of time. The sign of its time derivative indicates whether the spheres are approaching (incident) or escaping each other (recedent):

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t}\|\boldsymbol{y}(t)-x(t)\|=\frac{y-x}{\|y-x\|} \cdot(w-v)=(w-v) \cdot \boldsymbol{n} \tag{24}
\end{equation*}
$$

Clearly, whenever the time derivative above is positive (that is, the spheres are recedent), the dot-product $(\boldsymbol{w}-\boldsymbol{v}) \cdot n$ is positive, and vice versa. Following the authors of $[8,9,11]$, we use the condition in Equation (10) for $F$ and rearrange the collision integral in Equation (23), with help of the Heaviside step-function $\Theta(x)$, as follows:

$$
\begin{align*}
& \sigma^{2} \int F(t, x, x+\sigma n, v, w)(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w}= \\
& =\sigma^{2} \int F\left(t, x, x+\sigma n, \boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \Theta((\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w}+ \\
& \quad+\sigma^{2} \int F(t, x, x+\sigma n, \boldsymbol{v}, \boldsymbol{w})(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \Theta((\boldsymbol{v}-\boldsymbol{w}) \cdot \boldsymbol{n}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w}= \\
& =\sigma^{2} \int\left(F\left(t, x, x+\sigma n, \boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)-F(t, x, \boldsymbol{x}-\sigma n, \boldsymbol{v}, \boldsymbol{w})\right)(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \Theta((\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \tag{25}
\end{align*}
$$

Above, we split the collision integral into two hemispheres-one where $(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}>0$ (that is, the spheres are recedent), and another one where the spheres are incident. In the recedent hemisphere, we replace $F(\boldsymbol{v}, \boldsymbol{w})$ with $F\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$, and $\boldsymbol{v}^{\prime}=\boldsymbol{v}^{\prime}(\boldsymbol{v}, \boldsymbol{w}, \boldsymbol{n})$ and $\boldsymbol{w}^{\prime}=\boldsymbol{w}^{\prime}(\boldsymbol{v}, \boldsymbol{w}, \boldsymbol{n})$ are given via Equation (22). This is a valid rearrangement since the condition in Equation (10) requires $F(\boldsymbol{v}, \boldsymbol{w})$ and $F\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$ to be equal anywhere on the surface of integration. In the incident hemisphere, we change the sign of $n$ to the opposite. Then, we merge the integrals. The resulting BBGKY identity is given via

$$
\begin{equation*}
\frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=\sigma^{2} \int\left(F\left(t, x, x+\sigma n, v^{\prime}, w^{\prime}\right)-F(t, x, x-\sigma n, v, w)\right)(w-v) \cdot n \Theta((w-v) \cdot n) \mathrm{d} n \mathrm{~d} w \tag{26}
\end{equation*}
$$

Clearly, the BBGKY identity above is the same identity as in Equation (23), albeit rewritten in the form of Equation (26). The transition from Equation (23) to Equation (26) is justified for any $F$ which satisfies Equation (10) on the collision surface, with $f$ being the marginal distribution of $F$ given via Equation (11).

### 2.3. The Boltzmann Hierarchy and Boltzmann Equation

The derivation above can be extended to $K$ spheres [ $8,9,11$ ], assuming that $F$ is symmetric under an arbitrary permutation of the spheres. This results in the multiple BBGKY identities for marginal distributions of various orders, chain-linked to each other starting with the highest-order. The lowest-order identity in the BBGKY hierarchy for $K$ spheres is given via

$$
\begin{align*}
\frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \sigma^{2} \int\left(F^{(2)}( \right. & \left.t, x, x+\sigma n, v^{\prime}, w^{\prime}\right)- \\
& \left.-F^{(2)}(t, x, x-\sigma \boldsymbol{n}, \boldsymbol{v}, \boldsymbol{w})\right)(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \Theta((\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \tag{27}
\end{align*}
$$

where $F^{(2)}$ is the two-sphere marginal distribution of the full $K$-sphere density $F$ :

$$
\begin{equation*}
F^{(2)}\left(t, x_{1}, x_{2}, v_{1}, v_{2}\right)=\int F\left(t, x_{1}, \ldots x_{K}, v_{1}, \ldots, v_{K}\right) \mathrm{d} x_{3} \mathrm{~d} x_{K} \ldots \mathrm{~d} v_{3} \mathrm{~d} v_{K} \tag{28}
\end{equation*}
$$

The Boltzmann equation [7-10,12,13] is obtained from the BBGKY identity in Equation (27) using the following two steps. First, one assumes that

$$
\begin{equation*}
\sigma \rightarrow 0, \quad K \rightarrow \infty, \quad K \sigma^{2} \rightarrow \text { constant } \tag{29}
\end{equation*}
$$

which is known as the Boltzmann-Grad limit [10]. As $\sigma \rightarrow 0$, the following assumption is made [9,11]:

$$
\begin{equation*}
F^{(2)}(t, x, x-\sigma n, v, w) \approx F^{(2)}(t, x, x, v, w), \quad F^{(2)}\left(t, x, x+\sigma n, v^{\prime}, w^{\prime}\right) \approx F^{(2)}\left(t, x, x, v^{\prime}, w^{\prime}\right) \tag{30}
\end{equation*}
$$

This assumption transforms the BBGKY hierarchy into what is known as the Boltzmann hierarchy [9,11], whose lowest-order relation is given via

$$
\begin{equation*}
\frac{\partial f}{\partial t}+\boldsymbol{v} \cdot \frac{\partial f}{\partial x}=(K-1) \sigma^{2} \int\left(F^{(2)}\left(t, x, \boldsymbol{x}, \boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)-F^{(2)}(t, x, x, v, w)\right)(\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n} \Theta((\boldsymbol{w}-\boldsymbol{v}) \cdot \boldsymbol{n}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \tag{31}
\end{equation*}
$$

Next, the joint two-sphere marginal density $F^{(2)}$ above is approximated as follows:

$$
\begin{equation*}
F^{(2)}(t, x, x, v, w)=f(t, x, v) f(t, x, w), \quad F^{(2)}\left(t, x, x, v^{\prime}, w^{\prime}\right)=f\left(t, x, v^{\prime}\right) f\left(t, x, w^{\prime}\right) \tag{32}
\end{equation*}
$$

Substituting the approximations in Equation (32) into Equation (31), one arrives at

$$
\begin{equation*}
\frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \sigma^{2} \int\left(f\left(x, v^{\prime}\right) f\left(x, w^{\prime}\right)-f(x, v) f(x, w)\right)(w-v) \cdot n \Theta((w-v) \cdot n) \mathrm{d} n \mathrm{~d} w \tag{33}
\end{equation*}
$$

The relation above is known as the Boltzmann equation [7-10,12,13]. Unlike the BBGKY identity in Equation (27), built upon a (presumably known) solution of the Liouville problem for $K$ spheres, Equation (33) is treated as a closed, self-contained equation for the marginal distribution $f$. The factor $(K-1) \sigma^{2}$ in front of the integral above can be changed to $K \sigma^{2}$, since the Boltzmann equation is the result of the Boltzmann-Grad limit in Equation (29).

## 3. Inconsistencies in the Derivation of the Boltzmann Equation

Here, we point out some contradictions in the conventional derivation of boltzmann equation in Equation (33) from the Liouville problem in Equations (8)-(10), presented in the previous section. We start with a brief summary of the derivation.
(1) The Liouville equation in Equation (9) by itself does not have any collision effects. The effect of collision is imposed separately for every pair of spheres with coordinates $x$ and $y$, on the surface $\|x-y\|=\sigma$.
(2) Due to the effect on the collision surface, the probability density of states $F$ is discontinuous on this surface-it is zero for all $\|x-y\|<\sigma$ according to Equation (8) (as the spheres are impenetrable) and is generally nonzero otherwise.
(3) Due to the discontinuity of $F$ on the collision surface, the collision surface integrals emerge according to the Gauss theorem and the Leibniz rule, when the BBGKY hierarchy is constructed. The resulting identity is given by Equation (23)—observe that it is valid for any $F$, which satisfies Equations (8) and (9), and is discontinuous on the collision surface $\|\boldsymbol{x}-\boldsymbol{y}\|=\sigma$.
(4) In addition to the discontinuity, the velocity deflection condition in Equation (10) is imposed on $F$. This condition allows rewriting the collision integral in Equation (23) in the form of Equation (26). Note that it is the same exact integral as obtained originally in Equation (23) via Gauss theorem and Leibniz rule, only written in a different manner thanks to Equation (10). Similarly, one can obtain the $K$-sphere BBGKY identity in Equation (27) $[9,11]$.

After the BBGKY identity is obtained in Equation (26) (or its $K$-sphere version in Equation (27)), the following additional steps are taken to obtain the Boltzmann equation:
(a) The integrand in the collision integral of Equation (27) is modified via Equation (30), where the distance between the centers of colliding spheres is reduced to zero. The result is referred to as the Boltzmann hierarchy in Equation (31).
(b) The already modified integrand in the collision integral of Equation (31) is modified again via Equation (32), where the two-sphere marginal distribution is replaced with the product of two single-sphere marginal distributions. This results in the Boltzmann equation in Equation (33).

Remarkably, neither Step (a) nor Step (b) above is consistent with the conditions under which Steps (1)-(4) are valid. Below, we elaborate on the inconsistencies between Steps (a) and (b) and the conditions under which Steps (1)-(4) are obtained.

### 3.1. A Contradiction between the Liouville Problem and the Boltzmann Hierarchy

Observe that, if $F$ is a distribution of $K$ rigid spheres, then

$$
\begin{equation*}
F^{(2)}(t, x, x, v, w)=0, \quad \text { for all } x, v, w \text { and for any } \sigma \tag{34}
\end{equation*}
$$

The reason for this is the following. Observe that $F^{(2)}$ with the arguments given above is the two-sphere marginal distribution of the full probability density $F$, where the latter is computed at the point where the coordinates of the first and second spheres are identical. In such a state, the first and second spheres are fully overlapped for any value of the diameter $\sigma$. Due to the impenetrability requirement in Equation (8) (which can be extended onto $K$ spheres in a direct manner), the value of $F$ at this state is guaranteed to be zero irrespectively of the states of all other spheres:

$$
\begin{equation*}
F\left(t, x, x, x_{3}, \ldots, x_{K}, v_{1}, \ldots, v_{K}\right)=0, \quad \text { for all } x, x_{3}, \ldots, x_{K}, v_{1}, \ldots, v_{K} \tag{35}
\end{equation*}
$$

The identity above, in turn, means that the marginal integral in Equation (28) of such an $F$ is also zero, which leads to Equation (34).

Now, observe that the marginal distribution $F^{(2)}$, computed for the states of the form in Equation (34), is used in the collision integral of the lowest-order equation of the Boltzmann hierarchy in Equation (31), which was obtained from the BBGKY equation in Equation (27) via Equation (30). Comparing the right-hand side of Equation (31) with Equation (34), we find that the collision integral in Equation (31) must always be zero, if the impenetrability condition in Equation (8) is indeed satisfied.

Therefore, if the approximation of the BBGKY equation in Equation (27) via the Boltzmann hierarchy in Equation (31) is formally valid, this could only mean that any physical effects from the collision integral of the BBGKY equation in Equation (27) somehow vanish in the Boltzmann-Grad limit in Equation (29). However, even if that is indeed the case, it is not the dynamical regime which is
physically interesting or relevant—in typical applications of gas dynamics, the molecular collisions and the associated effects (such as viscosity and heat conductivity) are usually important.

### 3.2. A Contradiction between the Liouville Problem and the Boltzmann Closure

Let us assume that the inconsistency between the Liouville problem in Equations (8)-(10) and the Boltzmann hierarchy in Equation (31), described above, can somehow be "overlooked". Despite that, yet another contradiction arises when the integrand under the collision integral in Equation (31) is replaced with the product of the single-sphere marginal distributions via Equation (32). Recall that the identities in Equations (23) and (26) are the same exact identity, written in two different forms due to the fact the integrand in the collision integral obeys Equation (10)—in fact, the collision integral in Equation (26) is the same exact integral as in Equation (23), rewritten in a different manner thanks to Equation (10).

However, observe that the factorization in Equation (32) does not generally obey the conditions for which the rearrangement of the collision integrals between Equation (23) and Equation (26) is valid. Indeed, if the integrands of the collision integrals in Equations (23) and (26) are replaced with the factorization in Equation (32), for the same transition to be valid, one must have, according to Equation (10),

$$
\begin{equation*}
f(x, v) f(x, w)=f\left(x, v^{\prime}\right) f\left(x, w^{\prime}\right), \quad \text { for all } x, v, w, \text { and all unit vectors } n \tag{36}
\end{equation*}
$$

where $\boldsymbol{v}^{\prime}=\boldsymbol{v}^{\prime}(\boldsymbol{v}, \boldsymbol{w}, \boldsymbol{n})$ and $\boldsymbol{w}^{\prime}=\boldsymbol{w}^{\prime}(\boldsymbol{v}, \boldsymbol{w}, \boldsymbol{n})$ are given via Equation (22). However, solutions of the Boltzmann equation in Equation (33), which are not simple traveling waves, clearly may not have such a property, since Equation (36) automatically turns the Boltzmann collision integral in the right-hand side of Equation (33) into zero.

Additionally, observe that, since Equation (36) automatically sets the Boltzmann collision integral in Equation (33) to zero, any solution that is not a traveling wave is required to violate Equation (10). In other words, not only the violation of the deflection condition in Equation (10) is "overlooked" in the Boltzmann equation, but it is, in fact, used as the quintessential device to create its nontrivial solutions. This happens despite the fact that Equation (10) is the fundamental property of all solutions of the Liouville problem, whether stationary or not, which, together with Equation (8), leads to the BBGKY identity in Equation (26), from which the Boltzmann collision integral and the Boltzmann equation are subsequently derived.

Therefore, we conclude that what we observe in the conventional derivation of the Boltzmann equation is a logical inconsistency-first, a chain of identities is established under the assumption that certain conditions hold; then, the resultant identity is replaced with an equation whose nontrivial solutions must violate requisite conditions under which the identity was derived to begin with.

### 3.3. Reversibility and Loschmidt's Objection

Hypothetically, one can avoid the above arguments by assuming that the approximate transitions from BBGKY to the Boltzmann hierarchy in Equation (30) and further to the factorization in Equation (32) are only valid for incident velocities, but not recedent. More precisely, one may assume that, for recedent directions given via $n \cdot(\boldsymbol{v}-\boldsymbol{w})>0$,

$$
\begin{equation*}
F^{(2)}(t, x, x-\sigma n, v, w) \neq F^{(2)}(t, x, x, v, w) \neq f(t, x, v) f(t, y, w) \tag{37}
\end{equation*}
$$

so that the only option to approximate the value of $F^{(2)}$ in the recedent hemisphere is to first relate it to the incident hemisphere via $\boldsymbol{v}^{\prime}$ and $\boldsymbol{w}^{\prime}$, and then proceed with Equations (30) and (32). Under such an ansatz, the factorization in Equation (32) does not have to satisfy Equation (10), and the relation in Equation (36) does not have to apply between the incident and recedent directions. However, it is easy to see that such an irreversibility of the collision dynamics is also inconsistent with the formulation of the Liouville problem in Equations (8)-(10).

First, observe that the ansatz in Equation (37) is undoubtedly correct, and not just for the recedent, but also for incident directions, since $F^{(2)}(t, x, x, v, w)=0$ for all $x, v, w$ and $\sigma$, as we already pointed out above in Equation (34). Second, even if one somehow "overlooks" this fact, still such a hypothetical dichotomy between the incident and recedent directions is not only unfounded, but is also clearly in contradiction with the original formulation of the Liouville problem in Equations (8)-(10). The reason is that in the Liouville problem all collisions are exactly reversible, and all states of the system reproduce themselves in reverse order if one uses the terminal coordinates and negatives of terminal velocities in place of an initial condition.

In the case of such a reversal, the incident density states become recedent and vice versa, yet the formulation of the Liouville problem in Equations (8)-(10) remains invariant. Thus, the same exact reasoning as above involving the BBGKY hierarchy, the Boltzmann hierarchy in Equation (30) and factorization in Equation (32) can subsequently be applied to the reversed states as well, in which case the formerly recedent states are expressed via Equations (30) and (32). Conversely, if the ansatz in Equation (37) indeed holds for recedent states, then it also automatically holds for the incident states, in which case the transition from Equation (27) to Equation (33) cannot be carried out via Equations (30) and (32) (which is in fact the case due to Equation (34)).

A similar objection to the form of the Boltzmann collision integral has been posed by Loschmidt [14], who pointed out that, due to the entropy inequality (also known as the $H$-theorem [12]), the Boltzmann equation is time-irreversible, while the underlying dynamics of hard spheres, from which the collision integral is derived, are time-reversible. It is commonly known as "Loschmidt's paradox", even though Loschmidt's observation is entirely valid and does not constitute a paradox by itself. Instead, the "paradox" appears due to the logical inconsistency hidden between the formulation of the Liouville problem for hard spheres in Equations (8)-(10) and the resultant Boltzmann collision integral in Equation (33), as we exposed above.

### 3.4. Our Proposal to Correct the Inconsistency

At the same time, there is no doubt that the Boltzmann equation in Equation (33) is a very accurate model of a dilute gas, which is confirmed by numerous observations and experiments. This means that the collision integral in the right-hand side of Equation (33) is a valid approximation of the statistical effects of molecular interactions in practical scenarios, even if there is a conceptual flaw in its conventional derivation [8,9,11]. In what follows, we propose a different model of hard sphere interactions in which such a problem does not manifest, and which also leads to the Boltzmann equation in the dilute gas approximation-albeit in a different way, via the Enskog equation [16-22].

From what is presented above, it is clear that the formulation of the dynamics of the spheres needs to be such that the collision integrals in the BBGKY equation in Equation (27) possess their form independently of what their integrand is, to avoid breaking any prior constraints with the introduction of a closure approximation. For that, we need to disentangle the instantaneous changes of sphere velocities from the properties of $F$ in the Liouville equation, and arrange for them to be the property of the equation itself. This is possible to do if the underlying dynamical process possesses an infinitesimal generator. Generally, let $\boldsymbol{z}(t)$ be a Markov process in a $d$-dimensional Euclidean space. Assuming that $z(t)=z$ is known at the time $t$, let us consider the conditional expectation $\mathbb{E}[\psi(z(t+\varepsilon))]$ of a suitable function $\psi: \mathbb{R}^{d} \rightarrow \mathbb{R}$, for some $\varepsilon>0$. Assume that the following limit exists:

$$
\begin{equation*}
\lim _{\varepsilon \rightarrow 0} \frac{\mathbb{E}[\psi(\boldsymbol{z}(t+\varepsilon))]-\psi(\boldsymbol{z})}{\varepsilon}=\left.\frac{\partial}{\partial \varepsilon} \mathbb{E}_{\mathcal{\varepsilon}}[\psi](\boldsymbol{z})\right|_{\varepsilon=0}=(\mathcal{L} \psi)(\boldsymbol{z}) \tag{38}
\end{equation*}
$$

where the (generally, integro-differential) linear operator $\mathcal{L}$, which is independent of $\psi$, is called the infinitesimal generator. Clearly, $\mathcal{L}$ describes the underlying dynamics of the process-if $z(t)$ is the phase state of the system of particles/spheres containing their coordinates and velocities, then $\mathcal{L}$ describes the interactions between the spheres. For example, for the point particles interacting via a long-range potential [2], $\mathcal{L}$ includes the gradient of the potential function of the force field.

With $\mathcal{L}$ specified explicitly, it is easy to see that the forward equation for the probability density $F(t, z)$ can be derived in a straightforward fashion [24,25] with the help of the adjoint operator $\mathcal{L}^{\dagger}$, assuming that the latter can also be explicitly obtained. Integrating the above identity over $\mathbb{R}^{d}$ against $F(t, z) \mathrm{d} z$, we have

$$
\begin{equation*}
\int F(t, z)\left(\left.\frac{\partial}{\partial \varepsilon} \mathbb{E}_{\varepsilon}[\psi](z)\right|_{\varepsilon=0}-(\mathcal{L} \psi)(z)\right) \mathrm{d} z=0 \tag{39}
\end{equation*}
$$

First, it can be shown [24] that

$$
\begin{equation*}
\left.\int F(t, z) \frac{\partial}{\partial \varepsilon} \mathbb{E}_{\varepsilon}[\psi](z)\right|_{\varepsilon=0} \mathrm{~d} z=\int \psi(z) \frac{\partial}{\partial t} F(t, z) \mathrm{d} z \tag{40}
\end{equation*}
$$

Next, with the help of the adjoint operator $\mathcal{L}^{\dagger}$, we write

$$
\begin{equation*}
\int F(t, \boldsymbol{z})(\mathcal{L} \psi)(\boldsymbol{z}) \mathrm{d} \boldsymbol{z}=\int \psi(\boldsymbol{z})\left(\mathcal{L}^{\dagger} F\right)(t, \boldsymbol{z}) \mathrm{d} \boldsymbol{z} \tag{41}
\end{equation*}
$$

Combining the last two identities and stripping the $\psi$-integral, we arrive at the general form of the forward Kolmogorov equation [24,26] (also known as the Fokker-Planck equation [27]) for $F$ alone:

$$
\begin{equation*}
\frac{\partial}{\partial t} F(t, z)-\left(\mathcal{L}^{\dagger} F\right)(t, z)=0 \tag{42}
\end{equation*}
$$

Since the infinitesimal generator $\mathcal{L}$ is independent of $F$, its adjoint $\mathcal{L}^{+}$is also independent of $F$. Thus, the form of the collision integral in the corresponding BBGKY hierarchy is defined entirely by $\mathcal{L}^{\dagger}$. This, in turn, will allow to replace $F$ with a suitable closure without violating any prior assumptions on the integrand of the collision integral.

Below, we introduce the infinitesimal generator into the hard sphere collision dynamics by randomizing the times at which the velocity jumps occur. More specifically, the velocity jumps still occur according to Equation (2), except that, rather then lying precisely at the collision surface, the coordinate $(x, y)$ of the jump is selected at random along the trajectory from within a small "tolerance" interval around the collision surface. In such a case, the conditional expectation $\mathbb{E}_{t}[\psi]$ is a differentiable function of $t$, even though the velocities of the spheres still change instantaneously. Since a possibility arises that the spheres become overlapped, we make provisions for their unimpeded subsequent separation. The aforementioned tolerance interval can be made as small as necessary, and the probability of the jump can be made as large as necessary, so as to mimic the deterministic collisions with prescribed accuracy. In addition, Loschmidt's objection [14] is avoided since the underlying dynamics are inherently irreversible.

One can argue that the process we propose is a pure abstraction-clearly, the actual atoms in a gas are not known to interact randomly. However, note that the "hard sphere collision" is also an abstraction-in reality, no observable physical object can change its velocity instantaneously. In particular, the atoms in a real gas do not collide with each other instantaneously, but instead interact via the Lennard-Jones potential [1]. Therefore, neither the deterministic hard spheres, nor our random process are completely accurate models of molecular interaction on a microscopic level. In the present context, the quality of a model is defined by its ability to describe macroscopic effects in a gas while being logically consistent with its own abstract foundation.

## 4. Random Dynamics of Hard Spheres

Here, we present the details of the random dynamical system which models collisions of hard spheres, and at the same time possesses the infinitesimal generator. As before, first we consider the dynamics of two spheres with coordinates $x$ and $y$, and velocities $v$ and $w$, respectively. The dynamics of coordinates are, of course, given via Equation (1). For the dynamics of velocities, we consider two separate configurations:

1. Collision configuration. The collision configuration is given by the following two criteria, which must hold concurrently:
(a) The distance $\|x-y\|$ between the centers of spheres satisfies

$$
\begin{equation*}
1-\alpha<\frac{\|x-y\|}{\sigma}<1+\alpha \tag{43}
\end{equation*}
$$

where $0<\alpha \ll 1$ is a constant parameter. The condition above signifies that $\|\boldsymbol{x}-\boldsymbol{y}\| / \sigma \approx$ 1 , within $\pm \alpha$-tolerance (that is, the spheres are separated approximately by their diameter). We say that two spheres are in the "contact zone" whenever the condition in Equation (43) holds for the coordinates of their centers.
(b) The distance $\|x-y\|$ between the centers of spheres also diminishes in time, that is,

$$
\begin{equation*}
(x-y) \cdot(v-w)<0 \tag{44}
\end{equation*}
$$

This condition signifies that the spheres are approaching each other.
In the collision configuration, the velocities $(v, w)$ may be randomly transformed according to Equation (2), with a specified probability. For now, we write informally that in the collision configuration the velocities evolve according to

$$
\begin{equation*}
\mathrm{d} v(t)=-\mathrm{d} w(t)=\text { random jump process, } \tag{45}
\end{equation*}
$$

which changes the velocities instantaneously according to Equation (2). The jump process must be random for the expectation of a jump to be a continuous function of $t$, despite the fact that the velocity jumps themselves remain instantaneous. The continuity of the expectation is the key property which allows the process to possess the infinitesimal generator. The exact representation of the requisite jump process is provided below.
2. Free-flight configuration. If the two conditions above in Equations (43) and (44) do not hold, then the spheres are in the free-flight configuration. In this case, the velocities are constant in time:

$$
\begin{equation*}
\mathrm{d} \boldsymbol{v}(t)=-\mathrm{d} \boldsymbol{w}(t)=0 \tag{46}
\end{equation*}
$$

At this point, we need to choose an appropriate type of random jump process, which is used as a "trigger" to change the velocities instantaneously. The simplest process which is suitable for the given task is the point process [28]-that is, a scalar, piecewise-constant random process which starts at 0 and occasionally increments itself by 1 , randomly and independently of previous increments. Before describing the random hard sphere process in detail, we formulate a general dynamical system driven by a point process, and compute the explicit form of its infinitesimal generator.

### 4.1. A Dynamical System Driven by an Inhomogeneous Point Process

Here, we follow the theory in Section 7.4 of [28] (the original results are presented in [29]). Let $(\Omega, \mathcal{F}, \mathbb{P})$ be the probability space, equipped with a filtration $\left\{\mathcal{F}_{t}\right\}, t \in \mathbb{R}_{\geq 0}$, such that for $0 \leq t_{1} \leq$ $t_{2}<\infty$, the corresponding sigma-algebras are nested as $\mathcal{F}_{t_{1}} \subseteq \mathcal{F}_{t_{2}} \subseteq \mathcal{F}$. Let $h: \mathbb{R}_{\geq 0} \rightarrow \mathbb{R}_{>0}$ be a bounded, strictly positive random variable adapted to $\left\{\mathcal{F}_{t}\right\}$. Let $m: \mathbb{R}_{\geq 0} \rightarrow \mathbb{Z}_{\geq 0}$ be a $\left\{\mathcal{F}_{t}\right\}$-adapted inhomogeneous point process with the conditional intensity $h(t)$, and the compensator $\tau(t)$, given via

$$
\begin{equation*}
\tau(t)=\int_{0}^{t} h(s) \mathrm{d} s \tag{47}
\end{equation*}
$$

By choosing the random variable $h(t)$ appropriately, we can regulate the "temporal density" of jump points of $m(t)$; indeed, depending on the magnitude of $h(t)$, the jump points of $m(t)$ can either arrive in a statistically rapid succession, or, to the contrary, disperse farther away from each other. For what is to follow, it is a necessary requirement that the intensity $h(t)$ is a random variable-both $m(t)$ and $h(t)$ are, however, adapted to the same filtration $\left\{\mathcal{F}_{t}\right\}$, so that if the sequence of values of $h$ up to $t$ is given, so is the sequence of values of $m$.

We now denote the corresponding jump process of $m(t)$ via $\Delta m(t)$ :

$$
\begin{equation*}
\Delta m(t)=m(t)-m(t-) \tag{48}
\end{equation*}
$$

Above, the notation " $t-$ " denotes the left-limit at $t$. Let $M(t, \cdot)$ be the corresponding random measure of $\Delta m(t)$ :

$$
\begin{equation*}
M(t, A)=\text { number of values of } \Delta m(s) \in A \subset \mathbb{R}_{>0} \text { in } 0<s \leq t \tag{49}
\end{equation*}
$$

Let us now define a stochastic process $\boldsymbol{z}(t)$ on a Euclidean space $\mathbb{R}^{d}$ as follows:

$$
\begin{equation*}
z(t)=\boldsymbol{z}(0)+\int_{0}^{t} f(s) \mathrm{d} s+\int_{0}^{t} \int_{\mathbb{R}_{>0}} \xi \boldsymbol{g}(s) M(\mathrm{~d} s, \mathrm{~d} \xi) \tag{50}
\end{equation*}
$$

where $f, g: \mathbb{R}_{\geq 0} \rightarrow \mathbb{R}^{d}$ are suitable (for the purpose of stochastic integration above) random variables, adapted to $\left\{\mathcal{F}_{t}\right\}$. Clearly, $\boldsymbol{z}(t)$ is also $\left\{\mathcal{F}_{t}\right\}$-adapted by construction.

Our task here is to compute the infinitesimal generator of $\boldsymbol{z}(t)$, that is, for a test function $\psi(\boldsymbol{z})$, we would like to compute

$$
\begin{equation*}
\mathcal{L}[\psi](t)=\lim _{\varepsilon \rightarrow 0} \frac{1}{\varepsilon}\left(\mathbb{E}\left[\psi(\boldsymbol{z}(t+\varepsilon)) \mid \mathcal{F}_{t}\right]-\psi(\boldsymbol{z}(t))\right) \tag{51}
\end{equation*}
$$

To compute the expectation above, we need to adapt the integral form of $z$ in Equation (50) to the Itô formula for Lévy-type stochastic integrals (see Chapter 4 of [25]). However, the problem is that the random measure in the right-hand side of Equation (50) is not that of the standard Poisson point process with constant intensity, but that of the point process with random intensity $h(t)$, defined above.

Our first step here is, therefore, the transformation of the stochastic integral in the right-hand side of Equation (50) to an integral against the random measure of a standard Poisson point process. According to Theorem 7.4.I of [28] (also see [29]), the random point process $n: \mathbb{R}_{\geq 0} \rightarrow \mathbb{Z}_{\geq 0}$, defined via

$$
\begin{equation*}
n(t)=m\left(\tau^{-1}(t)\right) \tag{52}
\end{equation*}
$$

is the standard Poisson point process with intensity 1 , where $\tau(t)$ is the random, albeit $\left\{\mathcal{F}_{t}\right\}$-adapted, compensator process of $m(t)$, defined above in Equation (47). Subsequently, we can write the stochastic integral in Equation (50) from $t$ to $t+\varepsilon$ as

$$
\begin{align*}
\int_{t}^{t+\varepsilon} \int_{\mathbb{R}_{>0}} \xi \boldsymbol{g}(s) M(\mathrm{~d} s, \mathrm{~d} \tilde{\xi}) & =\sum_{\substack{t<\leq \leq t+\varepsilon \\
\Delta m(s)>0}} \Delta m(s) \boldsymbol{g}(s)=\sum_{\substack{t<s \leq t+\varepsilon \\
\Delta n(\tau(s))>0}} \Delta n(\tau(s)) \boldsymbol{g}(s)= \\
& =\sum_{\substack{\tau(t)<s \leq \tau(t+\varepsilon) \\
\Delta n(s)>0}} \Delta n(s) \boldsymbol{g}\left(\tau^{-1}(s)\right)=\int_{\tau(t)}^{\tau(t+\varepsilon)} \int_{\mathbb{R}_{>0}} \xi \boldsymbol{g}\left(\tau^{-1}(s)\right) N(\mathrm{~d} s, \mathrm{~d} \xi), \tag{53}
\end{align*}
$$

with $N(t, \cdot)$ being the random measure of the standard Poisson point process with intensity 1 ; in particular, its intensity measure $\mathbb{E} N[25]$ is given via

$$
\begin{equation*}
\mathbb{E} N(\mathrm{~d} s, \mathrm{~d} \xi)=\delta(\xi-1) \mathrm{d} \xi \mathrm{~d} s \tag{54}
\end{equation*}
$$

Substituting the above integral into Equation (50), we write

$$
\begin{equation*}
z(t+\varepsilon)=\boldsymbol{z}(t)+\int_{t}^{t+\varepsilon} f(s) \mathrm{d} s+\int_{\tau(t)}^{\tau(t+\varepsilon)} \int_{\mathbb{R}_{>0}} \xi \boldsymbol{g}\left(\tau^{-1}(s)\right) N(\mathrm{~d} s, \mathrm{~d} \xi) \tag{55}
\end{equation*}
$$

Next, recalling the Itô formula for Lévy-type stochastic integrals in Chapter 4 of [25], we write

$$
\begin{align*}
& \psi(\boldsymbol{z}(t+\varepsilon))-\psi(\boldsymbol{z}(t))=\int_{t}^{t+\varepsilon} \psi^{\prime}(\boldsymbol{z}(s-)) \boldsymbol{f}(s) \mathrm{d} s+ \\
& \quad+\int_{\tau(t)}^{\tau(t+\varepsilon)} \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi) \tag{56}
\end{align*}
$$

Applying the conditional expectation on both sides, we obtain, for the left-hand side and the first term in the right-hand side,

$$
\begin{gather*}
\mathbb{E}\left[\psi(\boldsymbol{z}(t+\varepsilon))-\psi(\boldsymbol{z}(t)) \mid \mathcal{F}_{t}\right]=\mathbb{E}\left[\psi(\boldsymbol{z}(t+\varepsilon)) \mid \mathcal{F}_{t}\right]-\psi(\boldsymbol{z}(t)),  \tag{57a}\\
\mathbb{E} \int_{t}^{t+\varepsilon} \psi^{\prime}(\boldsymbol{z}(s-)) \boldsymbol{f}(s) \mathrm{d} s=\mathbb{E} \int_{0}^{\varepsilon} \psi^{\prime}(\boldsymbol{z}(t+s-)) \boldsymbol{f}(t+s) \mathrm{d} s=\boldsymbol{\varepsilon} \psi^{\prime}(\boldsymbol{z}(t)) \boldsymbol{f}(t)+o(\varepsilon), \tag{57b}
\end{gather*}
$$

where in the second expression the conditional expectation disappears in the leading order term because the $\boldsymbol{z}(t)$ and $\boldsymbol{f}(t)$ are both $\left\{\mathcal{F}_{t}\right\}$-adapted.

The expectation of the stochastic integral is somewhat more complicated. Observe that

$$
\begin{equation*}
\tau(t+\varepsilon)=\tau(t)+\varepsilon h(t)+h^{*}(t, \varepsilon) \tag{58}
\end{equation*}
$$

where in the right-hand side both leading order terms are $\left\{\mathcal{F}_{t}\right\}$-adapted, and the last term is given via

$$
\begin{equation*}
h^{*}(t, \varepsilon)=\int_{0}^{\varepsilon}(h(t+s)-h(t)) \mathrm{d} s \tag{59}
\end{equation*}
$$

At this point, we further assume that the probability of $h(t+s)$ having a discontinuity for $0 \leq s \leq \varepsilon$ is $O(\varepsilon)$, which implies that

$$
\begin{equation*}
\mathbb{E}\left[h^{*}(t, \varepsilon) \mid \mathcal{F}_{t}\right]=o(\varepsilon) \tag{60}
\end{equation*}
$$

The stochastic integral above can then be split into two parts as

$$
\begin{align*}
\int_{\tau(t)}^{\tau(t+\varepsilon)} & \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi)= \\
& =\int_{\tau(t)+\varepsilon h(t)}^{\tau(t)+\varepsilon h} \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi)+ \\
& +\int_{\tau(t)+\varepsilon h(t)}^{\tau(t)+\varepsilon h(t)+h^{*}(t, \varepsilon)} \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi), \tag{61}
\end{align*}
$$

where the conditional expectation of the last integral is $o(\varepsilon)$.
The expectation of the first integral in the right-hand side above can be written as

$$
\begin{align*}
& \mathbb{E} \int_{\tau(t)}^{\tau(t)+\varepsilon h(t)} \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi)= \\
& \quad=\int_{0}^{\varepsilon h(t)} \int_{\mathbb{R}_{>0}} \mathbb{E}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(t+s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(t+s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(t+s)-\right)\right)\right) \mathbb{E} N(\mathrm{~d} s, \mathrm{~d} \xi), \tag{62}
\end{align*}
$$

where we make use of the fact that the limits of integration are $\left\{\mathcal{F}_{t}\right\}$-adapted, and the integrand is non-anticipative, so that the expectation can be carried into the integral and split into the product of the corresponding expectations of the integrand and the Poisson random measure.

For the expectation of the integrand, we use the same argument as above for the limits of integration-namely, we assume that the probability of a discontinuity appearing in the integrand for $0 \leq s \leq \varepsilon h(t)$ is $O(\varepsilon)$. For the expectation of the Poisson random measure, we recall Equation (54). This further leads to

$$
\begin{align*}
\mathbb{E} \int_{\tau(t)}^{\tau(t)+\varepsilon h(t)} \int_{\mathbb{R}_{>0}}\left(\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)+\xi \boldsymbol{g}\left(\tau^{-1}(s)\right)\right)-\psi\left(\boldsymbol{z}\left(\tau^{-1}(s)-\right)\right)\right) N(\mathrm{~d} s, \mathrm{~d} \xi)= \\
=\int_{0}^{\varepsilon h(t)} \int_{\mathbb{R}_{>0}}(\psi(\boldsymbol{z}(t)+\boldsymbol{\xi} \boldsymbol{g}(t))-\psi(\boldsymbol{z}(t))) \delta(\boldsymbol{\xi}-1) \mathrm{d} \xi \mathrm{~d} s+o(\varepsilon)= \\
=\operatorname{sh}(t)(\psi(\boldsymbol{z}(t)+\boldsymbol{g}(t))-\psi(\boldsymbol{z}(t)))+o(\varepsilon) \tag{63}
\end{align*}
$$

Assembling the terms together, we obtain the infinitesimal generator in the form

$$
\begin{equation*}
\mathcal{L}[\psi]=\lim _{\varepsilon \rightarrow 0} \frac{\mathbb{E}\left[\psi(\boldsymbol{z}(t+\varepsilon)) \mid \mathcal{F}_{t}\right]-\psi(\boldsymbol{z}(t))}{\varepsilon}=\psi^{\prime}(\boldsymbol{z}) \boldsymbol{f}(t)+h(t)(\psi(\boldsymbol{z}+\boldsymbol{g}(t))-\psi(\boldsymbol{z})), \tag{64}
\end{equation*}
$$

where in the last identity we denote $z=z(t)$, for brevity.
It is worth noting that the form of the infinitesimal generator above extends naturally onto the free-flight configuration of the dynamics-it suffices to set the variable intensity $h(t)=0$ above. In this case, $z(t)$ in Equation (50) is driven solely by the integral over the vector field $f(t)$ alone.

### 4.2. Random Dynamics of Two Spheres

To adapt the general stochastic process in Equation (50) to the dynamics of spheres, we need to relate $\boldsymbol{z}(t), \boldsymbol{f}(t), \boldsymbol{g}(t)$ and $h(t)$ to the variables of the dynamics. Obviously, $\boldsymbol{z}(t)$ is the state vector of the system, and thus it incorporates the coordinates and velocities of both spheres:

$$
\begin{equation*}
\boldsymbol{z}(t)=(x(t), \boldsymbol{y}(t), \boldsymbol{v}(t), \boldsymbol{w}(t))^{T} \tag{65}
\end{equation*}
$$

Subsequently, $f(t)$ is related to the deterministic component of the dynamics, which is the evolution of the coordinates $\boldsymbol{x}(t)$ and $\boldsymbol{y}(t)$ for given velocities $\boldsymbol{v}$ and $\boldsymbol{w}$ according to Equation (1):

$$
\begin{equation*}
\boldsymbol{f}(t)=\boldsymbol{f}(\boldsymbol{z}(t-))=(\boldsymbol{v}(t-), \boldsymbol{w}(t-), \mathbf{0}, \mathbf{0})^{T} \tag{66}
\end{equation*}
$$

To specify $\boldsymbol{g}(t)$, we observe that the instantaneous change of velocities in Equation (2) can be written, with help of the jump process $\Delta m(t)$, as

$$
\begin{equation*}
\boldsymbol{v}(t)-\boldsymbol{v}(t-)=-(\boldsymbol{w}(t)-\boldsymbol{w}(t-))=\Delta m(t)(\boldsymbol{w}(t-)-\boldsymbol{v}(t-)) \cdot(\boldsymbol{x}-\boldsymbol{y}) \frac{\boldsymbol{x}-\boldsymbol{y}}{\|\boldsymbol{x}-\boldsymbol{y}\|^{2}} \tag{67}
\end{equation*}
$$

where $\boldsymbol{x}=\boldsymbol{x}(t), \boldsymbol{y}=\boldsymbol{y}(t)$. Therefore, we can define $\boldsymbol{g}(t)$ via

$$
\begin{equation*}
g(t)=g(z(t-))=\frac{(x-y) \cdot(v(t-)-w(t-))}{\|x-y\|^{2}}(0,0, y-x, x-y)^{T} \tag{68}
\end{equation*}
$$

Then, we write the process in Equations (1) and (67) as the following stochastic differential equation [25]:

$$
\begin{equation*}
z(t)=z(0)+\int_{0}^{t} f(z(s-)) \mathrm{d} s+\int_{0}^{t} \int_{\mathbb{R}_{>0}} \xi g(z(s-)) M(\mathrm{~d} s, \mathrm{~d} \xi) \tag{69}
\end{equation*}
$$

with $\boldsymbol{z}, \boldsymbol{f}(\boldsymbol{z})$ and $\boldsymbol{g}(\boldsymbol{z})$ given via Equations (65), (66) and (68), respectively.

It remains to specify the variable intensity $h(t)$, which should activate the point process $m(t)$ when both Equations (43) and (44) hold concurrently, and be zero in the free-flight configuration. Here, we define $h(t)$ as

$$
\begin{equation*}
h(t)=h(z(t-))=\lambda \Theta((x-y) \cdot(w-v)) \delta_{\alpha \sigma}(\|x-y\|-\sigma) \frac{x-y}{\|x-y\|} \cdot(w-v) \tag{70}
\end{equation*}
$$

Above, $\lambda>0,0<\alpha \ll 1$ are constant parameters, and $\delta_{\alpha}(x)$ is the standard mollifier of the delta-function $\delta(x)$, given via

$$
\delta_{\alpha}(x)=\frac{1}{\alpha} \phi\left(\frac{x}{\alpha}\right), \quad \phi(x)=\left\{\begin{array}{ll}
c e^{-\frac{1}{1-x^{2}}}, & |x|<1  \tag{71}\\
0, & |x| \geq 1
\end{array} \quad \int_{-1}^{1} \phi(x) \mathrm{d} x=1\right.
$$

where the constant parameter $c$ ensures the proper normalization. For a function $\psi(x)$, which is continuous at zero, we thus have

$$
\begin{equation*}
\lim _{\alpha \rightarrow 0} \int_{-\infty}^{+\infty} \psi(x) \delta_{\alpha}(x) \mathrm{d} x=\psi(0) \tag{72}
\end{equation*}
$$

that is, $\delta_{\alpha}$ can serve as the delta-function in the limit $\alpha \rightarrow 0$, while remaining smooth for finitely small $\alpha$. We denote the anti-derivative of $\delta_{\alpha}$ as $\Theta_{\alpha}$ :

$$
\begin{equation*}
\Theta_{\alpha}(x)=\int_{-\infty}^{x} \delta_{\alpha}(y) \mathrm{d} y \tag{73}
\end{equation*}
$$

Clearly, as $\alpha \rightarrow 0, \Theta_{\alpha}(x)$ becomes the usual Heaviside step-function.
Observe that, in the collision configuration in Equations (43) and (44), the variable intensity of the point process in Equation (70), as required in $[28,29]$, is indeed $\left\{\mathcal{F}_{t}\right\}$-adapted, strictly positive and bounded, since, first, $z(t)$ is $\left\{\mathcal{F}_{t}\right\}$-adapted by construction, and, second, whenever Equations (43) and (44) hold, we have

$$
\begin{equation*}
0<h(\boldsymbol{z}) \leq \lambda \delta_{\alpha \sigma}(0)\|\boldsymbol{w}-\boldsymbol{v}\| \leq 2 \lambda \delta_{\alpha \sigma}(0) \sqrt{E} \tag{74}
\end{equation*}
$$

where $E$ is the constant energy of the system of two spheres. As the intensity of jumps is globally bounded, and the discontinuities in the solution of Equation (69) may occur solely as a result of these jumps, the assumption that the probability of discontinuity occurring on the time scale of $\varepsilon$ is $O(\varepsilon)$, used in the computation of the expectation of the stochastic integral in Equations (58)-(63), is indeed valid.

The dynamics in Equations (69) and (70) function as follows:

- If $\|x-y\|$ is away from $\sigma$, or if $\|x-y\|$ is growing in time (that is, the spheres are escaping each other), then the triggering point process is not present, and the spheres are moving with constant velocities according to Equation (1). Accordingly, the intensity $h(z(t))$ in Equation (70) is zero in the infinitesimal generator of the process, and thus the generator consists of the free-flight term only.
- Once $\|x-y\|$ is close enough to $\sigma$ and, at the same time $\|x-y\|$ is decreasing in time, the spheres enter the contact zone (both conditions in Equations (43) and (44) are satisfied). In this case, the collision-triggering point process becomes present, with the intensity $h(z(t))$ being strictly greater than zero. Then, there are two possibilities:
- A jump in the point process arrives so that the spheres "collide" according to Equation (2). In this case, the spheres start escaping each other (so that Equation (44) no longer holds) and the triggering point process is no longer present. In the infinitesimal generator, the Heaviside function becomes zero and so does the intensity $h(\boldsymbol{z}(t))$.
- A jump does not arrive, so that eventually $\delta_{\alpha \sigma}(\|x-y\|-\sigma)$ decays back to zero together with the intensity $h(\boldsymbol{z}(t))$ of the point process; in this case, the spheres pass through each other without interaction.

In either scenario, the point process $m(t)$ becomes dormant until the spheres approach each other again.

The existence of strong solutions to Equations (69) and (70) in the collision configuration in Equations (43) and (44) is a subject that merits a separate discussion. For the purpose of this work, we assume that bounded strong solutions are sufficiently generic for typical initial conditions, so that the corresponding statistical formulation (the forward Kolmogorov equation) of the dynamics is reliable enough for the description of large ensembles of solutions.

The definition of the jump intensity above in Equation (70) also indicates that the probability that the jump in the point process does not arrive during the collision window is $e^{-\lambda}$, regardless of the values of $\sigma$ or $\alpha$. Indeed, observe that one can write

$$
\begin{equation*}
h(t)=\lambda \Theta\left(-\frac{\mathrm{d}}{\mathrm{~d} t}\|x(t)-\boldsymbol{y}(t)\|\right) \frac{\mathrm{d}}{\mathrm{~d} t} \Theta_{\alpha \sigma}(\sigma-\|x(t)-\boldsymbol{y}(t)\|) \tag{75}
\end{equation*}
$$

which means that if the contact zone is traversed completely (that is, the jump has not arrived), then the compensator $\tau(t)$ in Equation (47) is incremented by $\lambda$. Clearly, to mimic the collisions of hard deterministic spheres in Section 2, one eventually needs to take $\alpha \rightarrow 0$ and $\lambda \rightarrow \infty$, so that, first, the contact zone Equation (43) becomes infinitely thin, and, second, the jump arrives with probability 1 whenever the spheres are in the collision configuration.

The corresponding infinitesimal generator of Equations (69) and (70) is given via

$$
\begin{equation*}
\mathcal{L}[\psi]=f(z) \cdot \frac{\partial \psi}{\partial z}+h(z)(\psi(z+g(z))-\psi(z)) \tag{76}
\end{equation*}
$$

Changing back to the original variables $x, y, v$ and $w$, we write the infinitesimal generator of Equation (69) in the form

$$
\begin{align*}
& \mathcal{L}[\psi]=v \cdot \frac{\partial \psi}{\partial x}+w \cdot \frac{\partial \psi}{\partial y}+\lambda \delta_{\alpha \sigma}(\|x-y\|-\sigma) \frac{x-y}{\|x-y\|} \cdot(\boldsymbol{w}-\boldsymbol{v}) \\
& \Theta((x-y) \cdot(\boldsymbol{w}-\boldsymbol{v}))\left(\psi\left(x, y, v^{\prime}, w^{\prime}\right)-\psi(x, y, v, w)\right) \tag{77}
\end{align*}
$$

where $v^{\prime}$ and $w^{\prime}$ are the functions of $x, y, v$ and $w$ given in Equation (2). The jump portion of the generator above in Equation (77) is not translationally invariant, and thus the process in Equation (69) is not a Lévy process. However, it is a Lévy-type Feller process [25,30], whose infinitesimal generator can be reformulated in the Courrège form [31] via an appropriate change of variables.

The next step is to obtain the corresponding forward Kolmogorov equation $[24,26,27]$ for the probability density of states of the system, which is easily achieved via the integration by parts. Let $F(t, \boldsymbol{x}, \boldsymbol{y}, \boldsymbol{v}, \boldsymbol{w})$ be the corresponding probability distribution of the random process above. We can then integrate Equation (77) against $F$ and obtain, with help of Equation (40),

$$
\begin{align*}
& \int\left(\psi \frac{\partial F}{\partial t}-F v \cdot \frac{\partial \psi}{\partial x}-F w \cdot \frac{\partial \psi}{\partial y}-F \lambda\left(\psi\left(x, y, v^{\prime}, w^{\prime}\right)-\psi(x, y, v, w)\right)\right. \\
&\left.\frac{x-y}{\|x-y\|} \cdot(w-v) \Theta((x-y) \cdot(\boldsymbol{w}-v)) \delta_{\alpha \sigma}(\|x-y\|-\sigma)\right) \mathrm{d} V_{2} \mathrm{~d} S_{2}=0 \tag{78}
\end{align*}
$$

where $\mathrm{d} V_{2}$ is the volume element of the coordinate space, and $\mathrm{d} S_{2}$ is the area element of the sphere of zero momentum and constant energy (the subscript denotes the number of spheres in the system).

Above, the terms with spatial derivatives in $x$ and $y$ can be integrated by parts, with the condition that the boundary effects are not present. For the part with $\psi\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$ we can write, for fixed $x$ and $y$,

$$
\begin{align*}
& \int \psi\left(x, y, v^{\prime}, w^{\prime}\right) F(x, y, v, w) \frac{x-y}{\|x-y\|} \cdot(w-v) \Theta((x-y) \cdot(w-v)) \mathrm{d} S_{2}= \\
& =\int \psi\left(x, y, v^{\prime}, w^{\prime}\right) F(x, y, v, w) \frac{x-y}{\|x-y\|} \cdot\left(v^{\prime}-w^{\prime}\right) \Theta\left((x-y) \cdot\left(v^{\prime}-w^{\prime}\right)\right) \mathrm{d} S_{2}= \\
& =-\int \psi(x, y, v, w) F\left(x, y, v^{\prime}, w^{\prime}\right) \frac{x-y}{\|x-y\|} \cdot(w-v) \Theta((x-y) \cdot(v-w)) \mathrm{d} S_{2} \tag{79}
\end{align*}
$$

where we used Equations (5)-(7) (note that $v^{\prime}$ and $w^{\prime}$ remain on the same zero momentum-constant energy sphere), and in the last identity renamed $v^{\prime} \rightarrow v, w^{\prime} \rightarrow \boldsymbol{w}$ and vice versa, since the integral occurs over the same velocity sphere. As a result, we can recombine the terms as

$$
\begin{align*}
& \int \psi\left(\frac{\partial F}{\partial t}+v \cdot \frac{\partial F}{\partial x}+w \cdot \frac{\partial F}{\partial y}+\lambda \delta_{\alpha \sigma}(\|x-y\|-\sigma) \frac{x-y}{\|x-y\|} \cdot(w-v)\right. \\
& \left.\quad\left[F(x, y, v, w) \Theta((x-y) \cdot(w-v))+F\left(x, y, v^{\prime}, w^{\prime}\right) \Theta((x-y) \cdot(v-w))\right]\right) \mathrm{d} V_{2} \mathrm{~d} S_{2}=0 \tag{80}
\end{align*}
$$

Assuming that $\psi$ is arbitrary, we can strip the integral over $\psi$ and obtain the equation for $F$ alone:

$$
\begin{align*}
& \frac{\partial F}{\partial t}+v \cdot \frac{\partial F}{\partial x}+w \cdot \frac{\partial F}{\partial y}+\lambda \delta_{\alpha \sigma}(\|x-y\|-\sigma) \frac{x-y}{\|x-y\|} \cdot(w-v) \\
& \quad\left[F(x, y, v, w) \Theta((x-y) \cdot(w-v))+F\left(x, y, v^{\prime}, w^{\prime}\right) \Theta((x-y) \cdot(v-w))\right]=0 . \tag{81}
\end{align*}
$$

Unlike the Liouville problem in Equations (8)-(10), here observe that the effect of collisions is present in the equation itself, and is not contingent upon additional properties imposed on $F$.

### 4.3. Extension to Many Spheres

Here, we extend the previously formulated dynamics onto $K$ spheres, with the corresponding coordinates $\boldsymbol{x}_{i}(t)$ and velocities $v_{i}(t), 1 \leq i \leq K$. Observe that we have $K(K-1) / 2$ possible pairs of spheres. To define their random interactions, we introduce $K(K-1) / 2$ independent instances of the point process, each assigned to the pair of $i$ th and $j$ th spheres.

For $1 \leq i<j \leq K$, let us define

$$
\begin{gather*}
\mathbf{Z}=\left(x_{1}, \ldots, x_{K}, v_{1}, \ldots, \boldsymbol{v}_{K}\right)^{T}, \quad z_{i j}=\left(x_{i}, x_{j}, v_{i}, v_{j}\right)^{T}, \quad \boldsymbol{F}(\mathbf{Z})=\left(v_{1}, \ldots, v_{K}, \mathbf{0}, \ldots, \mathbf{0}\right)^{T},  \tag{82}\\
G_{i j}(\mathbf{Z})=\frac{\left(x_{i}-x_{j}\right) \cdot\left(v_{i}-v_{j}\right)}{\left\|x_{i}-x_{j}\right\|^{2}}\left(\mathbf{0}, \ldots, \mathbf{0}, x_{j}-\boldsymbol{x}_{i}, \mathbf{0}, \ldots, \mathbf{0}, x_{i}-x_{j}, \mathbf{0}, \ldots, \mathbf{0}\right)^{T} \tag{83}
\end{gather*}
$$

where the two nonzero entries in the $\boldsymbol{G}_{i j}$-vector above are in the $(K+i)$ th and $(K+j)$ th slots. Let $m_{i j}(t)$ be the set of $K(K-1) / 2$ independent inhomogeneous point processes with conditional intensities $h_{i j}(t)$, given via

$$
\begin{equation*}
h_{i j}(t)=h\left(z_{i j}(t)\right) \tag{84}
\end{equation*}
$$

where $h(z)$ is defined in Equation (70). Let $M_{i j}(t, \cdot)$ be the set of corresponding random measures for $m_{i j}(t)$. Then, the $K$-sphere dynamics is defined via the following system of stochastic differential equations:

$$
\begin{equation*}
\boldsymbol{Z}(t)=\boldsymbol{Z}(0)+\int_{0}^{t} \boldsymbol{F}(\mathbf{Z}(s-)) \mathrm{d} s+\sum_{i=1}^{K-1} \sum_{j=k+1}^{K} \int_{0}^{t} \int_{\mathbb{R}_{>0}} \xi \boldsymbol{G}_{i j}(\mathbf{Z}(s-)) M_{i j}(\mathrm{~d} s, \mathrm{~d} \xi) \tag{85}
\end{equation*}
$$

The process above in Equation (85) is also a Lévy-type Feller process [25,30], which lives on the sphere of zero momentum and constant energy

$$
\begin{equation*}
\sum_{i=1}^{K} \boldsymbol{v}_{i}=\mathbf{0}, \quad E=\frac{1}{2} \sum_{i=1}^{K}\left\|\boldsymbol{v}_{i}\right\|^{2} \tag{86}
\end{equation*}
$$

As above for two spheres, here we assume that the total momentum of the system is zero without loss of generality. The infinitesimal generator of such process is, apparently, given via

$$
\begin{equation*}
\mathcal{L}[\psi]=\boldsymbol{F}(\mathbf{Z}) \cdot \frac{\partial \psi}{\partial \boldsymbol{Z}}+\sum_{i=1}^{K-1} \sum_{j=i+1}^{K} h\left(\boldsymbol{z}_{i j}\right)\left(\psi\left(\mathbf{Z}+\boldsymbol{G}_{i j}(\mathbf{Z})\right)-\psi(\mathbf{Z})\right) \tag{87}
\end{equation*}
$$

In the $\boldsymbol{x}_{i}$ and $\boldsymbol{v}_{i}$ variables, this translates into

$$
\begin{array}{r}
\mathcal{L}[\psi]=\sum_{i=1}^{K} v_{i} \cdot \frac{\partial \psi}{\partial x_{i}}+\sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \lambda \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{j}-v_{i}\right) \\
\Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right)\left(\psi\left(v_{i}^{\prime}, v_{j}^{\prime}\right)-\psi\right) \tag{88}
\end{array}
$$

where the notation $\psi\left(v_{i}^{\prime}, v_{j}^{\prime}\right)$ specifies that all velocity arguments in $\psi$ are set to the corresponding velocities $v_{k}$, except for $i$ th and $j$ th, which are set to $v_{i}^{\prime}$ and $v_{j}^{\prime}$ via Equation (2).

Observe that the dynamics in Equation (85) is a direct extension of the dynamics of two spheres in Equation (69) onto multiple spheres-the evolution of the coordinates is governed by the same equations, and the velocities of each sphere are coupled to all other spheres via the independent point processes $m_{i j}(t)$. In particular, there is no provision for a collision of more than two spheres at once (which is often discussed in the literature [7,9]); however, given the fact that the collisions in the hard sphere model are instantaneous, we assume that the event of a three-sphere collision is improbable for a "generic" initial condition.

It is interesting to note that the properties of an infinitesimal generator similar to Equation (88) are studied in [21] in the same context (that is, a system of $K$ particles interacting according to Equation (2)). However, the collision part of the infinitesimal generator in [21] is scaled differently, as if the intensity parameter $\lambda$ in Equation (88) is set to $(\alpha \sigma)^{3}$. Thus, as $\alpha \rightarrow 0$, the particles described by the generator in [21] cease colliding upon contact. It is, however, unclear where the form of the infinitesimal generator in [21] comes from; the generator in [21] appears to be postulated, rather than derived from an underlying SDE.

To obtain the corresponding forward equation, we follow the same principle as for the two spheres above. First, we integrate Equation (88) against the probability density $F$ and obtain

$$
\begin{align*}
\int\left(\psi \frac{\partial F}{\partial t}+\psi \sum_{i=1}^{K} v_{i} \cdot \frac{\partial F}{\partial x_{i}}-F \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \lambda\left(\psi\left(v_{i}^{\prime}, v_{j}^{\prime}\right)-\psi\right)\right. \\
\left.\frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{j}-v_{i}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right) \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right)\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}=0 \tag{89}
\end{align*}
$$

where $\mathrm{d} V_{K}$ is the volume element of the coordinate space of the $K$ spheres, $\mathrm{d} S_{K}$ is the area element of the corresponding velocity sphere of zero momentum and constant energy, and the term with the
spatial derivatives is integrated by parts assuming that the boundary terms vanish. Now, for all terms with $v_{i}^{\prime}$ and $v_{j}^{\prime}$, we have, for fixed coordinates,

$$
\begin{align*}
& \int F\left(v_{i}, v_{j}\right) \psi\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \frac{x_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(v_{j}-v_{i}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right) \mathrm{d} S_{K}= \\
& =\int F\left(v_{i}, v_{j}\right) \psi\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}^{\prime}-\boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(x_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}^{\prime}-\boldsymbol{v}_{j}^{\prime}\right)\right) \mathrm{d} S_{K}= \\
& =-\int F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \psi\left(\boldsymbol{v}_{i}, v_{j}\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-v_{j}\right)\right) \mathrm{d} S_{K} \tag{90}
\end{align*}
$$

where we use Equations (5)-(7), and observe that, for fixed coordinates, the variables $v_{i}^{\prime}$ and $v_{j}^{\prime}$ sample the same zero momentum and constant energy sphere as do $v_{i}$ and $v_{j}$. Finally, stripping the integral over $\psi \mathrm{d} V_{K} \mathrm{~d} S_{K}$, we arrive at the forward Kolmogorov equation in the form

$$
\begin{align*}
& \frac{\partial F}{\partial t}+\sum_{i=1}^{K} \boldsymbol{v}_{i} \cdot \frac{\partial F}{\partial x_{i}}+\sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \lambda \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{j}-v_{i}\right) \\
& {\left[F \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right)+F\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{i}-v_{j}\right)\right)\right]=0 } \tag{91}
\end{align*}
$$

Note that Equation (91) admits solutions that are symmetric under the reordering of the spheres. These solutions are "physical", that is, they correspond to real-world scenarios where it is impossible to statistically tell the spheres apart.

## 5. Some Properties of the Random Sphere Dynamics

### 5.1. A Two-Sphere Solution along a Characteristic

Above, we imply that sending $\alpha \rightarrow 0$ and $\lambda \rightarrow \infty$ in Equations (69) and (70) should result in a reasonable approximation of the conventional hard sphere dynamics described in Equations (1) and (2). Here, we examine the behavior of the solutions of the Kolmogorov equation in Equation (81) and compare it to the behavior of the solutions of Equations (8)-(10).

We solve Equation (81) using the method of characteristics; namely, we treat the advection term of Equation (81) as the ordinary, scalar spatial derivative in the direction of $(1, v, w)$, with the directional parameter denoted as $s$. With this, $t, x$ and $y$ are given via the straight line

$$
\begin{equation*}
(t(s), \boldsymbol{x}(s), \boldsymbol{y}(s))=\left(0, \boldsymbol{x}_{0}, \boldsymbol{y}_{0}\right)+s(1, \boldsymbol{v}, \boldsymbol{w}) \tag{92}
\end{equation*}
$$

On this straight line, Equation (81) becomes

$$
\begin{align*}
\frac{\mathrm{d} F}{\mathrm{~d} s}=\lambda \frac{\mathrm{d}}{\mathrm{~d} s} \Theta_{\alpha \sigma}(\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|-\sigma) & {[ }
\end{aligned} \begin{aligned}
& \Theta((\boldsymbol{x}(s)-\boldsymbol{y}(s)) \cdot(\boldsymbol{w}-\boldsymbol{v}))+ \\
& \left.+F\left(s, \boldsymbol{x}(s), \boldsymbol{y}(s), \boldsymbol{v}^{\prime}(s), \boldsymbol{w}^{\prime}(s)\right) \Theta((\boldsymbol{x}(s)-\boldsymbol{y}(s)) \cdot(\boldsymbol{v}-\boldsymbol{w}))\right] \tag{93}
\end{align*}
$$

Here, the two terms in the right-hand side are never nonzero simultaneously, due to the Heaviside step-functions. We assume that the initial condition $\left(x_{0}, y_{0}\right)$ satisfies $\left\|x_{0}-y_{0}\right\|>\sigma(1+\alpha)$, and $\left(x_{0}-\right.$ $\left.y_{0}\right) \cdot(\boldsymbol{v}-\boldsymbol{w})<0$, so that the spheres are not overlapping and on the "collision course". Then, as the characteristic traverses the contact zone for the first time, we have

$$
\begin{equation*}
\frac{\mathrm{d} F}{\mathrm{~d} s}=\lambda F \frac{\mathrm{~d}}{\mathrm{~d} s} \Theta_{\alpha \sigma}(\|x(s)-y(s)\|-\sigma) \tag{94}
\end{equation*}
$$

which results in the solution

$$
\begin{equation*}
F(s)=F_{0} \exp \left(\lambda \Theta_{\alpha \sigma}(\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|-\sigma)-\lambda \Theta_{\alpha \sigma}\left(\left\|x_{0}-y_{0}\right\|-\sigma\right)\right) \tag{95}
\end{equation*}
$$

Observe that $F(s)=F_{0}$ if $\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|>\sigma(1+\alpha)$, and $F(s)=e^{-\lambda} F_{0}$ if $\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|<\sigma(1-\alpha)$. In the original variables, the solution translates into

$$
\begin{align*}
& F(t, \boldsymbol{x}, \boldsymbol{y}, \boldsymbol{v}, \boldsymbol{w})=F_{0}(\boldsymbol{x}-t \boldsymbol{v}, \boldsymbol{y}-t \boldsymbol{w}, \boldsymbol{v}, \boldsymbol{w}) \\
& \quad \exp \left(\lambda \Theta_{\alpha \sigma}(\|\boldsymbol{x}-\boldsymbol{y}\|-\sigma)-\lambda \Theta_{\alpha \sigma}(\|\boldsymbol{x}-\boldsymbol{y}-t(\boldsymbol{v}-\boldsymbol{w})\|-\sigma)\right) . \tag{96}
\end{align*}
$$

Sending $\alpha \rightarrow 0$ has no effect other than making the contact zone infinitely thin, and, subsequently, the transition between $F_{0}$ and $e^{-\lambda} F_{0}$ instantaneous.

As we continue along the characteristic, it eventually traverses the contact zone again, and exits the overlapped state. In this configuration, we have

$$
\begin{equation*}
\frac{\mathrm{d} F}{\mathrm{~d} s}=\lambda F\left(\boldsymbol{x}(s), \boldsymbol{y}(s), \boldsymbol{v}^{\prime}(s), \boldsymbol{w}^{\prime}(s)\right) \frac{\mathrm{d}}{\mathrm{~d} s} \Theta_{\alpha \sigma}(\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|-\sigma) \tag{97}
\end{equation*}
$$

Here, the right-hand side contains the values of $F\left(\boldsymbol{v}^{\prime}, w^{\prime}\right)$, which are carried along the characteristic curves with directions given via

$$
v^{\prime}(s)=v^{\prime}(x(s), y(s), v, w), \quad w^{\prime}(s)=w^{\prime}(x(s), y(s), v, w)
$$

and should be treated as an external forcing. However, we already know from what is found above that we can express, for some $t^{*}$,

$$
\begin{align*}
& F\left(\boldsymbol{x}(s), \boldsymbol{y}(s), \boldsymbol{v}^{\prime}(s), \boldsymbol{w}^{\prime}(s)\right)=F_{0}\left(\boldsymbol{x}(s)-t^{*} \boldsymbol{v}^{\prime}(s), \boldsymbol{y}(s)-t^{*} \boldsymbol{w}^{\prime}(s), \boldsymbol{v}^{\prime}(s), \boldsymbol{w}^{\prime}(s)\right) \\
& \quad \exp \left(\lambda \Theta_{\alpha \sigma}(\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|-\sigma)-\lambda \Theta_{\alpha \sigma}\left(\left\|\boldsymbol{x}(s)-\boldsymbol{y}(s)-t^{*}\left(\boldsymbol{v}^{\prime}(s)-\boldsymbol{w}^{\prime}(s)\right)\right\|-\sigma\right)\right) . \tag{98}
\end{align*}
$$

Without loss of generality, we can take $t^{*}$ large enough so that the second term under the exponent is $\lambda$ (that is, the point back in time $t^{*}$ is sufficiently far away from the collision zone), which gives

$$
\begin{align*}
& \exp \left(\lambda\left(\Theta_{\alpha \sigma}(\|x(s)-y(s)\|-\sigma)-1\right)\right) \lambda \frac{\mathrm{d}}{\mathrm{~d} s} \Theta_{\alpha \sigma}(\|x(s)-\boldsymbol{y}(s)\|-\sigma)= \\
& \quad=\frac{\mathrm{d}}{\mathrm{~d} s} \exp \left(\lambda\left(\Theta_{\alpha \sigma}(\|\boldsymbol{x}(s)-\boldsymbol{y}(s)\|-\sigma)-1\right)\right) \tag{99}
\end{align*}
$$

The equation for $F$ thus becomes

$$
\begin{equation*}
\frac{\mathrm{d} F}{\mathrm{~d} s}=F_{0}\left(\boldsymbol{x}(s)-t^{*} v^{\prime}(s), \boldsymbol{y}(s)-t^{*} w^{\prime}(s), \boldsymbol{v}^{\prime}(s), w^{\prime}(s)\right) \frac{\mathrm{d}}{\mathrm{~d} s} \exp \left(\lambda\left(\Theta_{\alpha \sigma}(\|x(s)-\boldsymbol{y}(s)\|-\sigma)-1\right)\right) \tag{100}
\end{equation*}
$$

Assuming that the initial condition $F_{0}$ is continuous, and sending $\alpha \rightarrow 0$, we find, upon traversing the collision zone outward,

$$
\begin{equation*}
F=F_{\|x-y\|<\sigma}+\left(1-e^{-\lambda}\right) F_{0}^{\prime} \tag{101}
\end{equation*}
$$

where $F_{0}^{\prime}$ denotes the initial value of the solution that is carried from the outside to the point of the collision zone where our characteristic exits the overlapped region. Assembling the pieces together, we finally arrive at the following result in the limit as $\alpha \rightarrow 0$ : after the characteristic traverses the overlapped region completely, the solution is given via

$$
\begin{equation*}
F=e^{-\lambda} F_{0}+\left(1-e^{-\lambda}\right) F_{0}^{\prime} \tag{102}
\end{equation*}
$$

Subsequently, sending $\lambda \rightarrow \infty$ leads to $F_{0}$ being completely replaced by $F_{0}^{\prime}$ once the characteristic fully traverses the overlapped region $\|x-y\|<\sigma$. This is, obviously, the same solution as the one of the Liouville problem for hard spheres Equations (8)-(10). Thus, as $\alpha \rightarrow 0$ and $\lambda \rightarrow \infty$, the solution of Equations (69) and (70) approximates the solution of Equations (1) and (2) in the sense that, for the same initial condition, the corresponding probability densities of Equations (8)-(10) and (81) converge to each other on the same characteristic curve.

### 5.2. A Steady Solution for Many Spheres

To find some steady solutions of Equation (91) for $0<\lambda<\infty$ and $0<\alpha \ll 1$, we again use the method of characteristics-namely, we treat the advection term of Equation (91) as the derivative in the direction of $\left(1, v_{1}, \ldots, v_{K}\right)$, with the directional parameter denoted as $s$. With this, the variables $\left(t, x_{1}, \ldots, x_{K}\right)$ are given via the straight line

$$
\begin{equation*}
\left(t(s), x_{1}(s), \ldots, x_{K}(s)\right)=\left(0, x_{10}, \ldots, x_{K 0}\right)+s\left(1, v_{1}, \ldots, v_{K}\right) \tag{103}
\end{equation*}
$$

Then, along the straight line in Equation (103) we have

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} s} F\left(t(s), x_{1}(s), \ldots, x_{K}(s)\right)=\left(1, v_{1}, \ldots, v_{K}\right) \cdot \frac{\partial F}{\partial\left(t, x_{1}, \ldots, x_{K}\right)} \tag{104}
\end{equation*}
$$

The corresponding differential equation along the straight line is subsequently given via

$$
\begin{align*}
& \frac{\mathrm{d} F}{\mathrm{~d} s}+\sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \lambda \delta_{\alpha \sigma}\left(\left\|x_{i}-\boldsymbol{x}_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right) \\
& {\left[F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right)+F \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)\right)\right]=0 } \tag{105}
\end{align*}
$$

To compute steady solutions, it suffices to look for solutions with the property $F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)=F$, for all index pairs $(i, j)$. In such a situation, the Heaviside step-functions in Equation (91) are multiplied by identical $F^{\prime}$ s, and thus coalesce into 1 . The equation for $F$ on the characteristic in Equation (103) thus becomes

$$
\begin{equation*}
\frac{\mathrm{d} F}{\mathrm{~d} s}+F \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \lambda \delta_{\alpha \sigma}\left(\left\|\boldsymbol{x}_{i}(s)-\boldsymbol{x}_{j}(s)\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}(s)-\boldsymbol{x}_{j}(s)}{\left\|\boldsymbol{x}_{i}(s)-\boldsymbol{x}_{j}(s)\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)=0 \tag{106}
\end{equation*}
$$

Equation (106) can obviously be integrated via separation of variables, but before we proceed with that, let us recall Equation (75) and observe that the coefficient in front of $F$ in Equation (106) is by itself the time derivative of $\lambda \Theta_{\alpha \sigma}$, for each index pair $(i, j)$. Then, the separation of variables yields the equation

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} s} \ln F=-\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \frac{\mathrm{~d}}{\mathrm{~d} s} \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}(s)-x_{j}(s)\right\|\right) \tag{107}
\end{equation*}
$$

with the following solution on the straight line in Equation (103):

$$
\begin{equation*}
F=G\left(x_{10}, \ldots, x_{K 0}, v_{1}, \ldots, v_{K}\right) \exp \left(-\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}(s)-x_{j}(s)\right\|\right)\right) \tag{108}
\end{equation*}
$$

Above, $G$ is an arbitrary function of the starting point $\left(x_{10}, \ldots, x_{K 0}\right)$ and the fixed velocity vector $\left(\boldsymbol{v}_{1}, \ldots, \boldsymbol{v}_{K}\right)$.

To substitute the obtained $F$ into Equation (91), we change the notations back to the original variables:

$$
\begin{equation*}
F\left(t, x_{1}, \ldots, x_{K}, v_{1}, \ldots, v_{K}\right)=G\left(x_{1}-t v_{1}, \ldots, x_{K}-t v_{K}, v_{1}, \ldots, v_{K}\right) \bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right) \tag{109}
\end{equation*}
$$

where the subscript refers to the number of spheres in the system, and we denote

$$
\begin{gather*}
\bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right)=\frac{1}{Z_{K} S_{K}} \exp \left(-\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}-x_{j}\right\|\right)\right)  \tag{110a}\\
Z_{K}=\int \exp \left(-\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}-x_{j}\right\|\right)\right) \mathrm{d} V_{K} \tag{110b}
\end{gather*}
$$

Above, $S_{K}$ is the area of the sphere of zero momentum and constant energy for $K$ spheres. Observe that $\bar{F}_{K}$ is normalized to 1 and is a probability density by itself.

Now, to satisfy the condition $F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)=F\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right)$ for all index pairs $(i, j)$, we need to have

$$
\begin{equation*}
G\left(x_{i}-t v_{i}^{\prime}, x_{j}-t v_{j}^{\prime}, v_{i}^{\prime}, v_{j}^{\prime}\right)=G\left(x_{i}-t v_{i}, x_{j}-t v_{j}, v_{i}, v_{j}\right) \tag{111}
\end{equation*}
$$

for all $t, x_{i}, x_{j}, v_{i}$ and $v_{j}$, with $v_{i}^{\prime}$ and $v_{j}^{\prime}$ given via Equation (2). Recalling that switching $\left(v_{i}, v_{j}\right) \rightarrow$ $\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)$ preserves the momentum and energy in Equation (86), we have to express $G$ as a function of these two quantities, at the same time preserving the form in Equation (111). Apparently, the only form of $G$ which is consistent with both conditions is given via

$$
\begin{equation*}
G=G\left(\frac{1}{K} \sum_{i=1}^{K}\left(\boldsymbol{x}_{i}-t \boldsymbol{v}_{i}\right), \sum_{i=1}^{K} \boldsymbol{v}_{i}, \sum_{i=1}^{K}\left\|\boldsymbol{v}_{i}\right\|^{2}\right) \tag{112}
\end{equation*}
$$

where the division by $K$ is included for convenience. It is not difficult to see that $F$ of the form in Equation (109) with $G$ being of the form in Equation (111) turns the forward equation in Equation (91) into an identity, and the condition $F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)=F\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right)$ is satisfied for all index pairs $(i, j)$.

At this point, we recall that $v_{1}, \ldots, v_{K}$ belong to the set of zero momentum in Equation (86). This forces $G$ to lose the dependence on $t$ and on the second argument, while the third argument becomes the constant energy $E$ :

$$
\begin{equation*}
G=G\left(\frac{1}{K} \sum_{i=1}^{K} x_{i}, E\right) \tag{113}
\end{equation*}
$$

Thus, we find a family of solutions $F$ with $F\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right)=F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)$ for all pairs $(i, j)$, which are the steady states of Equation (91), uniform on the velocity sphere of zero momentum and constant energy. Apparently, there are infinitely many such steady states for a given system of hard spheres. However, note that $G$ above in Equation (113) is a function of the center of mass of the system of spheres, which is invariant under the dynamics of Equation (85).

As becomes important below, observe that there is a family of initial conditions of Equation (91) that has a uniform distribution of the center of mass of the spheres. For such a family of initial conditions, $G=1$, and we arrive at the steady solution, which consists purely of $\bar{F}_{K}$ in Equation (106). Qualitatively, $\bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right)$ behaves as follows:

- $\bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right)$ is constant outside contact zones (that is, the regions in the coordinate space where the mollifier $\left.\delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right)>0\right)$, and its transitions within contact zones are given via the exponents of anti-derivatives of the mollifier in Equation (71).
- Outside contact zones, the following condition holds:

$$
\begin{equation*}
\bar{F}_{\text {overlap }}=e^{-\lambda(n-1)} \bar{F}_{\text {non-overlap }} \tag{114}
\end{equation*}
$$

where $n$ is the number of the simultaneously overlapping spheres at a given point $\left(x_{1}, \ldots, x_{K}\right)$.
Observe that the preceding discussion implicitly assumes that the steady state of the form in Equations (110) and (113) is attracting for a given initial condition. This is certainly not the case for all initial conditions-there exist other steady states for which $F\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right) \neq F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)$; for example, those which are not supported in the contact zones, such that the spheres do not interact at all. However, below, we argue that the state $\bar{F}_{K}$ in Equation (110) is the most likely candidate for a "physical" (that is, statistically most common) steady state for Equation (91).

### 5.3. Entropy Inequality

For the $K$-sphere dynamics in Equation (91), the conventional Boltzmann (also Shannon [32]) entropy is given via

$$
\begin{equation*}
\mathcal{E}(F)=-\int F \ln F \mathrm{~d} V_{K} \mathrm{~d} S_{K} \tag{115}
\end{equation*}
$$

Here, we show that, while the Boltzmann entropy $\mathcal{E}$ can both increase and decrease in time, its appropriate modification, known as the Kullback-Leibler entropy [15,33-36],

$$
\begin{equation*}
P\left(F, G \bar{F}_{K}\right)=\int F \ln \left(\frac{F}{G \bar{F}_{K}}\right) \mathrm{d} V_{K} \mathrm{~d} S_{K} \tag{116}
\end{equation*}
$$

is a non-increasing function of time. In addition, unlike the Boltzmann entropy, the Kullback-Leibler entropy $P$ is invariant under arbitrary changes of variables, and is also nonnegative, that is, for two probability densities $F_{1}$ and $F_{2}$,

$$
\begin{equation*}
P\left(F_{1}, F_{2}\right) \geq 0 \tag{117}
\end{equation*}
$$

where the equality is achieved only when $F_{1}=F_{2}$ (or the difference between the two has zero volume measure on $V_{K} \times S_{K}$ ).

Let $\psi: \mathbb{R} \rightarrow \mathbb{R}$ be a differentiable function. Then, multiplying both sides of Equation (91) by its derivative $\psi^{\prime}$ with $F$ used for the argument, we obtain

$$
\begin{align*}
\left(\frac{\partial}{\partial t}+\sum_{i=1}^{K} v_{i} \cdot \frac{\partial}{\partial x_{i}}\right) \psi(F) & =\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{i}-v_{j}\right) \\
\psi^{\prime}(F) & {\left[F\left(v_{i}^{\prime}, v_{j}^{\prime}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{i}-v_{j}\right)\right)+F \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right)\right] . } \tag{118}
\end{align*}
$$

Integrating over $\mathrm{d} V_{K} \mathrm{~d} S_{K}$, we obtain

$$
\begin{align*}
& \frac{\partial}{\partial t} \int \psi(F) \mathrm{d} V_{K} \mathrm{~d} S_{K}=\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int \delta_{\alpha \sigma}\left(\left\|x_{i}-\boldsymbol{x}_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right) \\
& \psi^{\prime}(F)\left[F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right)+F \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)\right)\right] \mathrm{d} V_{K} \mathrm{~d} S_{K} . \tag{119}
\end{align*}
$$

where we denote $F^{\prime}=F\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$. Via Equations (5)-(7), we write, for any $i$ and $j$, the term with $F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)$ in the right-hand side as

$$
\begin{align*}
& \int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}-v_{j}\right) \psi^{\prime}(F) F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(x_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}= \\
& \quad=\int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}^{\prime}-\boldsymbol{v}_{i}^{\prime}\right) \psi^{\prime}(F) F\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}^{\prime}-\boldsymbol{v}_{i}^{\prime}\right)\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}= \\
& \quad=\int \delta_{\alpha \sigma}\left(\left\|x_{i}-\boldsymbol{x}_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right) \psi^{\prime}\left(F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)\right) F \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-v_{i}\right)\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}, \tag{120}
\end{align*}
$$

which leads to

$$
\begin{align*}
\frac{\partial}{\partial t} \int \psi(F) \mathrm{d} V_{K} \mathrm{~d} S_{K}=\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int & \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-v_{i}\right) \\
& \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)\right)\left[\psi^{\prime}\left(F\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right)\right)-\psi^{\prime}(F)\right] F \mathrm{~d} V_{K} \mathrm{~d} S_{K} \tag{121}
\end{align*}
$$

For the entropy in Equation (115), we substitute

$$
\begin{equation*}
\psi(F)=-F \ln F, \quad \psi^{\prime}(F)=-(1+\ln F) \tag{122}
\end{equation*}
$$

and arrive at the following equation for $\mathcal{E}$ :

$$
\begin{array}{r}
\frac{\partial \mathcal{E}}{\partial t}=\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-v_{i}\right) \\
\Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right) F \ln \left(\frac{F}{\left.F\left(v_{i}^{\prime}, v_{j}^{\prime}\right)\right)}\right) \mathrm{d} V_{K} \mathrm{~d} S_{K} . \tag{123}
\end{array}
$$

Next, recalling the inequality

$$
x \ln x \geq x-1
$$

observe that, for two probability densities $F_{1}$ and $F_{2}$, we have

$$
\begin{equation*}
F_{1} \ln \left(\frac{F_{1}}{F_{2}}\right)=F_{2} \frac{F_{1}}{F_{2}} \ln \left(\frac{F_{1}}{F_{2}}\right) \geq F_{2}\left(\frac{F_{1}}{F_{2}}-1\right)=F_{1}-F_{2} \tag{124}
\end{equation*}
$$

which, upon substitution into Equation (123), yields the following inequality:

$$
\begin{array}{r}
\frac{\partial \mathcal{E}}{\partial t} \geq \lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right) \\
\Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)\right)\left(F-F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)\right) \mathrm{d} V_{K} \mathrm{~d} S_{K} . \tag{125}
\end{array}
$$

For the part with $F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right)$, we observe that, for any $i$ and $j$, we have

$$
\begin{align*}
& -\int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{j}-v_{i}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right) F\left(v_{i}^{\prime}, v_{j}^{\prime}\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}= \\
& =-\int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}^{\prime}-\boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(x_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}^{\prime}-v_{j}^{\prime}\right)\right) F\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}= \\
& \quad=\int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-v_{i}\right) \Theta\left(\left(x_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-v_{j}\right)\right) F \mathrm{~d} V_{K} \mathrm{~d} S_{K} \tag{126}
\end{align*}
$$

which allows to coalesce the two Heaviside functions into 1 and results in

$$
\begin{equation*}
\frac{\partial \mathcal{E}}{\partial t} \geq \lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-v_{i}\right) F \mathrm{~d} V_{K} \mathrm{~d} S_{K} \tag{127}
\end{equation*}
$$

The right-hand side of Equation (127) above can, in general, be negative, and, therefore, it is possible for the Boltzmann entropy of the system to decrease in general. As an example, consider the uniform initial condition in both coordinates and velocities, which, under the normalization constraint, maximizes the entropy over all possible states. However, this state is not a steady state of

Equation (91); indeed, the states corresponding to the overlapped spheres are "drained", the solution becomes non-uniform in coordinates, and the entropy decreases as a result.

To examine the Kullback-Leibler entropy in Equation (116), let us look at the expression

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(F \ln \left(G \bar{F}_{K}\right)\right)=-\ln \left(G \bar{F}_{K}\right) \sum_{i=1}^{K} v_{i} \cdot \frac{\partial F}{\partial x_{i}}+\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \ln \left(G \bar{F}_{K}\right) \\
& \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\left[F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right)+F \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right)\right)\right] \tag{128}
\end{align*}
$$

where $G \bar{F}_{K}$ is the steady state from Equations (110) and (113). We rearrange the terms above as

$$
\begin{align*}
\frac{\partial}{\partial t}\left(F \ln \left(G \bar{F}_{K}\right)\right)=-\sum_{i=1}^{K} \boldsymbol{v}_{i} \cdot & \frac{\partial}{\partial x_{i}}\left(F \ln \left(G \bar{F}_{K}\right)\right)+\frac{F}{G \bar{F}_{K}} \sum_{i=1}^{K} v_{i} \cdot \frac{\partial}{\partial x_{i}}\left(G \bar{F}_{K}\right)+ \\
& +\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{i}-v_{j}\right) \\
& \ln \left(G \bar{F}_{K}\right)\left[F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right)+F \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right)\right] . \tag{129}
\end{align*}
$$

Using Equation (91) and the fact that $G \bar{F}_{K}$ is a steady state, we write

$$
\begin{equation*}
\sum_{i=1}^{K} v_{i} \cdot \frac{\partial}{\partial x_{i}}\left(G \bar{F}_{K}\right)=\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(v_{i}-v_{j}\right) G \bar{F}_{K} \tag{130}
\end{equation*}
$$

and subsequently obtain

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(F \ln \left(G \bar{F}_{K}\right)\right)=-\sum_{i=1}^{K} v_{i} \cdot \frac{\partial}{\partial x_{i}}\left(F \ln \left(G \bar{F}_{K}\right)\right)+\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \\
& \quad \cdot\left(v_{i}-v_{j}\right)\left(F+\ln \left(G \bar{F}_{K}\right)\left[F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{i}-v_{j}\right)\right)+F \Theta\left(\left(x_{i}-x_{j}\right) \cdot\left(v_{j}-v_{i}\right)\right)\right]\right) . \tag{131}
\end{align*}
$$

Upon the integration over $\mathrm{d} V_{K} \mathrm{~d} S_{K}$, the terms with $\ln \left(G \bar{F}_{K}\right)$ in the right-hand side above disappear, and we arrive at

$$
\begin{equation*}
\frac{\partial}{\partial t} \int F \ln \left(G \bar{F}_{K}\right) \mathrm{d} V_{K} \mathrm{~d} S_{K}=\lambda \sum_{i=1}^{K-1} \sum_{j=i+1}^{K} \int \delta_{\alpha \sigma}\left(\left\|x_{i}-x_{j}\right\|-\sigma\right) \frac{x_{i}-x_{j}}{\left\|x_{i}-x_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}-v_{j}\right) F \mathrm{~d} V_{K} \mathrm{~d} S_{K} \tag{132}
\end{equation*}
$$

Adding Equation (132) to Equation (127) and changing the sign on both sides, we arrive at

$$
\begin{equation*}
\frac{\partial}{\partial t} P\left(F, G \bar{F}_{K}\right) \leq 0 \tag{133}
\end{equation*}
$$

that is, the Kullback-Leibler entropy between a solution of Equation (91) and the steady state in Equations (110) and (113) is a nonnegative, non-increasing function of time. The inequality in Equation (133) is the analog of Boltzmann's $H$-theorem [9,10,13] for the Kolmogorov equation in Equation (91).

### 5.4. The Steady Solution for a System with Independently Distributed Initial States

Recall that the factor $G$ in Equation (113) is determined by the distribution of the center of mass of the spheres, which, in turn, is defined entirely by the corresponding initial condition for the forward equation in Equation (91). Thus far, we did not elaborate on the choice of the initial conditions, and thus $G$ was presumed to be largely arbitrary. However, from the perspective of physics, some of the
states are more realistic, while others are notably less so. As an example of a physically unrealistic state, one can arrange the spheres initially so they fly in straight lines without interacting. However, in real-world systems, such as gases and liquids, these states are not encountered in practice, and instead a strongly chaotic motion of frequently colliding molecules is observed.

Here, we propose the "most common" steady state for a system with large number of spheres, based on a "physicist's reasoning". Namely, in what follows, we assume that the spheres are initially distributed independently of each other. We compute the distribution of their center of mass, and, observing that it is invariant under the dynamics, conclude that the relevant steady state must have the same distribution of the center of mass. Note that the independence of distributions for each sphere suggests that the initial measures of overlapped states may be nonzero, however, in our random gas model, it is permissible for the spheres to overlap.

Let the distribution $F_{0}$ of the $K$ spheres in the system at the initial moment of time be given via the product of independent identical distributions $f_{0}$ for each sphere:

$$
\begin{equation*}
F_{0}\left(\boldsymbol{x}_{1}, \ldots, \boldsymbol{x}_{K}, \boldsymbol{v}_{1}, \ldots, \boldsymbol{v}_{K}\right)=\prod_{i=1}^{K} f_{0}\left(\boldsymbol{x}_{i}, \boldsymbol{v}_{i}\right) \tag{134}
\end{equation*}
$$

Let the position of the center of mass of the system be given via $y$ :

$$
\begin{equation*}
y=\frac{1}{K} \sum_{i=1}^{K} x_{i} \tag{135}
\end{equation*}
$$

Let us express the first coordinate, $x_{1}$, through the position of the center of mass $y$, and the rest of the coordinates:

$$
\begin{equation*}
x_{1}=K y-\sum_{i=2}^{K} x_{i} \tag{136}
\end{equation*}
$$

Then, the distribution of $y$ is given via

$$
\begin{align*}
& g(\boldsymbol{y})=K^{3} \int f_{0}\left(K \boldsymbol{y}-\sum_{i=2}^{K} \boldsymbol{x}_{i}, \boldsymbol{v}_{1}\right) \prod_{i=2}^{K} f_{0}\left(\boldsymbol{x}_{i}, \boldsymbol{v}_{i}\right) \mathrm{d} \boldsymbol{x}_{2} \ldots \mathrm{~d} \boldsymbol{x}_{K} \mathrm{~d} \boldsymbol{v}_{1} \ldots \mathrm{~d} \boldsymbol{v}_{K}= \\
&=K^{3} \int f_{0}^{x}\left(K \boldsymbol{y}-\sum_{i=2}^{K} x_{i}\right) \prod_{i=2}^{K} f_{0}^{x}\left(\boldsymbol{x}_{i}\right) \mathrm{d} \boldsymbol{x}_{2} \ldots \mathrm{~d} \boldsymbol{x}_{K} \tag{137}
\end{align*}
$$

where $f_{0}^{x}$ denotes the $x$-marginal of $f_{0}$ :

$$
\begin{equation*}
f_{0}^{x}(x)=\int f_{0}(x, v) \mathrm{d} v \tag{138}
\end{equation*}
$$

In what follows, we assume that $f_{0}^{x}$ is nondegenerate, that is, its support is of the same dimension as that of the whole $x$-space. Let $\hat{f}$ denote the characteristic function of $f_{0}^{x}$,

$$
\begin{equation*}
\hat{f}(\boldsymbol{k})=\int e^{i \boldsymbol{k} \cdot \boldsymbol{x}} f_{0}^{x}(\boldsymbol{x}) \mathrm{d} \boldsymbol{x}, \quad f_{0}^{x}(\boldsymbol{x})=\frac{1}{(2 \pi)^{3}} \int e^{-i \boldsymbol{k} \cdot \boldsymbol{x}} \hat{f}(\boldsymbol{k}) \mathrm{d} \boldsymbol{k} \tag{139}
\end{equation*}
$$

Clearly, $\hat{f}(\mathbf{0})=1$, since $f_{0}^{x}$ is a probability density. Additionally, due to the fact that $f_{0}^{x}$ is nondegenerate, $|\hat{f}(\boldsymbol{k})|<1$ for $\boldsymbol{k} \neq \mathbf{0}$ [37].

Next, let us express $g(\boldsymbol{y})$ via the characteristic function of $f_{0}^{x}$ from Equation (139):

$$
\begin{equation*}
g(\boldsymbol{y})=\frac{K^{3}}{(2 \pi)^{3 K}} \int \mathrm{~d} \boldsymbol{k}_{1} \ldots \mathrm{~d} \boldsymbol{k}_{K} \hat{f}\left(\boldsymbol{k}_{1}\right) \ldots \hat{f}\left(\boldsymbol{k}_{K}\right) e^{-i K \boldsymbol{k}_{1} \cdot \boldsymbol{y}} \int e^{i\left(\boldsymbol{k}_{1}-\boldsymbol{k}_{2}\right) \cdot \boldsymbol{x}_{2}} \ldots e^{i\left(\boldsymbol{k}_{1}-\boldsymbol{k}_{K}\right) \cdot \boldsymbol{x}_{K}} \mathrm{~d} \boldsymbol{x}_{2} \ldots \mathrm{~d} \boldsymbol{x}_{K} . \tag{140}
\end{equation*}
$$

Recalling that, in the generalized sense,

$$
\begin{equation*}
\int e^{i k \cdot x} \mathrm{~d} x=(2 \pi)^{3} \delta(\boldsymbol{k}) \tag{141}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
g(\boldsymbol{y})=\frac{K^{3}}{(2 \pi)^{3}} \int\left(e^{-i \boldsymbol{k} \cdot \boldsymbol{y}} \hat{f}(\boldsymbol{k})\right)^{K} \mathrm{~d} \boldsymbol{k} \tag{142}
\end{equation*}
$$

Observing that $\left|e^{-i \boldsymbol{k} \cdot \boldsymbol{y}} \hat{f}(\boldsymbol{k})\right|<1$ for $\boldsymbol{k} \neq \mathbf{0}$, we find that, in the limit $K \rightarrow \infty$, the integrand approaches zero for all $\boldsymbol{k} \neq \mathbf{0}$. Thus, as $K \rightarrow \infty, g(\boldsymbol{y})$ loses its dependence on $\boldsymbol{y}$ and becomes the uniform distribution. We thus conclude that the uniform distribution of the center of mass of the system of spheres is the most physical one. The corresponding most physical steady state is thus given via $G=1$ in Equation (113), as, under $\bar{F}_{K}$, any given state of the spheres is exactly as likely as its arbitrary parallel coordinate translation, and thus the center of mass of $\bar{F}_{K}$ alone is distributed uniformly.

Observe that the preceding derivation can be modified so that the spheres are distributed independently, but not identically (that is, any $i$ th sphere could be distributed independently, but with its own distribution $f_{0 i}$ ). If so, for the same reasoning to hold, we need all corresponding characteristic functions $\hat{f}_{i}$ of each $\boldsymbol{x}$-marginal to be bounded above by the same $h(\boldsymbol{k})$,

$$
\begin{equation*}
\left|\hat{f}_{i}(\boldsymbol{k})\right| \leq h(\boldsymbol{k}), \quad h(\mathbf{0})=1, \quad h(\boldsymbol{k} \neq \mathbf{0})<1 \tag{143}
\end{equation*}
$$

that is, the spheres should initially be distributed independently and "similarly", which is a reasonable requirement from the physical standpoint.

### 5.5. The Structure of Marginal Distributions of the Physical Steady State

As it becomes important below, we need to discuss the structure of the marginal distributions of the physical steady state $\bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right)$ in Equation (110), that is, the functions of the form

$$
\begin{equation*}
\bar{F}_{K}^{(n)}\left(\boldsymbol{x}_{1}, \ldots, \boldsymbol{x}_{n}, \boldsymbol{v}_{1}, \ldots, \boldsymbol{v}_{n}\right)=S_{n+1 \ldots K}\left(\boldsymbol{v}_{1}, \ldots, \boldsymbol{v}_{n}\right) \int \bar{F}_{K}\left(\boldsymbol{x}_{1}, \ldots, \boldsymbol{x}_{K}\right) \mathrm{d} V_{n+1 \ldots K} \tag{144}
\end{equation*}
$$

where $S_{n+1 \ldots K}\left(\boldsymbol{v}_{1}, \ldots, v_{n}\right)$ is the surface area of the subset of the constant energy sphere which corresponds to fixed velocities $v_{1}, \ldots, v_{n}$, and $\mathrm{d} V_{n+1 \ldots K}$ is the volume element of the subset of the volume which corresponds to fixed coordinates $x_{1}, \ldots, x_{n}$.

The integral over $\mathrm{d} V_{n+1 \ldots K}$ in the right-hand side of Equation (144) can be expressed in the form

$$
\begin{equation*}
\int \bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right) \mathrm{d} V_{n+1 \ldots K}=\bar{F}_{n}\left(x_{1}, \ldots, x_{n}\right) \frac{S_{n}}{S_{K}} R_{K}^{(n)}\left(x_{1}, \ldots, x_{n}\right) . \tag{145}
\end{equation*}
$$

Above, we introduce the spatial correlation function $R_{K}^{(n)}\left(x_{1}, \ldots, x_{n}\right)$ of the form

$$
\begin{align*}
& R_{K}^{(n)}\left(x_{1}, \ldots, x_{n}\right)=\frac{Z_{n}}{Z_{K}} \int\left(\prod_{i=n+1}^{K} e^{-\lambda\left(\Theta_{\alpha \sigma}\left(\sigma-\left\|x_{1}-x_{i}\right\|\right)+\ldots+\Theta_{\alpha \sigma}\left(\left\|x_{n}-x_{i}\right\|\right)\right)}\right. \\
&\left.\prod_{j=i+1}^{K} e^{-\lambda \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}-x_{j}\right\|\right)}\right) \mathrm{d} V_{n+1 \ldots K} \tag{146}
\end{align*}
$$

that is, the $R_{K}^{(n)}$ involves integration over all factors $e^{-\lambda \Theta_{\alpha \sigma}\left(\sigma-\left\|x_{i}-x_{j}\right\|\right)}$ which are part of $\bar{F}_{K}$, but not $\bar{F}_{n}$. Observe that the following normalization conditions hold concurrently:

$$
\begin{equation*}
S_{n} \int \bar{F}_{n}\left(x_{1}, \ldots, x_{n}\right) \mathrm{d} V_{n}=1 \tag{147a}
\end{equation*}
$$

$$
\begin{equation*}
S_{n} \int \bar{F}_{n}\left(x_{1}, \ldots, x_{n}\right) R_{K}^{(n)}\left(x_{1}, \ldots, x_{n}\right) \mathrm{d} V_{n}=S_{K} \int \bar{F}_{K}\left(x_{1}, \ldots, x_{K}\right) \mathrm{d} V_{K}=1 \tag{147b}
\end{equation*}
$$

so if $R_{K}^{(n)}$ does not vary "too much" throughout the domain, its magnitude should generally be of order 1. In situations where the overlapped states take a very small part of the total volume (which implies that the volume of the domain greatly exceeds the total volume of the spheres in it, the so-called "dilute gas"), the integrand of Equation (146) equals 1 almost everywhere, and, therefore, $R_{K}^{(n)} \approx 1$.

For reasons that become clear below, of particular importance are the special cases with $n=1$ and $n=2$. For $n=1$, we claim

$$
\begin{equation*}
\bar{F}_{K}^{(1)}(\boldsymbol{v})=\frac{1}{V} \frac{S_{2 \ldots K}(\boldsymbol{v})}{S_{K}}, \tag{148}
\end{equation*}
$$

where $V$ is the volume of the coordinate space of a single sphere, and the independence on $x$ follows from the fact that $\bar{F}_{K}$ depends only on the distances between the spheres, and not on their absolute positions.

For the special case $n=2$, we claim

$$
\begin{equation*}
R_{K}^{(2)}(x, y)=R_{K}^{(2)}(\|x-y\|), \tag{149}
\end{equation*}
$$

which follows from the symmetry and isotropy considerations. Indeed, observe that, since $\bar{F}_{K}$ depends only on distances between different spheres, $R_{K}^{(2)}$ can depend neither on the absolute positions of its two spheres, nor on the direction of their mutual orientation, which leaves only the distance as a possible dependence. Taking into account Equation (110), for the two-sphere marginal $\bar{F}_{K}^{(2)}$ we, therefore, arrive at

$$
\begin{equation*}
\bar{F}_{K}^{(2)}(\boldsymbol{x}, \boldsymbol{y}, \boldsymbol{v}, \boldsymbol{w})=\frac{1}{Z_{2}} e^{-\lambda \Theta_{\alpha \sigma}(\sigma-\|x-y\|)} R_{K}^{(2)}(\|\boldsymbol{x}-\boldsymbol{y}\|) \frac{S_{3 \ldots K}(\boldsymbol{v}, \boldsymbol{w})}{S_{K}} . \tag{150}
\end{equation*}
$$

Combining Equations (148) and (150), we can express

$$
\begin{equation*}
\bar{F}_{K}^{(2)}(\boldsymbol{x}, \boldsymbol{y}, \boldsymbol{v}, \boldsymbol{w})=\frac{V^{2}}{Z_{2}} \frac{S_{K} S_{3 \ldots K}(\boldsymbol{v}, \boldsymbol{w})}{S_{2 \ldots K}(\boldsymbol{v}) S_{2 \ldots K}(\boldsymbol{w})} e^{-\lambda \Theta_{\alpha \sigma}(\sigma-\|\boldsymbol{x}-\boldsymbol{y}\|)} R_{K}^{(2)}(\|\boldsymbol{x}-\boldsymbol{y}\|) \bar{F}_{K}^{(1)}(\boldsymbol{v}) \bar{F}_{K}^{(1)}(\boldsymbol{w}) \tag{151}
\end{equation*}
$$

### 5.6. The Marginal Distributions in the Limit of Infinitely Many Spheres

In what follows, we are interested in situations where the number of spheres is large, and thus we need to to examine the limit of Equation (151) as $K \rightarrow \infty$. In such a limit, first we observe that

$$
\begin{equation*}
\lim _{K \rightarrow \infty} \frac{V^{2}}{Z_{2}}=1 \tag{152}
\end{equation*}
$$

This happens due to the fact that the diameter $\sigma$ must decay as $K^{-1 / 3}$, in order for the total volume of the spheres to be bounded (to fit into the domain without overlapping). Therefore, the volume of the set of points where the integrand of $Z_{2}$ in Equation (110) equals $e^{-\lambda}$ is proportional to $K^{-1}$. This means that the ratio above behaves as $V^{2} / Z_{2}=1+O\left(K^{-1}\right)$, and approaches 1 as $K \rightarrow \infty$.

Second, it can be shown via geometric arguments (see $[36,38,39]$ and references therein) that, as $K \rightarrow \infty$,

$$
\begin{equation*}
\lim _{K \rightarrow \infty} \frac{S_{2 \ldots K}(\boldsymbol{v})}{S_{K}}=\frac{e^{-\|\boldsymbol{v}\|^{2} /(2 \theta)}}{(2 \pi \theta)^{3 / 2}}, \quad \lim _{K \rightarrow \infty} \frac{S_{3 \ldots K}(\boldsymbol{v}, \boldsymbol{w})}{S_{K}}=\frac{e^{-\left(\|\boldsymbol{v}\|^{2}+\|\boldsymbol{w}\|^{2}\right) /(2 \theta)}}{(2 \pi \theta)^{3}} \tag{153}
\end{equation*}
$$

where $\theta$ is the temperature of the system of the spheres, given via

$$
\begin{equation*}
E=\frac{3}{2} K \theta \tag{154}
\end{equation*}
$$

The relations in Equation (153) lead to

$$
\begin{equation*}
\lim _{K \rightarrow \infty} \frac{S_{K} S_{3 \ldots K}(\boldsymbol{v}, \boldsymbol{w})}{S_{2 \ldots K}(\boldsymbol{v}) S_{2 \ldots K}(\boldsymbol{w})}=1 \tag{155}
\end{equation*}
$$

Combining Equations (151), (152) and (155), we arrive, as $K \rightarrow \infty$, to

$$
\begin{equation*}
\bar{F}^{(2)}(\boldsymbol{x}, \boldsymbol{y}, \boldsymbol{v}, \boldsymbol{w})=e^{-\lambda \Theta_{\alpha \sigma}(\sigma-\|\boldsymbol{x}-\boldsymbol{y}\|)} R^{(2)}(\|\boldsymbol{x}-\boldsymbol{y}\|) \bar{F}^{(1)}(\boldsymbol{v}) \bar{F}^{(1)}(\boldsymbol{w}), \tag{156}
\end{equation*}
$$

where we dropped the subscript $K$ from $\bar{F}^{(1)}, \bar{F}^{(2)}$ and $R^{(2)}$ to signify that these quantities are related to infinitely many spheres.

## 6. The forward Equation for the Marginal Distribution of a Single Sphere

Let us consider an unsteady solution $F$ of the forward equation in Equation (91), which is invariant under the permutations of the spheres, and let us denote the two-sphere and one-sphere marginal distributions of $F$ as

$$
\begin{equation*}
F^{(2)}(t, x, y, v, w)=\int F \mathrm{~d} V_{3 \ldots K} \mathrm{~d} S_{3 \ldots K}, \quad f(t, x, v)=\int F^{(2)} \mathrm{d} y \mathrm{~d} w \tag{157}
\end{equation*}
$$

where the integral in $w$ occurs over the remaining area of $S_{K}$ for a fixed $v$. Then, integrating Equation (91) over all coordinate-velocity pairs $\left(\boldsymbol{x}_{i}, \boldsymbol{v}_{i}\right)$ but one, we arrive at

$$
\begin{align*}
& \frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \lambda \int \delta_{\alpha \sigma}(\|x-y\|-\sigma) \frac{x-y}{\|x-y\|} \cdot(v-w) \\
& \quad\left[F^{(2)}\left(x, y, v^{\prime}, w^{\prime}\right) \Theta((x-y) \cdot(v-w))+F^{(2)}(x, y, v, w) \Theta((x-y) \cdot(w-v))\right] \mathrm{d} y \mathrm{~d} w \tag{158}
\end{align*}
$$

To arrive at the above result in the collision term, observe that, for fixed $x$ and $y$, we, with the help of Equations (5) and (7) have, for $i, j>1, i \neq j$ :

$$
\begin{align*}
& \int F\left(\boldsymbol{v}_{i}, v_{j}\right) \psi\left(\boldsymbol{v}_{i}^{\prime}, v_{j}^{\prime}\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-v_{i}\right) \Theta\left(\left(x_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{j}-v_{i}\right)\right) \mathrm{d} S_{2 \ldots K}= \\
& =\int F\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right) \psi\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|x_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{i}^{\prime}-\boldsymbol{v}_{j}^{\prime}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}^{\prime}-\boldsymbol{v}_{j}^{\prime}\right)\right) \mathrm{d} S_{2 \ldots K}= \\
& =-\int F\left(\boldsymbol{v}_{i}^{\prime}, \boldsymbol{v}_{j}^{\prime}\right) \psi\left(\boldsymbol{v}_{i}, \boldsymbol{v}_{j}\right) \frac{\boldsymbol{x}_{i}-\boldsymbol{x}_{j}}{\left\|\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right\|} \cdot\left(\boldsymbol{v}_{j}-\boldsymbol{v}_{i}\right) \Theta\left(\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right) \cdot\left(\boldsymbol{v}_{i}-\boldsymbol{v}_{j}\right)\right) \mathrm{d} S_{2 \ldots K}, \tag{159}
\end{align*}
$$

which cancels out all terms in the collision that do not involve that particular sphere over which the integration does not occur. Next, we switch to the spherical coordinate system and replace

$$
\begin{equation*}
y=x \pm \sigma r n, \quad \mathrm{~d} y=\sigma^{3} r^{2} \mathrm{~d} r \mathrm{~d} n \tag{160}
\end{equation*}
$$

where $r$ is the non-dimensional distance, $n$ is the unit vector, $\mathrm{d} n$ is the area element of the unit sphere, " + " is used in the term with $F^{(2)}\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)$, while " - " is used in the term with $F^{(2)}(\boldsymbol{v}, \boldsymbol{w})$. In the new variables, the identity in Equation (2) becomes Equation (22), and Equation (158) becomes the leading order identity of the corresponding BBGKY hierarchy [3-5] for Equation (91):

$$
\begin{align*}
& \frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \lambda \sigma^{2} \int \delta_{\alpha}(r-1) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
& \quad\left[F^{(2)}\left(\boldsymbol{x}, \boldsymbol{x}+\sigma r \boldsymbol{n}, \boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)-F^{(2)}(\boldsymbol{x}, \boldsymbol{x}-\sigma r \boldsymbol{n}, \boldsymbol{v}, \boldsymbol{w})\right] r^{2} \mathrm{~d} r \mathrm{~d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \tag{161}
\end{align*}
$$

At this point, it might be tempting to assume that $\alpha$ is small enough so that one can neglect the variations in $F^{(2)}$ within the contact zone and integrate in $\mathrm{d} r$; however, it is clearly not the case due to e.g., Equation (156). Thus, to proceed further, we first need to find an appropriate closure to $F^{(2)}$ above in Equation (161) in terms of the one-sphere marginal $f$.

### 6.1. Approximating the Two-Sphere Marginal via One-Sphere Marginals

Above in Equation (161) we arrive at the ubiquitous closure problem of molecular dynamics-the two-sphere marginal $F^{(2)}$ must be approximated via the one-sphere marginal $f$ in order to be able to transform the leading order BBGKY identity in Equation (161) into a standalone closed equation for $f$. In the mean field approximation [2] for long-range potentials, the joint density of two particles is approximated via the product of single-particle distributions. However, in our situation, the relation in Equation (156) for the steady state marginals does not permit such a direct factorization.

Instead, here we choose to relate the unsteady marginals $f$ and $F^{(2)}$ in the same way their steady counterparts are related in Equation (156) in the limit of infinitely many spheres:

$$
\begin{equation*}
F^{(2)}(t, \boldsymbol{x}, \boldsymbol{y}, \boldsymbol{w}, \boldsymbol{v}) \approx e^{-\lambda \Theta_{\alpha \sigma}(\sigma-\|x-y\|)} R^{(2)}(\|\boldsymbol{x}-\boldsymbol{y}\|) f(t, \boldsymbol{x}, \boldsymbol{v}) f(t, \boldsymbol{y}, \boldsymbol{w}) \tag{162}
\end{equation*}
$$

where $R^{(2)}$ depends parametrically on $\alpha$ and $\lambda: R^{(2)}=R_{\lambda, \alpha^{*}}^{(2)}$. Observe that the closure in Equation (162) becomes the exact solution if $F^{(2)}$ is the marginal distribution of the steady state as $K \rightarrow \infty$. If $F^{(2)}$ is not a marginal distribution of the steady state, but sufficiently close to it, then we can estimate the improvement of Equation (162) over Equation (156) via the following appropriately generalized version of the marginal formula for the Kullback-Leibler divergence [40,41].

Let $F\left(z_{1}, \ldots, z_{K}\right)$ be a probability density, $\psi\left(z_{1}, \ldots, z_{K}\right)>0$ be an integrable function, and $p_{1}\left(z_{1}\right), \ldots, p_{K}\left(z_{K}\right)$ be a set of probability densities for $z_{i} \in \mathbb{R}^{d}, d>0$. Then, the following identity holds for the Kullback-Leibler divergence $P$ :

$$
\begin{equation*}
P\left(F, \psi \prod_{i=1}^{K} p_{i}\left(z_{i}\right)\right)=P\left(F, \psi \prod_{i=1}^{K} f_{i}\left(z_{i}\right)\right)+\sum_{i=1}^{N} P\left(f_{i}, p_{i}\right) \tag{163}
\end{equation*}
$$

where $f_{i}\left(\boldsymbol{z}_{i}\right)$ is the $i$ th marginal of $F$. Indeed, observe that

$$
\begin{align*}
P\left(F, \psi \prod_{i=1}^{K} p_{i}\left(\boldsymbol{z}_{i}\right)\right) & =\int F\left[\ln F-\ln \psi-\ln \left(\prod_{i=1}^{K} p_{i}\left(\boldsymbol{z}_{i}\right)\right)\right] \mathrm{d} \mathbf{Z}= \\
& =\int F\left[\ln F-\ln \left(\psi \prod_{i=1}^{K} f_{i}\left(\boldsymbol{z}_{i}\right)\right)+\ln \left(\prod_{i=1}^{K} f_{i}\left(\boldsymbol{z}_{i}\right)\right)-\ln \left(\prod_{i=1}^{K} p_{i}\left(\boldsymbol{z}_{i}\right)\right)\right] \mathrm{d} \mathbf{Z}, \tag{164}
\end{align*}
$$

whereupon Equation (163) follows from the first and second pair of terms in the integrand.
The second term in the right-hand side of Equation (163) is always nonnegative, which bounds the left-hand side below as:

$$
\begin{equation*}
P\left(F, \psi \prod_{i=1}^{K} p_{i}\left(\boldsymbol{z}_{i}\right)\right) \geq P\left(F, \psi \prod_{i=1}^{K} f_{i}\left(\boldsymbol{z}_{i}\right)\right) . \tag{165}
\end{equation*}
$$

We can apply the formula above to our set-up by setting $z_{1}=(x, v), z_{2}=(y, w), F=F^{(2)}$, $\psi=e^{-\lambda \Theta_{\alpha \sigma}(\sigma-\|x-y\|)} R^{(2)}(\|x-y\|)$, and $p_{1}=p_{2}=\bar{F}^{(1)}$ from Equation (156). One should note, however, that the right-hand side of Equation (165) is not necessarily the Kullback-Leibler divergence due to the fact that Equation (162) (which is the second argument of $P$ ) is not necessarily normalized to 1 for large $\sigma$. However, for a dilute gas (that is, when $\sigma$ is much smaller than the average distance between the spheres), the second argument of $P$ becomes normalized to 1 , so that the right-hand side of the Kullback-Leibler bound in Equation (165) indeed involves the probability densities in such a
case. Additionally, the closure in Equation (162) preserves the proportionality relations between the velocity moments in $v$ or $w$ variables separately.

Now, taking into account that $\|\boldsymbol{x}-\boldsymbol{y}\|=\sigma r$ above in Equation (161), we have

$$
\begin{equation*}
\Theta_{\alpha \sigma}(\sigma-\|\boldsymbol{x}-\boldsymbol{y}\|)=\Theta_{\alpha \sigma}(\sigma-\sigma r)=\Theta_{\alpha}(1-r) \tag{166}
\end{equation*}
$$

and, with the approximation in Equation (162), the forward equation for the marginals in Equation (161) becomes

$$
\begin{align*}
\frac{\partial f}{\partial t}+\boldsymbol{v} \cdot \frac{\partial f}{\partial x}=(K-1) \lambda \sigma^{2} \int R_{\lambda, \alpha}^{(2)}(\sigma r) e^{-\lambda \Theta_{\alpha}(1-r)} \delta_{\alpha}(r-1) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
{\left[f\left(\boldsymbol{x}, \boldsymbol{v}^{\prime}\right) f\left(\boldsymbol{x}+\sigma r \boldsymbol{n}, \boldsymbol{w}^{\prime}\right)-f(\boldsymbol{x}, \boldsymbol{v}) f(\boldsymbol{x}-\sigma r \boldsymbol{n}, \boldsymbol{w})\right] r^{2} \mathrm{~d} r \mathrm{~d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} } \tag{167}
\end{align*}
$$

### 6.2. Thin Contact Zone and Impenetrable Spheres

Above in Equation (167), we can formally assume that the contact zone is "thin", that is, $\alpha \rightarrow 0$ so that, for the values of $r$ for which $\delta_{\alpha}(r-1)>0$, we have $f(x \pm \sigma r n) \rightarrow f(x \pm \sigma n), R_{\lambda, \alpha}^{(2)}(\sigma r) \rightarrow R_{\lambda, 0}^{(2)}(\sigma)$. In such a case, the integral over $\mathrm{d} r$ involves only the mollifier $\delta_{\alpha}(r-1)$ and its antiderivative, and thus can be integrated across the thin contact zone exactly:

$$
\begin{equation*}
\lambda \int e^{-\lambda \Theta_{\alpha}(1-r)} \delta_{\alpha}(r-1) \mathrm{d} r=1-e^{-\lambda} \tag{168}
\end{equation*}
$$

This simplifies the forward equation in Equation (167) to

$$
\begin{align*}
& \frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \sigma^{2}\left(1-e^{-\lambda}\right) R_{\lambda, 0}^{(2)}(\sigma) \int \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
& {\left[f\left(\boldsymbol{x}, \boldsymbol{v}^{\prime}\right) f\left(\boldsymbol{x}+\sigma \boldsymbol{n}, \boldsymbol{w}^{\prime}\right)-f(\boldsymbol{x}, \boldsymbol{v}) f(\boldsymbol{x}-\sigma \boldsymbol{n}, \boldsymbol{w})\right] \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} } \tag{169}
\end{align*}
$$

Now, observe that the factor $\left(1-e^{-\lambda}\right)$ in front of the equation above is the probability that at least one jump arrives in the point process during the collision. Sending the intensity of the point process $\lambda \rightarrow \infty$ sets this probability to 1 , which means that the spheres can be considered impenetrable. The resulting forward equation for a single impenetrable sphere is given via

$$
\begin{align*}
& \frac{\partial f}{\partial t}+v \cdot \frac{\partial f}{\partial x}=(K-1) \sigma^{2} R_{\infty, 0}^{(2)}(\sigma) \int n \cdot(\boldsymbol{w}-v) \Theta(n \cdot(\boldsymbol{w}-v)) \\
& {\left[f\left(x, v^{\prime}\right) f\left(x+\sigma \boldsymbol{n}, \boldsymbol{w}^{\prime}\right)-f(x, v) f(x-\sigma n, w)\right] \mathrm{d} \boldsymbol{n} \mathrm{~d} w . } \tag{170}
\end{align*}
$$

This is a variant of the Enskog equation for hard spheres [16-20], which we henceforth refer to as "the Enskog equation", even though it is not identical to what was originally proposed by Enskog (in the conventional Enskog equation, the term in front of the collision integral in Equation (170) is a phenomenological correlation function).

Apparently, the collision integral closure above in the Enskog equation in Equation (170) becomes exact if $f$ is the single-sphere marginal distribution of the steady state $\bar{F}_{K}$ in Equation (110), in the limit as the number of spheres $K \rightarrow \infty$, the width of the mollifier $\alpha \rightarrow 0$, and $\lambda \rightarrow \infty$.

Observe that, if the gas is dilute (that is, the diameter $\sigma$ of a sphere is much smaller than the average distance between the spheres), then we can neglect the higher order terms in $\sigma$ in comparison to the leading order. This sets $R_{\infty, 0}^{(2)}(\sigma)=1, f(x \pm \sigma n, v)=f(x, v)$, and the Enskog equation in Equation (170) becomes the Boltzmann equation in Equation (33).

## 7. The Fluid Dynamics of the Enskog Equation in a Physical Hydrodynamic Limit

Above, we introduce a new, random model for hard sphere collision which avoids Loschmidt's objection [14], and where the form of the collision integral is a property of the forward Kolmogorov
equation of the full multisphere dynamics. We show that, under suitable assumptions, this new model leads to the Boltzmann equation in Equation (33). However, one of the key differences between our derivation of the Boltzmann equation and its conventional derivation [9,11] is that, as a precursor to the Boltzmann equation, we also obtain the Enskog equation in Equation (170) in a systematic fashion, by reducing and simplifying the forward Kolmogorov equation of the multisphere dynamics under suitable assumptions.

The hydrodynamic limit of the Boltzmann equation is a well-researched area (see, for example, [13] and references therein), and it is known that the equations for the velocity moments of the Boltzmann equation lead to the Euler [42] and Navier-Stokes [6,10] equations. However, the hydrodynamic limit of the Enskog equation is a less researched topic, although there are works on this subject as well (see [19] and references therein).

Below, we derive the fluid dynamics equations of the Enskog equation in Equation (170) in the hydrodynamic limit (that is, as the diameter $\sigma \rightarrow 0$ ) which, on the one hand, is physically plausible, and, on the other hand, produces additional non-vanishing terms in the resulting fluid dynamics equations.

### 7.1. The Mass-Weighted Equation and the Hydrodynamic Limit

Here, we endow each sphere with a mass density $\rho_{s p}$, which is a constant parameter. Above, $f$ is the distribution density of a single sphere in the $K$-sphere system, and thus is normalized to 1 . However, for the subsequent derivation of the fluid dynamics equations, it is more convenient to normalize the density by the total mass of the system. With this, we define the mass density $g$ via

$$
\begin{equation*}
g(t, x, v)=\frac{1}{6} K \pi \rho_{s p} \sigma^{3} f(t, x, v) \tag{171}
\end{equation*}
$$

where the factor in front of $f$ above is the total mass of the system of $K$ spheres, each of diameter $\sigma$ and density $\rho_{s p}$. The corresponding mass-weighted form of the Enskog equation for $K$ spheres in Equation (170) is, therefore,

$$
\begin{align*}
& \frac{\partial g}{\partial t}+\boldsymbol{v} \cdot \frac{\partial g}{\partial x}=\frac{K-1}{K} \frac{6 R_{\infty, 0}^{(2)}(\sigma)}{\pi \rho_{s p} \sigma} \int \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(n \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
& {\left[g\left(\boldsymbol{x}, \boldsymbol{v}^{\prime}\right) g\left(\boldsymbol{x}+\sigma \boldsymbol{n}, \boldsymbol{w}^{\prime}\right)-g(\boldsymbol{x}, \boldsymbol{v}) g(\boldsymbol{x}-\sigma \boldsymbol{n}, \boldsymbol{w})\right] \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} . } \tag{172}
\end{align*}
$$

Here, observe that, as $K \rightarrow \infty$, the ratio $(K-1) / K \rightarrow 1$, and we need to decide how the remaining coefficient in front of the integral behaves as $\sigma \rightarrow 0$. Here, we argue that the density $\rho_{s p}$ of the spheres should be kept constant; in Figure 1, we plot the cubic root of the atomic mass versus the atomic diameter for noble gases, so that the straight line of the least squares fit approximates the hypothetical situation with constant density of the spheres $\rho_{s p}$. Observe that, with the exception of neon, all atoms align on the least squares fit line, which indicates that the density of atoms of the actual noble gases tends to be constant. The resulting Enskog equation is given via

$$
\begin{align*}
& \frac{\partial g}{\partial t}+v \cdot \frac{\partial g}{\partial x}=\frac{6 R_{\infty, 0}^{(2)}(\sigma)}{\pi \rho_{s p} \sigma} \int n \cdot(\boldsymbol{w}-v) \Theta(n \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
& {\left[g\left(x, v^{\prime}\right) g\left(x+\sigma n, w^{\prime}\right)-g(x, v) g(x-\sigma n, w)\right] \mathrm{d} n \mathrm{~d} \boldsymbol{w}, } \tag{173}
\end{align*}
$$

where the factor in front of the collision integral above behaves as $\sim \sigma^{-1}$ in the constant-density hydrodynamic limit $K \rightarrow \infty, \sigma \rightarrow 0, \rho_{s p}=$ const.


Figure 1. Atomic diameter vs. atomic mass for noble gases. The values of the atomic diameter are taken from [23]. The values of the atomic mass are taken from the standard periodic table of elements.

### 7.2. The Euler Equations

With $\sigma$ being treated as a small parameter, a standard way to look for a solution of Equation (173) is to expand $g$ in powers of $\sigma$ [19]:

$$
\begin{equation*}
g(t, x, v, \sigma)=g_{0}(t, x, v)+\sigma g_{1}(t, x, v)+\sigma^{2} g_{2}(t, x, v)+\ldots \tag{174}
\end{equation*}
$$

and then recover the expansion terms sequentially starting from the leading order. We also note that $R_{\infty, 0}^{(2)}(0)=1$, and thus

$$
\begin{equation*}
R_{\infty, 0}^{(2)}(\sigma)=1+\left.\sigma \frac{\mathrm{d}}{\mathrm{~d} \sigma} R_{\infty, 0}^{(2)}(\sigma)\right|_{\sigma=0}+\left.\frac{\sigma^{2}}{2} \frac{\mathrm{~d}^{2}}{\mathrm{~d} \sigma^{2}} R_{\infty, 0}^{(2)}(\sigma)\right|_{\sigma=0}+\ldots \tag{175}
\end{equation*}
$$

As a result, in the leading order of $\sigma$, we find

$$
\begin{equation*}
\int \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left[g_{0}\left(\boldsymbol{v}^{\prime}\right) g_{0}\left(\boldsymbol{w}^{\prime}\right)-g_{0}(\boldsymbol{v}) g_{0}(\boldsymbol{w})\right] \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w}=0 \tag{176}
\end{equation*}
$$

which is the usual Boltzmann collision integral [6,9,10,13]. This means that the leading order expansion term $g_{0}$ is the standard Maxwell-Boltzmann thermodynamic equilibrium state,

$$
\begin{equation*}
g_{0}(t, x, v)=\frac{\rho}{(2 \pi \theta)^{3 / 2}} \exp \left(-\frac{\|v-u\|^{2}}{2 \theta}\right) \tag{177}
\end{equation*}
$$

where the mass density $\rho(t, x)$, average velocity $\boldsymbol{u}(t, x)$ and kinetic temperature $\theta(t, x)$ are given by the following mass-weighted moments of molecular velocity:

$$
\begin{equation*}
\rho=\int g_{0} \mathrm{~d} v, \quad \rho \boldsymbol{u}=\int \boldsymbol{v} g_{0} \mathrm{~d} v, \quad \rho\left(\frac{1}{2}\|\boldsymbol{u}\|^{2}+\frac{3}{2} \theta\right)=\rho \epsilon=\frac{1}{2} \int\|\boldsymbol{v}\|^{2} g_{0} \mathrm{~d} v \tag{178}
\end{equation*}
$$

Above, for further convenience, we additionally denote the average kinetic energy $\epsilon$ via

$$
\begin{equation*}
\epsilon=\frac{1}{2}\|\boldsymbol{u}\|^{2}+\frac{3}{2} \theta \tag{179}
\end{equation*}
$$

For the next order in $\sigma$, we have

$$
\begin{align*}
\frac{\partial g_{0}}{\partial t}+\boldsymbol{v} \cdot \frac{\partial g_{0}}{\partial \boldsymbol{x}} & =\frac{6}{\pi \rho_{s p}} \int \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left[\left(g_{0}\left(\boldsymbol{v}^{\prime}\right) g_{1}\left(\boldsymbol{w}^{\prime}\right)+g_{1}\left(\boldsymbol{v}^{\prime}\right) g_{0}\left(\boldsymbol{w}^{\prime}\right)-\right.\right. \\
& \left.\left.-g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})-g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})\right)+\boldsymbol{n} \cdot\left(g_{0}\left(\boldsymbol{v}^{\prime}\right) \frac{\partial g_{0}}{\partial \boldsymbol{x}}\left(\boldsymbol{w}^{\prime}\right)+g_{0}(\boldsymbol{v}) \frac{\partial g_{0}}{\partial \boldsymbol{x}}(\boldsymbol{w})\right)\right] \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \tag{180}
\end{align*}
$$

Observe that the form of the equation above generally does not guarantee that $g_{0}$ retains its Gaussian form in Equation (177), due to the additive collision integral in the right-hand side of Equation (180). Thus, we have to resort to a weaker formulation of Equation (180) in the form of velocity moments of $1, v$, and $\|v\|^{2}$, in the same manner as in $[10,19]$, which allows to exclude the terms with $g_{1}$ completely and close the equation with respect to $\rho, u$, and $\theta$.

First, for two arbitrary functions $\psi(\boldsymbol{v})$ and $h(v, w)$, the latter with the symmetry property $h(\boldsymbol{v}, \boldsymbol{w})=h(\boldsymbol{w}, \boldsymbol{v})$, observe the following identities (assuming that the integrals are bounded):

$$
\begin{align*}
& \int \psi(v) h(v, w) n \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} v= \\
& =\int \psi(\boldsymbol{w}) h(\boldsymbol{v}, \boldsymbol{w}) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}
\end{aligned} \quad \begin{aligned}
& \int \psi\left(\boldsymbol{v}^{\prime}\right) h(v, w) n \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=  \tag{181a}\\
& =\int \psi\left(\boldsymbol{w}^{\prime}\right) h(\boldsymbol{v}, \boldsymbol{w}) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}
\end{align*}
$$

Above, with help of the symmetry of $h(v, w)$, we rename $v$ as $w$ and vice versa, and also change the sign of $n$.

Second, observe the following chain of identities:

$$
\begin{align*}
& \int \psi(v) h\left(\boldsymbol{v}^{\prime}, w^{\prime}\right) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
&=\int \psi(\boldsymbol{v}) h\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right) \boldsymbol{n} \cdot\left(\boldsymbol{v}^{\prime}-\boldsymbol{w}^{\prime}\right) \Theta\left(\boldsymbol{n} \cdot\left(\boldsymbol{v}^{\prime}-\boldsymbol{w}^{\prime}\right)\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
&=\int \psi(\boldsymbol{v}) h\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right) \boldsymbol{n} \cdot\left(\boldsymbol{w}^{\prime}-\boldsymbol{v}^{\prime}\right) \Theta\left(\boldsymbol{n} \cdot\left(\boldsymbol{w}^{\prime}-\boldsymbol{v}^{\prime}\right)\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w}^{\prime} \mathrm{d} \boldsymbol{v}^{\prime}= \\
&=\int \psi\left(\boldsymbol{v}^{\prime}\right) h(\boldsymbol{v}, \boldsymbol{w}) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v} \tag{182}
\end{align*}
$$

Above, in the first identity, we use Equation (5); in the second identity, we use Equation (7) and also replace $\boldsymbol{n}$ with $-\boldsymbol{n}$, using the fact that $\boldsymbol{v}^{\prime}$ and $\boldsymbol{w}^{\prime}$ in Equation (22) are invariant with respect to the sign of $n$; and, in the last identity, we rename $v^{\prime}$ as $v, w^{\prime}$ as $w$, and vice versa.

Third, with help of Equations (181) and (182), we can write

$$
\begin{align*}
& \int \psi(\boldsymbol{v})\left(h\left(\boldsymbol{v}^{\prime}, \boldsymbol{w}^{\prime}\right)-h(\boldsymbol{v}, \boldsymbol{w})\right) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
& =\int\left(\psi\left(\boldsymbol{v}^{\prime}\right)-\psi(\boldsymbol{v})\right) h(\boldsymbol{v}, \boldsymbol{w}) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
& =\frac{1}{2} \int\left(\psi\left(\boldsymbol{v}^{\prime}\right)+\psi\left(\boldsymbol{w}^{\prime}\right)-\psi(\boldsymbol{v})-\psi(\boldsymbol{w})\right) h(\boldsymbol{v}, \boldsymbol{w}) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v} \tag{183}
\end{align*}
$$

For a constant $\psi$, the integral above is, obviously, zero regardless of the choice of $h(v, w)$. Additionally, recall that $v^{\prime}$ and $w^{\prime}$ in Equation (22) are chosen so that the momentum and energy of any two spheres are preserved during the collision. Thus, if $\psi(v)$ is set to $v$ or $\|v\|^{2}$ (or any linear
combination thereof, including an additive constant), then the integral above in Equation (183) is also zero irrespective of the choice of $h(\boldsymbol{v}, \boldsymbol{w})$.

Computing the velocity moments in Equation (178) of both sides of Equation (180), and setting $h(\boldsymbol{v}, \boldsymbol{w})=g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})+g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})$, we arrive at

$$
\begin{gather*}
\frac{\partial \rho}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho \boldsymbol{u})=0, \quad \frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\theta \boldsymbol{I}\right)\right)=\mathcal{C}[\boldsymbol{v}]  \tag{184a}\\
\frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho(\epsilon+\theta) \boldsymbol{u})=\frac{1}{2} \mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right] \tag{184b}
\end{gather*}
$$

where by $\mathcal{C}$ we denote the corresponding collision integrals

$$
\begin{gather*}
\mathcal{C}[\boldsymbol{v}]=\frac{6}{\pi \rho_{s p}} \int\left(\boldsymbol{v}-\boldsymbol{v}^{\prime}\right) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial x} g_{0}(\boldsymbol{v}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}  \tag{185a}\\
\mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right]=\frac{6}{\pi \rho_{s p}} \int\left(\|\boldsymbol{v}\|^{2}-\left\|\boldsymbol{v}^{\prime}\right\|^{2}\right) \boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) n \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial x} g_{0}(\boldsymbol{v}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v} \tag{185b}
\end{gather*}
$$

Next, recalling from Equation (22) that

$$
\begin{equation*}
v-v^{\prime}=-(n \cdot(w-v)) n, \quad\|v\|^{2}-\left\|v^{\prime}\right\|^{2}=-(n \cdot(w+v))(n \cdot(w-v)) \tag{186}
\end{equation*}
$$

we further obtain

$$
\begin{gather*}
\mathcal{C}[\boldsymbol{v}]=-\frac{6}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial x} g_{0}(\boldsymbol{v}) \boldsymbol{n} \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}  \tag{187a}\\
\mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right]=-\frac{6}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}+\boldsymbol{v}))(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}} g_{0}(\boldsymbol{v}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v} \tag{187b}
\end{gather*}
$$

With $g_{0}$ provided explicitly in Equation (177), the expressions above are integrated exactly into

$$
\begin{equation*}
\mathcal{C}[\boldsymbol{v}]=-\frac{4}{\rho_{s p}} \frac{\partial}{\partial x}\left(\rho^{2} \theta\right), \quad \mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right]=-\frac{8}{\rho_{s p}} \frac{\partial}{\partial x} \cdot\left(\rho^{2} \theta \boldsymbol{u}\right) \tag{188}
\end{equation*}
$$

Substituting Equation (188) into Equation (184), we arrive at

$$
\begin{gather*}
\frac{\partial \rho}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho \boldsymbol{u})=0, \quad \frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\left(1+\frac{4 \rho}{\rho_{s p}}\right) \theta \boldsymbol{I}\right)\right)=0  \tag{189a}\\
\frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial x} \cdot\left(\rho\left(\epsilon+\left(1+\frac{4 \rho}{\rho_{s p}}\right) \theta\right) \boldsymbol{u}\right)=0 \tag{189b}
\end{gather*}
$$

The equations above are also obtained in [19], where they are referred to as the "Enskog-Euler" equations. With $\rho / \rho_{s p} \ll 1$ (dilute gas), the equations in Equation (189) become the conventional Euler equations [13,42]. However, observe that, in the physical constant-density limit, considered here, the additional terms are formally non-vanishing despite the fact that the diameter of the sphere $\sigma \rightarrow 0$.

### 7.3. The Newton and Fourier Laws

Here, we examine the next-order term $g_{1}$ from Equation (174). With $g_{0}$ already computed above in Equation (177), we can use Equation (180) to determine the properties of $g_{1}$. However, observe that, while the form of $g_{0}$ is given uniquely by the leading order identity in Equation (176), such a strict constraint on $g_{1}$ is not present-indeed, the next-order equation in Equation (180) does not call for any particular restriction on $g_{1}$. Therefore, the form of $g_{1}$ in the expansion in Equation (174) can be chosen as necessary.

Below, we follow Grad [10] and assume that the full density $g$ can be expressed in the form of the Hermite polynomial expansion around the Gaussian density $g_{0}$. This form is chosen solely due to convenience-as shown below, all subsequent moment integrals that arise are expressed in terms of elementary functions. However, this is not a unique choice for $g$; for an alternative, see, for example, the work in [43], where the form of density is chosen to maximize the Boltzmann entropy under given moment constraints. The downside of the latter approach is that the moment integrals of the solutions of constrained maximum entropy optimization problems cannot, in general, be expressed in terms of elementary functions.

Following Grad [10], we choose $g_{1}$ in the form of the Hermite polynomial, which includes only the following powers of $v: v$ itself, $v v^{T}$, and $\|v\|^{2} v$. We impose the following orthogonality conditions on $g_{1}$ :

$$
\begin{equation*}
\int g_{1} \mathrm{~d} v=0, \quad \int v g_{1} \mathrm{~d} v=0, \quad \int\|v\|^{2} g_{1} \mathrm{~d} v=0 \tag{190}
\end{equation*}
$$

The following higher-order centered moments of $g_{1}$ is used as the new variables in addition to $\rho$, $\boldsymbol{u}$ and $\theta$ :

$$
\begin{equation*}
\int(v-u)(v-u)^{T} g_{1} \mathrm{~d} v=\rho \boldsymbol{S}, \quad \frac{1}{2} \int\|\boldsymbol{v}-\boldsymbol{u}\|^{2}(v-u) g_{1} \mathrm{~d} v=\rho \boldsymbol{q} \tag{191}
\end{equation*}
$$

Above, $\boldsymbol{S}$ and $\boldsymbol{q}$ are called the stress and heat flux, respectively. Note that the symmetric matrix $\boldsymbol{S}$ has zero trace due to the orthogonality requirements above. Then, according to the calculations in $[10,44], g_{1}$ is given by the following Hermite polynomial in $v$ :

$$
\begin{equation*}
g_{1}=\frac{1}{\theta^{2}} g_{0}\left((\boldsymbol{u}-\boldsymbol{v}) \cdot \boldsymbol{q}+\frac{1}{2}(\boldsymbol{v}-\boldsymbol{u})^{T} \boldsymbol{S}(\boldsymbol{v}-\boldsymbol{u})+\frac{1}{5 \theta}\|\boldsymbol{v}-\boldsymbol{u}\|^{2}(\boldsymbol{v}-\boldsymbol{u}) \cdot \boldsymbol{q}\right) . \tag{192}
\end{equation*}
$$

With the ansatz in Equation (192), we only need to relate the stress $\boldsymbol{S}$ and heat flux $\boldsymbol{q}$ to the leading order approximation $g_{0}$ (that is, to $\rho, \boldsymbol{u}$ and $\theta$ ). To achieve that, we proceed via integrating Equation (180) against the velocity moments $v v^{T}$ and $\|v\|^{2} v$ :

$$
\begin{align*}
\frac{\partial}{\partial t} \int\left(\boldsymbol{v} \boldsymbol{v}^{T}\right) g_{0} \mathrm{~d} v+\frac{\partial}{\partial \boldsymbol{x}} \cdot \int \boldsymbol{v}\left(\boldsymbol{v} \boldsymbol{v}^{T}\right) g_{0} \mathrm{~d} \boldsymbol{v} & =\mathcal{C}\left[\boldsymbol{v} \boldsymbol{v}^{T}\right]  \tag{193a}\\
\frac{\partial}{\partial t} \int\|\boldsymbol{v}\|^{2} \boldsymbol{v} g_{0} \mathrm{~d} \boldsymbol{v}+\frac{\partial}{\partial \boldsymbol{x}} \cdot \int\|\boldsymbol{v}\|^{2}\left(\boldsymbol{v} \boldsymbol{v}^{T}\right) g_{0} \mathrm{~d} \boldsymbol{v} & =\mathcal{C}\left[\|\boldsymbol{v}\|^{2} \boldsymbol{v}\right] \tag{193b}
\end{align*}
$$

Then, we switch to the centered moments:

$$
\begin{gather*}
\boldsymbol{v} \boldsymbol{v}^{T}=(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}+(\boldsymbol{v}-\boldsymbol{u}) \boldsymbol{u}^{T}+\boldsymbol{u}(\boldsymbol{v}-\boldsymbol{u})^{T}+\boldsymbol{u} \boldsymbol{u}^{T}  \tag{194a}\\
\|\boldsymbol{v}\|^{2} \boldsymbol{v}=\|\boldsymbol{v}-\boldsymbol{u}\|^{2}(\boldsymbol{v}-\boldsymbol{u})+2(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T} \boldsymbol{u}+\|\boldsymbol{v}\|^{2} \boldsymbol{u}+\|\boldsymbol{u}\|^{2} \boldsymbol{v}-\|\boldsymbol{u}\|^{2} \boldsymbol{u} \tag{194b}
\end{gather*}
$$

Substituting the above decompositions and using the moment definitions of $\rho, \boldsymbol{u}$ and $\theta$, we arrive at

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\theta \boldsymbol{I}\right)\right)+\frac{\partial(\rho \theta \boldsymbol{u})}{\partial \boldsymbol{x}}+\frac{\partial(\rho \theta \boldsymbol{u})^{T}}{\partial \boldsymbol{x}}+\left(\frac{\partial}{\partial \boldsymbol{x}} \cdot(\rho \theta \boldsymbol{u})\right) \boldsymbol{I}+\frac{\partial}{\partial \boldsymbol{x}} \cdot(\rho \boldsymbol{u} \otimes \boldsymbol{u} \otimes \boldsymbol{u})= \\
&=\mathcal{C}\left[(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right]+\mathcal{C}[\boldsymbol{v}] \boldsymbol{u}^{T}+\boldsymbol{u C}[\boldsymbol{v}]^{T} \tag{195a}
\end{align*}
$$

$$
\begin{align*}
2 \frac{\partial}{\partial t}(\rho(\epsilon+\theta) \boldsymbol{u})+\frac{\partial}{\partial \boldsymbol{x}} & \cdot\left(5 \rho \theta^{2} \boldsymbol{I}+7 \rho \theta \boldsymbol{u} \boldsymbol{u}^{T}+\rho \theta\|\boldsymbol{u}\|^{2} \boldsymbol{I}+\rho\|\boldsymbol{u}\|^{2} \boldsymbol{u} \boldsymbol{u}^{T}\right)= \\
& =\mathcal{C}\left[\|\boldsymbol{v}-\boldsymbol{u}\|^{2}(\boldsymbol{v}-\boldsymbol{u})\right]+2 \mathcal{C}\left[(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right] \boldsymbol{u}+\mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right] \boldsymbol{u}+\|\boldsymbol{u}\|^{2} \mathcal{C}[\boldsymbol{v}] \tag{195b}
\end{align*}
$$

We exclude the time derivatives above via the Enskog-Euler equations in Equation (189):

$$
\begin{align*}
& \mathcal{C}\left[(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right]= \rho \theta\left(\frac{\partial u}{\partial x}+\frac{\partial u^{T}}{\partial x}-\frac{2}{3}\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right)\right)+\frac{1}{3} \mathcal{C}\left[\|\boldsymbol{v}\|^{2}\right] \boldsymbol{I}-\frac{2}{3} \boldsymbol{u}^{T} \mathcal{C}[v] \boldsymbol{I},  \tag{196a}\\
& \mathcal{C}\left[\|v-u\|^{2}(v-u)\right]=5 \rho \theta \frac{\partial \theta}{\partial x}+5 \theta \mathcal{C}[v] \tag{196b}
\end{align*}
$$

Recalling the form of the collision terms, and using Equation (22), we further obtain

$$
\begin{gather*}
\frac{6}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2}\left((\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n} \boldsymbol{n}^{T}+\boldsymbol{n}(\boldsymbol{v}-\boldsymbol{u})^{T}+(\boldsymbol{v}-\boldsymbol{u}) \boldsymbol{n}^{T}\right) \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
\left(g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})+g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})-\boldsymbol{n} \cdot g_{0}(\boldsymbol{v}) \frac{\partial g_{0}}{\partial \boldsymbol{x}}(\boldsymbol{w})\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
=\rho \theta\left(\frac{\partial \boldsymbol{u}}{\partial \boldsymbol{x}}+\frac{\partial \boldsymbol{u}^{T}}{\partial \boldsymbol{x}}-\frac{2}{3}\left(\frac{\partial}{\partial \boldsymbol{x}} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right)-\frac{8}{3 \rho_{s p}} \rho^{2} \theta\left(\frac{\partial}{\partial \boldsymbol{x}} \cdot \boldsymbol{u}\right) \boldsymbol{I},  \tag{197a}\\
\frac{6}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2}\left[(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \boldsymbol{n}+(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left(\boldsymbol{I}+2 \boldsymbol{n} \boldsymbol{n}^{T}\right)(\boldsymbol{v}-\boldsymbol{u})+\right. \\
\left.+\left(\|\boldsymbol{v}-\boldsymbol{u}\|^{2} \boldsymbol{I}+2(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right) \boldsymbol{n}\right] \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
\left(g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})+g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})-\boldsymbol{n} \cdot g_{0}(\boldsymbol{v}) \frac{\partial g_{0}}{\partial \boldsymbol{x}}(\boldsymbol{w})\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=5 \rho \theta \frac{\partial \theta}{\partial \boldsymbol{x}}-\frac{20}{\rho_{s p}} \theta \frac{\partial\left(\rho^{2} \theta\right)}{\partial \boldsymbol{x}} \tag{197b}
\end{gather*}
$$

Taking into account the explicit formulas for $g_{0}$ in Equation (177) and $g_{1}$ in Equation (192), we obtain the following exact relations via direct integration:

$$
\begin{align*}
& \int(n \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2}\left((\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n} \boldsymbol{n}^{T}+\boldsymbol{n}(\boldsymbol{v}-\boldsymbol{u})^{T}+(\boldsymbol{v}-\boldsymbol{u}) \boldsymbol{n}^{T}\right) \\
& \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left(g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})+g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=-\frac{16 \sqrt{\pi}}{5} \rho^{2} \sqrt{\theta} \boldsymbol{S},  \tag{198a}\\
& \int(n \cdot(\boldsymbol{w}-v))^{2}\left((n \cdot(\boldsymbol{w}-v)) n n^{T}+\boldsymbol{n}(v-u)^{T}+(v-u) n^{T}\right) \\
& \Theta(n \cdot(\boldsymbol{w}-\boldsymbol{v})) n \cdot g_{0}(\boldsymbol{v}) \frac{\partial g_{0}}{\partial x}(\boldsymbol{w}) \mathrm{d} \boldsymbol{n} \mathrm{~d} w \mathrm{~d} v=\frac{4 \pi}{15} \rho^{2} \theta\left(\frac{\partial u}{\partial x}+\frac{\partial \boldsymbol{u}^{T}}{\partial x}+\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right),  \tag{198b}\\
& \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2}\left[(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \boldsymbol{n}+(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left(\boldsymbol{I}+2 \boldsymbol{n} \boldsymbol{n}^{T}\right)(\boldsymbol{v}-\boldsymbol{u})+\right. \\
& \left.+\left(\|\boldsymbol{v}-\boldsymbol{u}\|^{2} \boldsymbol{I}+2(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right) n\right] \Theta(n \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
& \left(g_{0}(\boldsymbol{v}) g_{1}(\boldsymbol{w})+g_{1}(\boldsymbol{v}) g_{0}(\boldsymbol{w})\right) \mathrm{d} n \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=-\frac{64 \sqrt{\pi}}{15} \rho^{2} \sqrt{\theta} \boldsymbol{q},  \tag{198c}\\
& \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2}\left[(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \boldsymbol{n}+(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))\left(\boldsymbol{I}+2 \boldsymbol{n} \boldsymbol{n}^{T}\right)(\boldsymbol{v}-\boldsymbol{u})+\right. \\
& \left.+\left(\|v-\boldsymbol{u}\|^{2} \boldsymbol{I}+2(\boldsymbol{v}-\boldsymbol{u})(\boldsymbol{v}-\boldsymbol{u})^{T}\right) n\right] \Theta(n \cdot(\boldsymbol{w}-v)) \\
& n \cdot g_{0}(\boldsymbol{v}) \frac{\partial g_{0}}{\partial x}(\boldsymbol{w}) \mathrm{d} n \mathrm{~d} w \mathrm{~d} v=\frac{10 \pi}{3} \theta \frac{\partial\left(\rho^{2} \theta\right)}{\partial x}+2 \pi \rho^{2} \theta \frac{\partial \theta}{\partial x} . \tag{198d}
\end{align*}
$$

Substituting the identities in Equation (198) into Equation (197), we express the stress $\boldsymbol{S}$ and heat flux $\boldsymbol{q}$ in terms of $\rho, \boldsymbol{u}$ and $\theta$ :

$$
\begin{align*}
\rho \boldsymbol{S} & =\left(1+\frac{8 \rho}{5 \rho_{s p}}\right)[\rho \boldsymbol{S}]_{B}=-\left(1+\frac{8 \rho}{5 \rho_{s p}}\right) \mu\left(\frac{\partial \boldsymbol{u}}{\partial x}+\frac{\partial \boldsymbol{u}^{T}}{\partial x}-\frac{2}{3}\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right)  \tag{199a}\\
\rho \boldsymbol{q} & =\left(1+\frac{12 \rho}{5 \rho_{s p}}\right)[\rho \boldsymbol{q}]_{B}=-\left(1+\frac{12 \rho}{5 \rho_{s p}}\right) \frac{15}{4} \mu \frac{\partial \theta}{\partial \boldsymbol{x}}, \quad \mu=\frac{5 \sqrt{\pi} \rho_{s p} \sigma}{96} \sqrt{\theta} \tag{199b}
\end{align*}
$$

where $\mu$ is the usual viscosity for the hard sphere gas $[10]$, and $[\rho \boldsymbol{S}]_{B},[\rho \boldsymbol{q}]_{B}$ denote the conventional Newton and Fourier law expressions for a dilute gas ( $\rho / \rho_{s p} \ll 1$ ), respectively, derived purely from the Boltzmann equation [6,10]. Observe that the viscosity $\mu$ is $O(\sigma)$. Therefore, to include viscous effects into the fluid dynamics equations, we have to retain the $O(\sigma)$-term in the expansion in Equation (174).

### 7.4. The Navier-Stokes Equations

Above, we compute the expansion terms $g_{0}$ and $g_{1}$ of Equation (174). The $g_{0}$ is given via the Maxwell-Boltzmann state in Equation (177), while the first-order correction $g_{1}$ is given via the Hermite polynomial in Equation (192), with the parameters $\boldsymbol{S}$ and $\boldsymbol{q}$ provided via the Newton and Fourier laws in Equation (199).

Now, instead of proceeding with the formal hydrodynamic limit as the sphere diameter $\sigma \rightarrow 0$, we assume that $\sigma$ is a finitely small constant. In such a case, we can truncate Equation (174) to the first two leading order terms (assuming that all $O\left(\sigma^{2}\right)$ terms are small enough so that they can be neglected). The result is known as the Grad state [10]:

$$
\begin{equation*}
g_{G r a d}=g_{0}+\sigma g_{1} . \tag{200}
\end{equation*}
$$

Then, we follow the same procedure for $g_{G r a d}$ as above for $g_{0}$; namely, we substitute $g_{G r a d}$ into Equation (180), and compute the transport equations for the mass, momentum and energy moments. Since $g_{G r a d}$ contains $O(\sigma)$ terms, the terms of the same order must be retained in the collision term of Equation (180). The resulting transport equation for the density is unchanged, however, the equations for the velocity and energy are now given via

$$
\begin{array}{r}
\begin{array}{r}
\frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\theta \boldsymbol{I}+\sigma \boldsymbol{S}\right)\right)=- \\
\rho_{s p}\left(1+\sigma R_{*}\right) \frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho^{2} \theta\right)+ \\
\\
+\frac{6 \sigma}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \boldsymbol{n} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
\left(-g_{1}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}}-g_{0}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{1}(\boldsymbol{w})}{\partial \boldsymbol{x}}+\frac{1}{2} g_{0}(\boldsymbol{v}) \boldsymbol{n}^{T} \frac{\partial^{2} g_{0}(\boldsymbol{w})}{\partial x^{2}} \boldsymbol{n}\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v} \\
\frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot(\rho((\epsilon+\theta) \boldsymbol{u}+\sigma \boldsymbol{S} \boldsymbol{u}+\sigma \boldsymbol{q}))=-\frac{4}{\rho_{s p}}\left(1+\sigma R_{*}\right) \frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho^{2} \theta \boldsymbol{u}\right)+ \\
\quad+\frac{3 \sigma}{\pi \rho_{s p}} \int(\boldsymbol{n} \cdot(\boldsymbol{w}+\boldsymbol{v}))(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \\
\left(-g_{1}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}}-g_{0}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{1}(\boldsymbol{w})}{\partial \boldsymbol{x}}+\frac{1}{2} g_{0}(\boldsymbol{v}) \boldsymbol{n}^{T} \frac{\partial^{2} g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}^{2}} \boldsymbol{n}\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}
\end{array}
\end{array}
$$

Above, for convenience, the constant $R_{*}$ denotes the derivative of $R_{\infty, 0}^{(2)}$ at zero:

$$
\begin{equation*}
R_{*}=\left.\frac{\mathrm{d}}{\mathrm{~d} \sigma} R_{\infty, 0}^{(2)}(\sigma)\right|_{\sigma=0} \tag{202}
\end{equation*}
$$

Due to the fact that both $g_{0}$ and $g_{1}$ are given explicitly via Equations (177) and (192), respectively, all integrals in the right-hand side of Equation (201) are computable in terms of elementary functions. In particular, we arrive at the following exact relations via direct integration:

$$
\begin{align*}
& \int(n \cdot(w-v))^{2} n \Theta(n \cdot(w-v)) \\
& \left(g_{1}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}}+g_{0}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{1}(\boldsymbol{w})}{\partial \boldsymbol{x}}\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=\frac{4 \pi}{15} \frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho^{2} \boldsymbol{S}\right),  \tag{203a}\\
& \int(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \boldsymbol{n} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n}^{T} \frac{\partial^{2} g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}^{2}} \boldsymbol{n} g_{0}(\boldsymbol{v}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
& =\frac{8 \sqrt{\pi}}{15} \frac{\partial}{\partial x} \cdot\left(\rho^{2} \sqrt{\theta}\left(\frac{\partial u}{\partial x}+\frac{\partial u^{T}}{\partial x}+\left(\frac{\partial}{\partial x} \cdot u\right) \boldsymbol{I}\right)\right),  \tag{203b}\\
& \int(n \cdot(w+v))(n \cdot(w-v))^{2} \Theta(n \cdot(w-v)) \\
& \left(g_{1}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}}+g_{0}(\boldsymbol{v}) \boldsymbol{n} \cdot \frac{\partial g_{1}(\boldsymbol{w})}{\partial \boldsymbol{x}}\right) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}=\frac{8 \pi}{15} \frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho^{2}\left(\boldsymbol{S} \boldsymbol{u}+\frac{3}{2} \boldsymbol{q}\right)\right),  \tag{203c}\\
& \int(\boldsymbol{n} \cdot(\boldsymbol{w}+\boldsymbol{v}))(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v}))^{2} \Theta(\boldsymbol{n} \cdot(\boldsymbol{w}-\boldsymbol{v})) \boldsymbol{n}^{T} \frac{\partial^{2} g_{0}(\boldsymbol{w})}{\partial \boldsymbol{x}^{2}} \boldsymbol{n} g_{0}(\boldsymbol{v}) \mathrm{d} \boldsymbol{n} \mathrm{~d} \boldsymbol{w} \mathrm{~d} \boldsymbol{v}= \\
& =\frac{16 \sqrt{\pi}}{15} \frac{\partial}{\partial x} \cdot\left(\rho^{2} \sqrt{\theta}\left[\left(\frac{\partial u}{\partial x}+\frac{\partial u^{T}}{\partial x}+\left(\frac{\partial}{\partial x} \cdot u\right) \boldsymbol{I}\right) \boldsymbol{u}+\frac{5}{2} \frac{\partial \theta}{\partial x}\right]\right) . \tag{203d}
\end{align*}
$$

With the expressions above, the transport equations in Equation (201) become

$$
\begin{align*}
& \frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\theta \boldsymbol{I}\right)\right)=-\frac{4}{\rho_{s p}}\left(1+\sigma R_{*}\right) \frac{\partial}{\partial x} \cdot\left(\rho^{2} \theta\right)- \\
& \quad-\sigma \frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\left(1+\frac{8 \rho}{5 \rho_{s p}}\right) \rho \boldsymbol{S}\right)+\frac{8 \sigma}{5 \sqrt{\pi} \rho_{s p}} \frac{\partial}{\partial x} \cdot\left(\rho^{2} \sqrt{\theta}\left(\frac{\partial \boldsymbol{u}}{\partial \boldsymbol{x}}+\frac{\partial \boldsymbol{u}^{T}}{\partial \boldsymbol{x}}+\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right)\right) \tag{204a}
\end{align*}
$$

$$
\begin{align*}
& \frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho(\epsilon+\theta) \boldsymbol{u})=-\frac{4}{\rho_{s p}}\left(1+\sigma R_{*}\right) \frac{\partial}{\partial x} \cdot\left(\rho^{2} \theta \boldsymbol{u}\right)- \\
&-\sigma \frac{\partial}{\partial x} \cdot\left(\left(1+\frac{8 \rho}{5 \rho_{s p}}\right) \rho \boldsymbol{S} \boldsymbol{u}\right)-\sigma \frac{\partial}{\partial x} \cdot\left(\left(1+\frac{12 \rho}{5 \rho_{s p}}\right) \rho \boldsymbol{q}\right)+ \\
&+\frac{8 \sigma}{5 \sqrt{\pi} \rho_{s p}} \frac{\partial}{\partial x} \cdot\left(\rho^{2} \sqrt{\theta}\left[\left(\frac{\partial \boldsymbol{u}}{\partial x}+\frac{\partial \boldsymbol{u}^{T}}{\partial x}+\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right) \boldsymbol{u}+\frac{5}{2} \frac{\partial \theta}{\partial x}\right]\right) . \tag{204b}
\end{align*}
$$

Substituting the Newton and Fourier laws for $\boldsymbol{S}$ and $\boldsymbol{q}$ from Equation (199), we close Equation (204) under $\rho, u$ and $\theta$ :

$$
\begin{align*}
\frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial \boldsymbol{x}} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\theta \boldsymbol{I}\right)\right) & =-\frac{4}{\rho_{s p}}\left(1+\sigma R_{*}\right) \frac{\partial\left(\rho^{2} \theta\right)}{\partial x}+ \\
& +\frac{\partial}{\partial x} \cdot\left(\mu\left(\left(1+a_{1}\right)\left(\frac{\partial \boldsymbol{u}}{\partial x}+\frac{\partial \boldsymbol{u}^{T}}{\partial x}\right)-\frac{2}{3}\left(1+a_{2}\right)\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right)\right)  \tag{205a}\\
\frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho(\epsilon+\theta) \boldsymbol{u})= & -\frac{4}{\rho_{s p}}\left(1+\sigma R_{*}\right) \frac{\partial}{\partial x} \cdot\left(\rho^{2} \theta \boldsymbol{u}\right)+\frac{15}{4} \frac{\partial}{\partial x} \cdot\left(\mu\left(1+a_{3}\right) \frac{\partial \theta}{\partial x}\right)+ \\
& +\frac{\partial}{\partial x} \cdot\left(\mu\left(\left(1+a_{1}\right)\left(\frac{\partial \boldsymbol{u}}{\partial x}+\frac{\partial \boldsymbol{u}^{T}}{\partial x}\right)-\frac{2}{3}\left(1+a_{2}\right)\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right) \boldsymbol{u}\right) . \tag{205b}
\end{align*}
$$

Above in Equation (205), the expressions for $a_{1}, a_{2}$ and $a_{3}$ are given via

$$
\begin{align*}
& a_{1}\left(\frac{\rho}{\rho_{s p}}\right)=\frac{16 \rho}{5 \rho_{s p}}\left(1+\frac{4 \rho}{5 \rho_{s p}}\left(1+\frac{12}{\pi}\right)\right)  \tag{206a}\\
& a_{2}\left(\frac{\rho}{\rho_{s p}}\right)=\frac{16 \rho}{5 \rho_{s p}}\left(1+\frac{4 \rho}{5 \rho_{s p}}\left(1-\frac{18}{\pi}\right)\right)  \tag{206b}\\
& a_{3}\left(\frac{\rho}{\rho_{s p}}\right)=\frac{24 \rho}{5 \rho_{s p}}\left(1+\frac{2 \rho}{15 \rho_{s p}}\left(9+\frac{32}{\pi}\right)\right) \tag{206c}
\end{align*}
$$

Observe that there are two kinds of $O(\sigma)$-terms in Equation (205): the viscous terms with $\mu$, and the advection terms with $\sigma R_{*}$. Clearly, the viscous terms cannot be neglected, since it is common in practice for the velocity and temperature to have large second derivatives, and thus the contribution from those terms can be measurable even if $\sigma$ is small. On the other hand, observe that the advection terms with $\sigma R_{*}$ contain only the first derivatives, with the product $\sigma R_{*}$ being a constant correction of order $\sigma$ to 1 . From the practical standpoint, the $\sigma R_{*}$-correction is unlikely to affect the solution to a measurable extent in most practical situations, and thus can be discarded in a typical real-world scenario. The resulting transport equations become

$$
\begin{equation*}
\frac{\partial \rho}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho \boldsymbol{u})=0 \tag{207a}
\end{equation*}
$$

$$
\begin{align*}
& \frac{\partial(\rho \boldsymbol{u})}{\partial t}+\frac{\partial}{\partial x} \cdot\left(\rho\left(\boldsymbol{u} \boldsymbol{u}^{T}+\left(1+\frac{4 \rho}{\rho_{s p}}\right) \theta \boldsymbol{I}\right)\right)= \\
&  \tag{207b}\\
& =\frac{\partial}{\partial x} \cdot\left(\mu\left(\left(1+a_{1}\right)\left(\frac{\partial \boldsymbol{u}}{\partial x}+\frac{\partial u^{T}}{\partial x}\right)-\frac{2}{3}\left(1+a_{2}\right)\left(\frac{\partial}{\partial x} \cdot \boldsymbol{u}\right) \boldsymbol{I}\right)\right)
\end{align*}
$$

$$
\begin{align*}
\frac{\partial(\rho \epsilon)}{\partial t}+\frac{\partial}{\partial x} \cdot(\rho(\epsilon+(1 & \left.\left.\left.+\frac{4 \rho}{\rho_{s p}}\right) \theta\right) \boldsymbol{u}\right)=\frac{15}{4} \frac{\partial}{\partial x} \cdot\left(\mu\left(1+a_{3}\right) \frac{\partial \theta}{\partial x}\right)+ \\
& +\frac{\partial}{\partial x} \cdot\left(\mu\left(\left(1+a_{1}\right)\left(\frac{\partial u}{\partial x}+\frac{\partial u^{T}}{\partial x}\right)-\frac{2}{3}\left(1+a_{2}\right)\left(\frac{\partial}{\partial x} \cdot u\right) \boldsymbol{I}\right) \boldsymbol{u}\right) \tag{207c}
\end{align*}
$$

Following the convention set forth in [19], we name the equations above the "Enskog-Navier-Stokes" equations. It is interesting that similar equations were obtained in [45], but the terms which correspond to the Newton and Fourier laws in Equation (199) were missing. For
$\rho / \rho_{s p} \ll 1$ (dilute gas), the coefficients $a_{1}, a_{2}, a_{3}$ disappear, and the equations in Equation (207) become the conventional Navier-Stokes equations [13,42].

Above, in Equation (207), observe that the presence of viscous terms is entirely due to the fact that we treat the sphere diameter $\sigma$ as being finitely small; indeed, if we formally send $\sigma \rightarrow 0$ as in the constant-density hydrodynamic limit above, the viscous terms in Equation (207) disappear, and the Enskog-Euler equations in Equation (189) emerge instead.

## 8. Summary

In the current work, we start by examining the conventional derivation of the Boltzmann equation $[10,12,13]$ from the Liouville equation for the hard sphere dynamics [ $8,9,11]$. The conventional derivation consists, roughly, of two parts-the first part is the derivation of the BBGKY hierarchy [3-5] from the Liouville equation under the assumption that the solution of the Liouville equation is known, while the second part is the derivation of the Boltzmann equation from the leading order BBGKY identity by, first, altering the integrand under the collision integral (Boltzmann hierarchy), and, second, factorizing the joint probability distribution of two spheres into the product of the single-sphere distributions. We observe that the second part of the derivation contradicts the assumptions under which the first part of the derivation was carried out. We also note that some of the observed contradictions are similar to the objection pointed out by Loschmidt [14].

At the same time, it is known from observations and experiments that the Boltzmann equation is de facto an excellent model for a dilute gas in practical scenarios. We, therefore, propose that there should exist a different underlying dynamical model of the hard sphere gas, which does not violate Loschmidt's objection and, at the same time, leads to the Boltzmann equation in the dilute gas approximation. We subsequently formulate a random process with an infinitesimal generator, which triggers the necessary velocity jumps via a suitable point process [28,29] whenever the collision condition is detected. This process is a Lévy-type Feller process [25,30], and the general form of its characteristic function is given by Courrège's theorem [31]. We subsequently formulate the forward Kolmogorov equation [24] for the probability density of the random dynamics, compute those of its steady states that are uniform in the velocity variables, and show that the corresponding Kullback-Leibler entropy [15] of the complete system of spheres is a non-increasing function of time.

We find that, in the case of many spheres, which are distributed independently and identically at the initial time, the corresponding steady state is uniform not only in velocities, but also in the coordinates of the spheres, except for the "contact zones" (that is, the sets of coordinates which satisfy the collision condition). With the help of the computed marginal distributions of these steady states, we derive the forward equation for the dynamics of a single sphere under the assumption that the distribution of the full system is invariant under the reordering of the spheres.

As the total number of spheres in the system becomes large, we find that, for the limiting dynamics of thin contact zones and impenetrable spheres, the forward equation becomes a variant of the Enskog equation [16-20]. Further, assuming that the gas is dilute, we arrive at the Boltzmann equation [7-10,12,13]. Finally, we find that, in the hydrodynamic limit of constant-density spheres, the resulting Enskog-Euler and Enskog-Navier-Stokes equations acquire additional effects in both the advective and viscous terms, which are absent in the conventional Euler and Navier-Stokes equations derived from the Boltzmann equation.

Funding: The work was supported by the Office of Naval Research grant N00014-15-1-2036.
Acknowledgments: The author thanks Ibrahim Fatkullin and Peter Kramer for helpful discussions.
Conflicts of Interest: The author declares no conflict of interest.

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