

## Article

# Simulation and Analysis of the Influence of Sounding Rocket Outgassing on In-Situ Atmospheric Detection

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**Abstract:** The Meridian Project's sounding rocket mission uses a mass spectrometer to conduct in-situ atmospheric detection. In order to assess the influence of surface material outgassing and the attitude control jet on the spectrometer's detection, a sounding rocket platform was modeled and simulated. Using the physical field simulation software COMSOL and the Monte Carlo method, this study investigated whether the gas molecules from the two cases could enter the in-situ atmospheric mass spectrometer's sensor sampling port after colliding with the background atmosphere. The simulation results show that the influence of surface material outgassing on the in-situ atmospheric detection is very small, even under the conditions of medium solar activity and medium geomagnetic activity, while the influence of the attitude control jet on the in-situ atmospheric detection is large but can be reduced by reducing the low-altitude attitude control operation and decreasing the transmission probability. Through simulation optimization and according to engineering needs, increasing the nozzle outlet cross-sectional area, increasing the temperature of the gas used for attitude control, increasing the nozzle rotation angle, increasing the nozzle outlet angle, or increasing the nozzle center height can reduce the transmission probability. This model can simulate and analyze the influence of both surface material outgassing and attitude control jets on in-situ atmospheric detection, optimize relevant parameters, and provide new ideas for relevant work.

**Keywords:** Meridian Project; sounding rocket; in-situ detection; outgassing; COMSOL; Monte Carlo



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## 1. Introduction

Sounding rockets are a powerful method for the in-situ detection of the atmosphere and ionosphere at 20–200 km. This region is beyond the capability of sounding balloons and is too low for satellites to operate for long periods of time. This makes sounding rockets the best platform for collecting data and observing phenomena. Sounding rockets have the advantages of a simple structure, mature technology, low cost, and controllable risk; meanwhile, they are highly flexible and can be launched quickly for detection when solar activity and geomagnetic activity are intense. Therefore, developed countries, such as Europe [1–5] and the United States [6], as well as Japan [7] and India [8], have been actively developing sounding rockets and conducting missions at an early stage and have achieved many scientific results [9–13]. Today's advanced international sounding rockets can even reach an altitude of 1000 km. Additionally, China began sounding rocket research in 1958, and three generations of sounding rockets have been developed in 50 years [14,15].

The Meridian Project, a major national scientific project, was started by the Chinese Academy of Sciences (CAS) in 2008. The Meridian Project is the most comprehensive, full-featured, and longest-spanning space environment ground-based monitoring network in

the world [16,17]. Additionally, through the efforts of scientists from China and around the world, it has gradually developed into the International Meridian Circle Program [18]. The Meridian Project uses sounding rockets to conduct in-situ detection within the atmospheric environment at 20–320 km to complement ground-based detection, to explore the dynamical processes and mechanisms of the atmosphere, ionosphere, and solar-terrestrial space [19,20], and to provide the necessary environmental parameters for atmospheric science, scientific research [21,22] and engineering projects related to the atmosphere.

Spacecraft contamination has always been an important aspect to consider in spacecraft design and has been studied by many researchers [23–27]. Spacecraft contamination is mainly caused by particles of surface materials outgassing from platforms and payloads, as well as particles generated by thrusters and attitude control operations. Spacecraft contamination affects not only optical instruments but also in-situ mass spectrometers, especially in a thin atmosphere. For optical instruments [28], deposition of contamination can lead to degradation of instrument performance; for mass spectrometers [29–31], contamination does not degrade instrument performance but can produce significant background signals, especially for detection missions in thin atmospheres, which affect the reliability and authenticity of the detection results.

So far, many scholars have used simulation analysis [32–36], in-orbit monitoring [37], and adsorption [38,39] control to study spacecraft contamination. Sounding rocket missions are not very long (typically 5–20 min), and the detection region is the atmosphere. Therefore, simulation and analysis before missions are particularly important to address the issue of sounding rocket contamination. Simulation and analysis can be used to improve and optimize the structure and location of rockets and payloads and to filter noise from the result data after missions to reduce the effects of contamination. This paper focuses on the influence of surface material outgassing and the attitude control jet on the mass spectrometer used for in-situ atmospheric detection in the context of the Meridian Project sounding rocket mission and investigates whether the gas molecules from the above two cases can enter the mass spectrometer's sensor sampling port after colliding with the background atmosphere. The physical field simulation software COMSOL is used to model and simulate the collision process with the Monte Carlo method to explore the influence of different solar and geomagnetic activities, and to support the development of the Meridian Project sounding rocket mission.

## 2. Theory and Methods

### 2.1. Gas Molecular Collision Model

Monte Carlo methods are often used to simulate collision processes. The direct simulation Monte Carlo (DSMC) method is suitable for solving the Boltzmann equation to study rarefied gas flow [32,40,41] and is relatively easy to apply in terms of programming effort and computational time. Additionally, atmospheric motion above 120 km can be regarded as free molecular flow, which is the equilibrium Maxwell distribution and can be described and calculated by the Boltzmann equation. DSMC uses a large number of simulated molecules ( $10^4$ – $10^6$  or more) and records the position, velocity, internal energy, and collision process of each simulated molecule, as well as its state after the collision, so as to simulate the complex physical phenomena of the atmosphere. Collisions between molecules are handled based on probability.

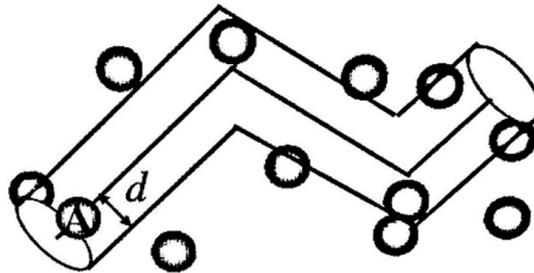
The surface material of the sounding rocket platform can adsorb gas molecules, which can escape when the altitude increases and the pressure decreases. The sounding rocket platform is controlled in flight by jetting gas, and these two cases of gas molecules will collide with the background atmosphere. In 1858, Clausius developed the elastic sphere model of gas molecules, introducing the average collision frequency, i.e., the average number of collisions of a molecule with other molecules per unit of time. This model simplifies molecular motion:

1. The molecular repulsion range is set to a rigid sphere of diameter  $d$ ;
2. The collisions between molecules are elastic;

3. The background molecules are approximately stationary;
4. The incident molecules move at average speed  $\bar{v}$ .

The incident molecular A's motion is shown in Figure 1. The cross-sectional area of the cylinder, i.e., the area of molecular collision, is  $\sigma$ . The distance traveled by A in time  $t$  is  $\bar{v}t$ . The numerical density of gas molecules in the background atmosphere is  $n$ . The total number of molecules centered in the cylinder, i.e., the number of molecules colliding with A, is  $n\sigma\bar{v}t$ . The average collision frequency is shown in Equation (1).

$$\bar{z} = \frac{n\sigma\bar{v}t}{t} = n\sigma\bar{v} \quad (1)$$



**Figure 1.** Diagram of the incident molecular motion.

In 1860, Maxwell used probability statistics to derive the speed distribution function of equilibrium gas molecules, i.e., Maxwell's distribution, and used the distribution to derive the relationship between average speed and absolute temperature, and modified Clausius' formula, as shown in Equation (2):

$$\bar{z} = n\sigma\sqrt{(\bar{v}_A^2 + \bar{v}_B^2)} = n\pi d^2\sqrt{(\bar{v}_A^2 + \bar{v}_B^2)} \quad (2)$$

The diameter  $d$  is given by the Van der Waals radius of molecules; the average speed of background molecules is obtained from the Maxwell distribution; and different molecules' numerical density in the background atmosphere is calculated using the NRLMSISE-00 model [42,43]. The NRLMSISE-00 model was developed by the U.S. Naval Laboratory based on satellite mass spectrometer data and incoherent scattering radar temperature measurements and is one of the most advanced and well-established empirical atmospheric models in the world, which is widely used in data analysis and simulation.

## 2.2. Surface Material Outgassing

Generally, any solid material can adsorb gas molecules. Additionally, when the solid material is placed in a high-vacuum environment, the gas will be desorbed and escape. According to the literature [44,45], Equation (3) is obtained, where  $q$  is the material outgassing rate in  $\text{Pa}\cdot\text{L}/(\text{s}\cdot\text{cm}^2)$ ;  $C$  is the pumping speed in  $\text{L}/\text{s}$  for the pressure change (the following text uses  $\text{m}^3/\text{s}$  as the unit);  $p_1$  and  $p_2$  are the values before and after the pressure change in  $\text{Pa}$ ; and  $S$  is the surface area of the material in  $\text{cm}^2$ . According to the mission, the material outgassing rate is  $533.2 \times 10^{-6} \text{ Pa}\cdot\text{L}/(\text{s}\cdot\text{cm}^2)$ .

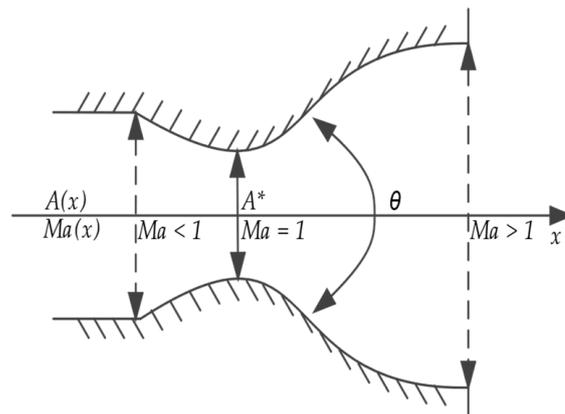
$$q = \frac{C(p_1 - p_2)}{S} \quad (3)$$

## 2.3. Attitude Control Jet

According to the mission, the attitude control uses a de Laval nozzle; its structure is shown in Figure 2, and its design parameters are shown in Table 1. The cold gas propulsion method is used, and the gas is nitrogen ( $\text{N}_2$ ). From the principle of the de Laval nozzle,

it is known that the gas speed is accelerated from subsonic to supersonic as the cross-sectional area of the nozzle outlet changes. The relationship is shown in Equation (4) and Figure 3, where  $\gamma$  is the adiabatic index (1.66 for monatomic molecules and 1.40 for diatomic molecules).

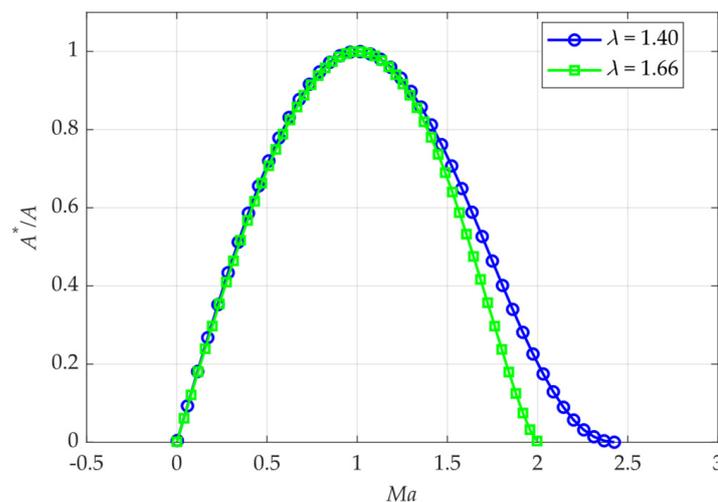
$$\frac{A^*}{A} = \left(\frac{\gamma + 1}{2}\right)^{\frac{1}{\gamma-1}} Ma \left(1 - \frac{\gamma - 1}{\gamma + 1} Ma^2\right)^{\frac{1}{\gamma-1}} \tag{4}$$



**Figure 2.** Structure of the de Laval nozzle. The three straight arrows are the inlet, throat and outlet from left to right, where  $A^*$  is the cross-sectional area of throat and the speed is  $1 Ma$  at this point. The only curved arrow indicates the outlet angle  $\theta$ .

**Table 1.** Design parameters of the de Laval nozzle.

Parameter	Value
Inlet radius	1.30 mm
Throat radius	0.75 mm
Outlet radius	2.60 mm
Outlet angle	30°



**Figure 3.** Relationship between  $Ma$  and  $A^*/A(x)$ . The x-axis represents the Mach number, which is the ratio of the jet speed to the local speed of sound (the blue line intersects the x-axis at  $\sqrt{6}$ , and the green line intersects the x-axis at  $\sqrt{4389/33}$ ). The y-axis represents the ratio of the cross-sectional area of the throat to any cross-sectional area of the nozzle.

According to aerodynamics, the speed of sound is calculated as shown in Equation (5), where  $\gamma$  is the adiabatic index (1.66 for monatomic molecules and 1.40 for diatomic molecules),  $R$  is the gas constant,  $T$  is the absolute temperature, and  $M$  is molar mass. Combining Table 1, Equation (4), and Equation (5), the speed of jetting  $N_2$ —whose temperature is 303 K—to the nozzle outlet in the mission design is about 765 m/s.

$$c = \sqrt{\gamma \frac{p}{\rho}} = \sqrt{\frac{\gamma RT}{M}} \tag{5}$$

2.4. Simulation Model

COMSOL Multiphysics (COMSOL) [46] is a powerful piece of physical field simulation software based on the finite element method that can be used to solve systems of partial differential equations to simulate physical phenomena and can be used to analyze electromagnetic, structural, fluid, heat transfer, and engineering problems. COMSOL’s free molecular flow module has been used to simulate and analyze spacecraft contamination [47–49]. However, the particle tracking module is better suited to deal with the collision process. The particle tracking module [50] supports the calculation of particle trajectories in fluid or electromagnetic fields, including particle–particle, particle–fluid, and particle–field interactions. The particle tracking module has a velocity reinitialization feature that applies random forces to particles during the simulation, allowing various Monte Carlo simulations to be performed. The module has collision nodes for adding random collisions, and further sub-nodes for elasticity, ionization, and excitation can be selected.

The flowchart of the building model is shown in Figure 4. The structure of the model is shown in Figure 5, and the geometrical parameters are shown in Table 2. The direction of the outer normal vector of the nozzle outlet is the same as the positive direction of the x-axis. Figure 5a is placed in the center of Figure 5d.

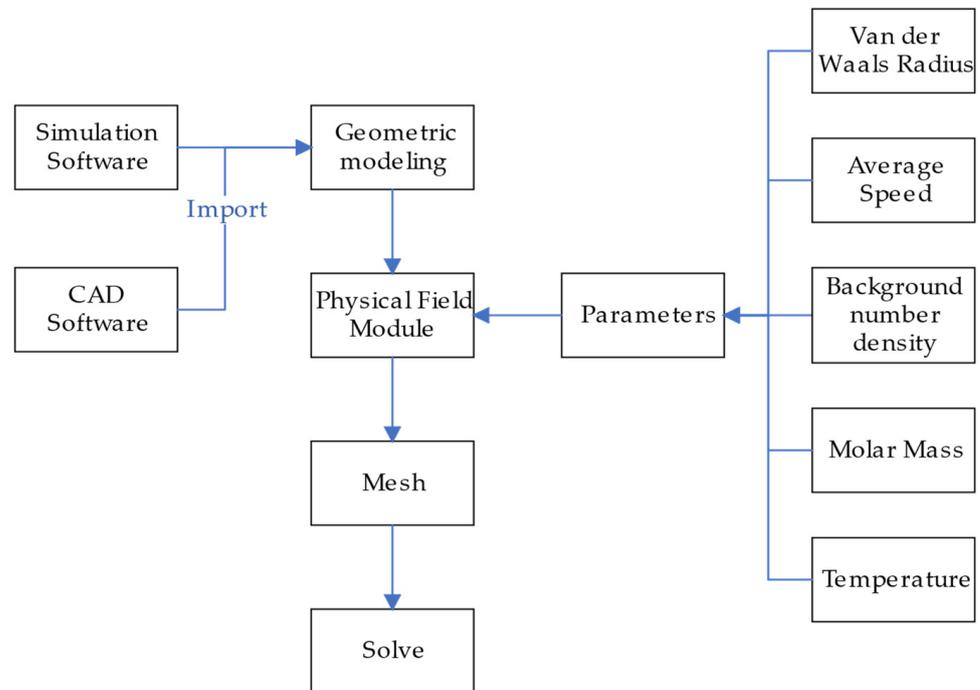
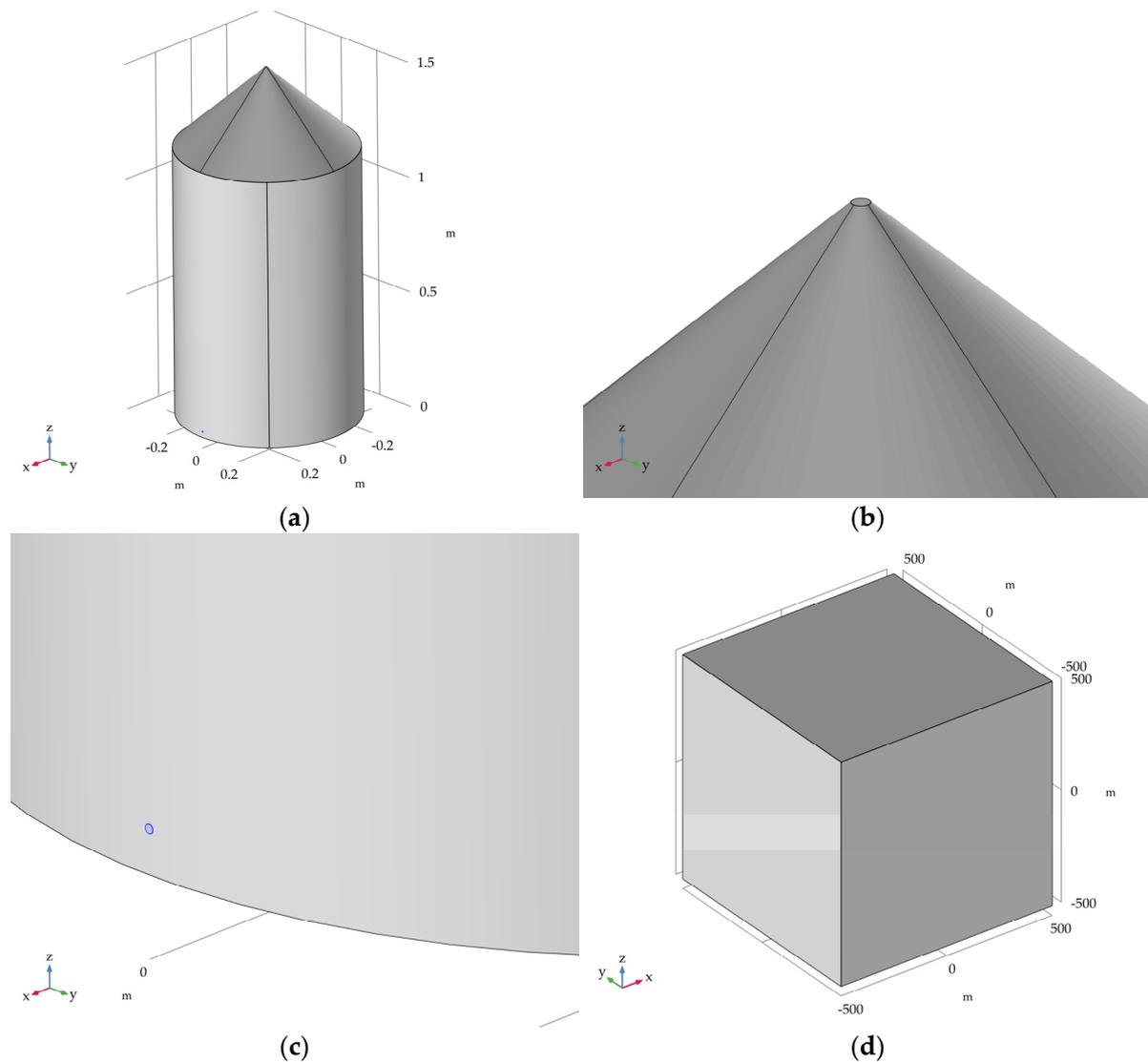


Figure 4. Flowchart of building model.



**Figure 5.** Structure of the model: (a) the sounding rocket platform; (b) the sensor sampling port of the mass spectrometer for in-situ atmospheric detection; (c) the outlet of the de Laval nozzle for the attitude control jet, which is the blue area; and (d) the collision domain used to simulate the collision process.

**Table 2.** Geometric parameters of the model.

Domain	Parameter	Value
Frustum of a cone of platform	Top radius	2.50 mm
	Bottom radius	0.375 m
	Height	0.345 m
Cylinder of platform	Radius	0.375 m
	Height	1.115 m
The outlet of the de Laval nozzle of the platform	Radius	2.60 mm
	Center height	25 mm
Collision	Edge length	500 m

The particle tracking module was selected to simulate the collision process between the gas molecules from the outgassing or jet and the background atmosphere. The points to be considered when setting up the particle tracking module were:

1. The wall condition of the platform surface was set to diffuse scattering, and the wall condition of the collision domain and sensor sampling port was set to frozen;
2. The gas molecules of outgassing were N<sub>2</sub>, oxygen (O<sub>2</sub>), and argon (Ar), and were set on the surface of the platform (the effective surface area is 3.323 m<sup>2</sup>) at the ratio of 78:21:1 and at the same speed as the sounding rocket;
3. The gas molecules of the jet were N<sub>2</sub> and were set at the nozzle outlet with the calculated velocity of the three-dimensional cone distribution during jetting; the jet duration was set to 1 s, and the number of jets was set to 50 to simulate the continuous state;
4. The collision nodes were added, and the background parameters were input, such as numerical density, molar mass, temperature, and average collision frequency;
5. The simulated molecules in the model were only the particles of outgassing or the jet, not the background atmosphere;
6. Due to the limited computing power, the maximum number of simulated molecules in the model was 10<sup>5</sup>.

The model was meshed and solved using a transient study in the time domain.

### 3. Results and Analysis

The mission was planned to detect altitudes of 120–320 km. COMSOL was used to investigate the collision process with different solar and geomagnetic activities. The simulation parameters are shown in Table 3, and the background atmospheric data calculated by the NRLMSISE-00 model under different solar activity, geomagnetic activity, and altitudes are shown in Tables 4–7. The transmission probability obtained from the simulation is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

**Table 3.** Simulation parameters.

Parameter	Value
Altitude	120, 160, 200, 240, 280, 320 km
Jet speed	765 m/s
Medium/low solar activity index F10.7	200/100 sfu
Medium/low geomagnetic activity index Ap	48/8
Average speed of platform at 120 km in simulation	1925.5 m/s
Average speed of platform at 160 km in simulation	1723.5 m/s
Average speed of platform at 200 km in simulation	1498.5 m/s
Average speed of platform at 240 km in simulation	1236.5 m/s
Average speed of platform at 280 km in simulation	908.5 m/s
Average speed of platform at 320 km in simulation	375.5 m/s

**Table 4.** Background atmospheric temperature and numerical density of various molecules at different altitudes under low F10.7 and low Ap environments (LL).

Altitude	Temperature	He	O	N <sub>2</sub>	O <sub>2</sub>	Ar	H
120 km	383 K	$3.77 \times 10^{13} \text{ m}^{-3}$	$7.82 \times 10^{16} \text{ m}^{-3}$	$2.54 \times 10^{17} \text{ m}^{-3}$	$3.90 \times 10^{16} \text{ m}^{-3}$	$1.01 \times 10^{15} \text{ m}^{-3}$	$5.67 \times 10^{12} \text{ m}^{-3}$
160 km	634 K	$2.06 \times 10^{13} \text{ m}^{-3}$	$1.21 \times 10^{16} \text{ m}^{-3}$	$1.46 \times 10^{16} \text{ m}^{-3}$	$1.33 \times 10^{15} \text{ m}^{-3}$	$1.78 \times 10^{13} \text{ m}^{-3}$	$1.02 \times 10^{12} \text{ m}^{-3}$
200 km	680 K	$1.50 \times 10^{13} \text{ m}^{-3}$	$3.80 \times 10^{15} \text{ m}^{-3}$	$2.07 \times 10^{15} \text{ m}^{-3}$	$1.55 \times 10^{14} \text{ m}^{-3}$	$1.12 \times 10^{12} \text{ m}^{-3}$	$5.55 \times 10^{11} \text{ m}^{-3}$
240 km	689 K	$1.15 \times 10^{13} \text{ m}^{-3}$	$1.34 \times 10^{15} \text{ m}^{-3}$	$3.39 \times 10^{14} \text{ m}^{-3}$	$2.03 \times 10^{13} \text{ m}^{-3}$	$8.47 \times 10^{10} \text{ m}^{-3}$	$4.68 \times 10^{11} \text{ m}^{-3}$
280 km	691 K	$8.96 \times 10^{12} \text{ m}^{-3}$	$4.89 \times 10^{14} \text{ m}^{-3}$	$5.81 \times 10^{13} \text{ m}^{-3}$	$2.74 \times 10^{12} \text{ m}^{-3}$	$6.82 \times 10^9 \text{ m}^{-3}$	$4.31 \times 10^{11} \text{ m}^{-3}$
320 km	691 K	$6.99 \times 10^{12} \text{ m}^{-3}$	$1.81 \times 10^{14} \text{ m}^{-3}$	$1.02 \times 10^{13} \text{ m}^{-3}$	$3.76 \times 10^{11} \text{ m}^{-3}$	$5.68 \times 10^8 \text{ m}^{-3}$	$4.04 \times 10^{11} \text{ m}^{-3}$

**Table 5.** Background atmospheric temperature and numerical density of various molecules at different altitudes under low F10.7 and medium Ap environments (LM).

Altitude	Temperature	He	O	N <sub>2</sub>	O <sub>2</sub>	Ar	H
120 km	389 K	$4.00 \times 10^{13} \text{ m}^{-3}$	$8.56 \times 10^{16} \text{ m}^{-3}$	$2.59 \times 10^{17} \text{ m}^{-3}$	$4.02 \times 10^{16} \text{ m}^{-3}$	$1.09 \times 10^{15} \text{ m}^{-3}$	$4.84 \times 10^{12} \text{ m}^{-3}$
160 km	659 K	$2.24 \times 10^{13} \text{ m}^{-3}$	$1.35 \times 10^{16} \text{ m}^{-3}$	$1.56 \times 10^{16} \text{ m}^{-3}$	$1.47 \times 10^{15} \text{ m}^{-3}$	$2.20 \times 10^{13} \text{ m}^{-3}$	$8.27 \times 10^{11} \text{ m}^{-3}$
200 km	714 K	$1.64 \times 10^{13} \text{ m}^{-3}$	$4.39 \times 10^{15} \text{ m}^{-3}$	$2.38 \times 10^{15} \text{ m}^{-3}$	$1.86 \times 10^{14} \text{ m}^{-3}$	$1.53 \times 10^{12} \text{ m}^{-3}$	$4.45 \times 10^{11} \text{ m}^{-3}$
240 km	727 K	$1.27 \times 10^{13} \text{ m}^{-3}$	$1.63 \times 10^{15} \text{ m}^{-3}$	$4.24 \times 10^{14} \text{ m}^{-3}$	$2.68 \times 10^{13} \text{ m}^{-3}$	$1.31 \times 10^{11} \text{ m}^{-3}$	$3.74 \times 10^{11} \text{ m}^{-3}$
280 km	730 K	$9.99 \times 10^{12} \text{ m}^{-3}$	$6.24 \times 10^{14} \text{ m}^{-3}$	$7.96 \times 10^{13} \text{ m}^{-3}$	$4.01 \times 10^{12} \text{ m}^{-3}$	$1.20 \times 10^{10} \text{ m}^{-3}$	$3.45 \times 10^{11} \text{ m}^{-3}$
320 km	730 K	$7.90 \times 10^{12} \text{ m}^{-3}$	$2.43 \times 10^{14} \text{ m}^{-3}$	$1.53 \times 10^{13} \text{ m}^{-3}$	$6.13 \times 10^{11} \text{ m}^{-3}$	$1.15 \times 10^9 \text{ m}^{-3}$	$3.24 \times 10^{11} \text{ m}^{-3}$

**Table 6.** Background atmospheric temperature and numerical density of various molecules at different altitudes under medium F10.7 and low Ap environments (ML).

Altitude	Temperature	He	O	N <sub>2</sub>	O <sub>2</sub>	Ar	H
120 km	398 K	$3.97 \times 10^{13} \text{ m}^{-3}$	$9.01 \times 10^{16} \text{ m}^{-3}$	$2.65 \times 10^{17} \text{ m}^{-3}$	$3.46 \times 10^{16} \text{ m}^{-3}$	$1.03 \times 10^{15} \text{ m}^{-3}$	$3.28 \times 10^{12} \text{ m}^{-3}$
160 km	758 K	$2.09 \times 10^{13} \text{ m}^{-3}$	$1.51 \times 10^{16} \text{ m}^{-3}$	$1.68 \times 10^{16} \text{ m}^{-3}$	$1.10 \times 10^{15} \text{ m}^{-3}$	$2.10 \times 10^{13} \text{ m}^{-3}$	$3.34 \times 10^{11} \text{ m}^{-3}$
200 km	863 K	$1.54 \times 10^{13} \text{ m}^{-3}$	$5.56 \times 10^{15} \text{ m}^{-3}$	$3.23 \times 10^{15} \text{ m}^{-3}$	$1.54 \times 10^{14} \text{ m}^{-3}$	$2.06 \times 10^{12} \text{ m}^{-3}$	$1.48 \times 10^{11} \text{ m}^{-3}$
240 km	895 K	$1.24 \times 10^{13} \text{ m}^{-3}$	$2.41 \times 10^{15} \text{ m}^{-3}$	$7.71 \times 10^{14} \text{ m}^{-3}$	$2.92 \times 10^{13} \text{ m}^{-3}$	$2.68 \times 10^{11} \text{ m}^{-3}$	$1.20 \times 10^{11} \text{ m}^{-3}$
280 km	906 K	$1.01 \times 10^{13} \text{ m}^{-3}$	$1.10 \times 10^{15} \text{ m}^{-3}$	$1.98 \times 10^{14} \text{ m}^{-3}$	$6.13 \times 10^{12} \text{ m}^{-3}$	$3.86 \times 10^{10} \text{ m}^{-3}$	$1.11 \times 10^{11} \text{ m}^{-3}$
320 km	909 K	$8.36 \times 10^{12} \text{ m}^{-3}$	$5.15 \times 10^{14} \text{ m}^{-3}$	$5.24 \times 10^{13} \text{ m}^{-3}$	$1.34 \times 10^{12} \text{ m}^{-3}$	$5.79 \times 10^9 \text{ m}^{-3}$	$1.05 \times 10^{11} \text{ m}^{-3}$

**Table 7.** Background atmospheric temperature and numerical density of various molecules at different altitudes under medium F10.7 and medium Ap environments (MM).

Altitude	Temperature	He	O	N <sub>2</sub>	O <sub>2</sub>	Ar	H
120 km	404 K	$4.21 \times 10^{13} \text{ m}^{-3}$	$9.87 \times 10^{16} \text{ m}^{-3}$	$2.70 \times 10^{17} \text{ m}^{-3}$	$3.57 \times 10^{16} \text{ m}^{-3}$	$1.12 \times 10^{15} \text{ m}^{-3}$	$2.80 \times 10^{12} \text{ m}^{-3}$
160 km	779 K	$2.27 \times 10^{13} \text{ m}^{-3}$	$1.68 \times 10^{16} \text{ m}^{-3}$	$1.78 \times 10^{16} \text{ m}^{-3}$	$1.20 \times 10^{15} \text{ m}^{-3}$	$2.55 \times 10^{13} \text{ m}^{-3}$	$2.73 \times 10^{11} \text{ m}^{-3}$
200 km	894 K	$1.68 \times 10^{13} \text{ m}^{-3}$	$6.31 \times 10^{15} \text{ m}^{-3}$	$3.56 \times 10^{15} \text{ m}^{-3}$	$1.76 \times 10^{14} \text{ m}^{-3}$	$2.65 \times 10^{12} \text{ m}^{-3}$	$1.19 \times 10^{11} \text{ m}^{-3}$
240 km	931 K	$1.36 \times 10^{13} \text{ m}^{-3}$	$2.81 \times 10^{15} \text{ m}^{-3}$	$8.91 \times 10^{14} \text{ m}^{-3}$	$3.52 \times 10^{13} \text{ m}^{-3}$	$3.70 \times 10^{11} \text{ m}^{-3}$	$9.65 \times 10^{10} \text{ m}^{-3}$
280 km	943 K	$1.12 \times 10^{13} \text{ m}^{-3}$	$1.32 \times 10^{15} \text{ m}^{-3}$	$2.40 \times 10^{14} \text{ m}^{-3}$	$7.82 \times 10^{12} \text{ m}^{-3}$	$5.72 \times 10^{10} \text{ m}^{-3}$	$8.92 \times 10^{10} \text{ m}^{-3}$
320 km	947 K	$9.29 \times 10^{12} \text{ m}^{-3}$	$6.35 \times 10^{14} \text{ m}^{-3}$	$6.71 \times 10^{13} \text{ m}^{-3}$	$1.82 \times 10^{12} \text{ m}^{-3}$	$9.26 \times 10^9 \text{ m}^{-3}$	$8.47 \times 10^{10} \text{ m}^{-3}$

### 3.1. The Influence of Surface Material Outgassing on Detection

According to Table 3, the NRLMSISE-00 model, the Ideal Gas Law, and Equation (3), the temperature, density, average molar mass, pressure, pumping speed, and the actual number of outgassing particles for different solar activities, geomagnetic activities, and altitudes can be obtained as shown in Tables 8–11. In Tables 8–11, as the altitude increases, the pressure decreases and the actual number of outgassing particles increases. Due to the limited computing power, the number of outgassing particles at each altitude in the simulation is set based on the proportions in Tables 8–11.

**Table 8.** Data related to outgassing at different altitudes under the LL case.

Altitude	Density	Average Molar Mass	Pressure	Pumping Speed	Actual Number of Outgassing Particles	Proportion
100 km	—	—	$2.99 \times 10^{-2} \text{ Pa}$	—	—	—
120 km	$1.60 \times 10^{-8} \text{ kg/m}^3$	25.921 g/mol	$1.97 \times 10^{-3} \text{ Pa}$	$6.33 \times 10^{-1} \text{ m}^3/\text{s}$	$2.36 \times 10^{17}$	0.230
160 km	$1.07 \times 10^{-9} \text{ kg/m}^3$	22.992 g/mol	$2.46 \times 10^{-4} \text{ Pa}$	$1.03 \times 10^1 \text{ m}^3/\text{s}$	$2.89 \times 10^{17}$	0.282
200 km	$2.06 \times 10^{-10} \text{ kg/m}^3$	20.509 g/mol	$5.68 \times 10^{-5} \text{ Pa}$	$9.37 \times 10^1 \text{ m}^3/\text{s}$	$5.67 \times 10^{17}$	0.553
240 km	$5.27 \times 10^{-11} \text{ kg/m}^3$	18.512 g/mol	$1.63 \times 10^{-5} \text{ Pa}$	$4.38 \times 10^2 \text{ m}^3/\text{s}$	$7.51 \times 10^{17}$	0.732
280 km	$1.59 \times 10^{-11} \text{ kg/m}^3$	17.163 g/mol	$5.34 \times 10^{-6} \text{ Pa}$	$1.61 \times 10^3 \text{ m}^3/\text{s}$	$9.03 \times 10^{17}$	0.880
320 km	$5.37 \times 10^{-12} \text{ kg/m}^3$	16.244 g/mol	$1.90 \times 10^{-6} \text{ Pa}$	$5.15 \times 10^3 \text{ m}^3/\text{s}$	$1.03 \times 10^{18}$	1.000

**Table 9.** Data related to outgassing at different altitudes under the LM case.

Altitude	Density	Average Molar Mass	Pressure	Pumping Speed	Actual Number of Outgassing Particles	Proportion
100 km	—	—	$2.85 \times 10^{-2}$ Pa	—	—	—
120 km	$1.65 \times 10^{-8}$ kg/m <sup>3</sup>	25.775 g/mol	$2.07 \times 10^{-3}$ Pa	$6.69 \times 10^{-1}$ m <sup>3</sup> /s	$2.58 \times 10^{17}$	0.247
160 km	$1.16 \times 10^{-9}$ kg/m <sup>3</sup>	22.881 g/mol	$2.79 \times 10^{-4}$ Pa	$9.89 \times 10^0$ m <sup>3</sup> /s	$3.03 \times 10^{17}$	0.290
200 km	$2.38 \times 10^{-10}$ kg/m <sup>3</sup>	20.510 g/mol	$6.88 \times 10^{-5}$ Pa	$8.45 \times 10^1$ m <sup>3</sup> /s	$5.89 \times 10^{17}$	0.564
240 km	$6.46 \times 10^{-11}$ kg/m <sup>3</sup>	18.602 g/mol	$2.10 \times 10^{-5}$ Pa	$3.71 \times 10^2$ m <sup>3</sup> /s	$7.75 \times 10^{17}$	0.742
280 km	$2.06 \times 10^{-11}$ kg/m <sup>3</sup>	17.293 g/mol	$7.23 \times 10^{-6}$ Pa	$1.29 \times 10^3$ m <sup>3</sup> /s	$9.25 \times 10^{17}$	0.885
320 km	$7.29 \times 10^{-12}$ kg/m <sup>3</sup>	16.408 g/mol	$2.70 \times 10^{-6}$ Pa	$3.91 \times 10^3$ m <sup>3</sup> /s	$1.05 \times 10^{18}$	1.000

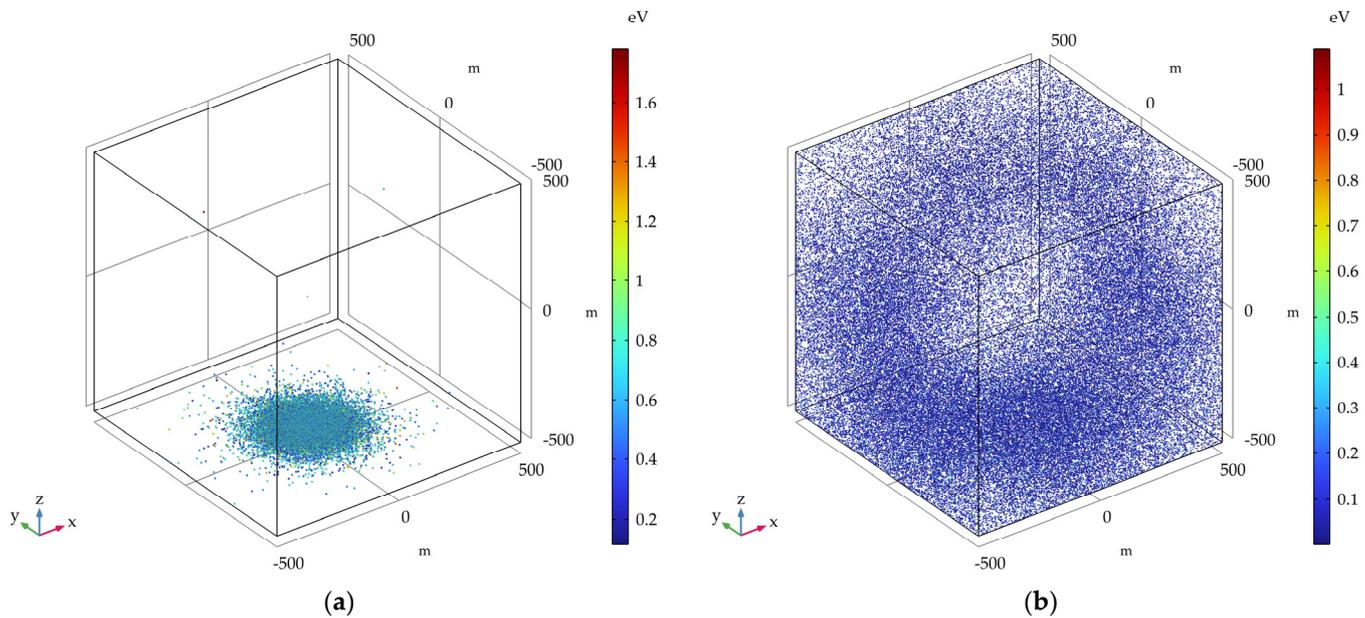
**Table 10.** Data related to outgassing at different altitudes under the ML case.

Altitude	Density	Average Molar Mass	Pressure	Pumping Speed	Actual Number of Outgassing Particles	Proportion
100 km	—	—	$2.96 \times 10^{-2}$ Pa	—	—	—
120 km	$1.66 \times 10^{-8}$ kg/m <sup>3</sup>	25.607 g/mol	$2.15 \times 10^{-3}$ Pa	$6.47 \times 10^{-1}$ m <sup>3</sup> /s	$2.52 \times 10^{17}$	0.227
160 km	$1.24 \times 10^{-9}$ kg/m <sup>3</sup>	22.651 g/mol	$3.46 \times 10^{-4}$ Pa	$9.84 \times 10^0$ m <sup>3</sup> /s	$3.25 \times 10^{17}$	0.293
200 km	$3.07 \times 10^{-10}$ kg/m <sup>3</sup>	20.604 g/mol	$1.07 \times 10^{-4}$ Pa	$7.41 \times 10^1$ m <sup>3</sup> /s	$6.64 \times 10^{17}$	0.598
240 km	$1.02 \times 10^{-10}$ kg/m <sup>3</sup>	19.006 g/mol	$3.99 \times 10^{-5}$ Pa	$2.65 \times 10^2$ m <sup>3</sup> /s	$8.54 \times 10^{17}$	0.769
280 km	$3.90 \times 10^{-11}$ kg/m <sup>3</sup>	17.831 g/mol	$1.65 \times 10^{-5}$ Pa	$7.57 \times 10^2$ m <sup>3</sup> /s	$9.97 \times 10^{17}$	0.898
320 km	$1.63 \times 10^{-11}$ kg/m <sup>3</sup>	17.005 g/mol	$7.25 \times 10^{-6}$ Pa	$1.92 \times 10^3$ m <sup>3</sup> /s	$1.11 \times 10^{18}$	1.000

**Table 11.** Data related to outgassing at different altitudes under the MM case.

Altitude	Density	Average Molar Mass	Pressure	Pumping Speed	Actual Number of Outgassing Particles	Proportion
100 km	—	—	$2.82 \times 10^{-2}$ Pa	—	—	—
120 km	$1.71 \times 10^{-8}$ kg/m <sup>3</sup>	25.451 g/mol	$2.26 \times 10^{-3}$ Pa	$6.84 \times 10^{-1}$ m <sup>3</sup> /s	$2.77 \times 10^{17}$	0.246
160 km	$1.34 \times 10^{-9}$ kg/m <sup>3</sup>	22.510 g/mol	$3.86 \times 10^{-4}$ Pa	$9.45 \times 10^0$ m <sup>3</sup> /s	$3.39 \times 10^{17}$	0.301
200 km	$3.43 \times 10^{-10}$ kg/m <sup>3</sup>	20.531 g/mol	$1.24 \times 10^{-4}$ Pa	$6.77 \times 10^1$ m <sup>3</sup> /s	$6.81 \times 10^{17}$	0.605
240 km	$1.18 \times 10^{-10}$ kg/m <sup>3</sup>	19.003 g/mol	$4.81 \times 10^{-5}$ Pa	$2.33 \times 10^2$ m <sup>3</sup> /s	$8.73 \times 10^{17}$	0.775
280 km	$4.69 \times 10^{-11}$ kg/m <sup>3</sup>	17.873 g/mol	$2.06 \times 10^{-5}$ Pa	$6.42 \times 10^2$ m <sup>3</sup> /s	$1.01 \times 10^{18}$	0.901
320 km	$2.02 \times 10^{-11}$ kg/m <sup>3</sup>	17.073 g/mol	$9.33 \times 10^{-6}$ Pa	$1.58 \times 10^3$ m <sup>3</sup> /s	$1.13 \times 10^{18}$	1.000

The simulation results of the influence of surface material outgassing on detection for different solar activities, geomagnetic activities, and altitudes were zero for all cases, except for the MM case, where there was a transmission probability of 0.0011% at 280 km. The simulation results in the MM case at 120 km and 320 km are shown in Figure 6. According to the simulation results and Table 5, it can be estimated that the number of outgassing particles entering the sampling port at 280 km in the MM case can reach  $1.11 \times 10^{13}$ . According to the sampling port area, rocket speed, simulation time, and numerical density of various molecules in the background atmosphere, it can be calculated that the number of atmospheric particles entering the sampling port at 280 km in the MM case is  $8.45 \times 10^{14}$ . Combining the simulation results and the above two calculation results, it can be concluded that the influence of surface material outgassing on detection is very small.



**Figure 6.** Simulation results of surface material outgassing in the MM case: (a) 120 km; (b) 320 km.

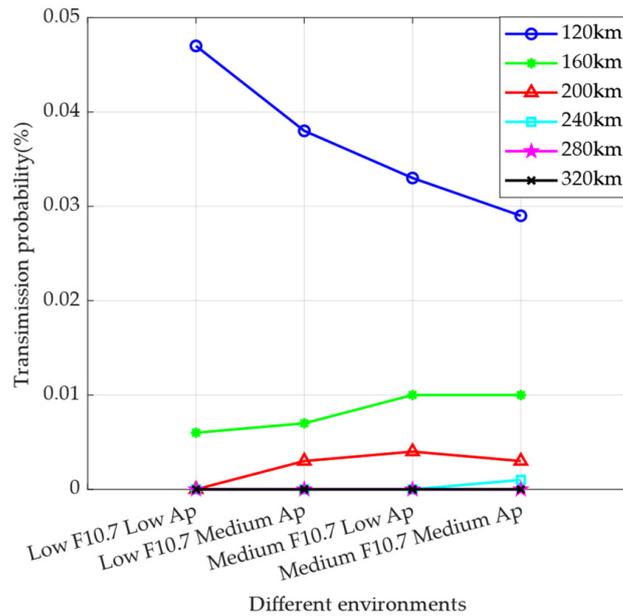
### 3.2. The Influence of Attitude Control Jet on Detection

The simulation results of the influence of the attitude control jet on detection for different solar activity, geomagnetic activity, and altitude are shown in Table 12 and Figure 7, and the simulation results in the LL case at 120 km and 320 km are shown in Figure 8. According to the results, it can be observed that the transmission probability gradually decreases at 120 km as the solar activity and geomagnetic activity become intense, while the opposite situation is observed at 160 km, 200 km, and 240 km, where the transmission probability increases as the solar activity and geomagnetic activity become intense; above 240 km, the transmission probability is 0 and there is no trend.

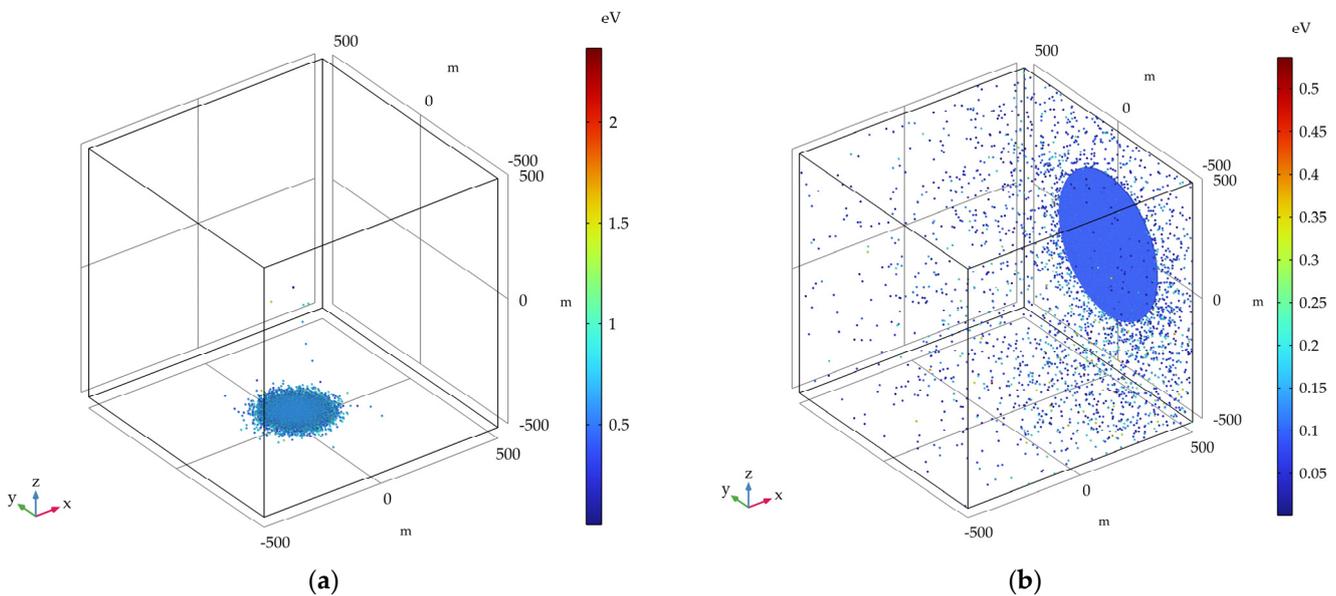
**Table 12.** Simulation results of the attitude control jet for different solar activities, geomagnetic activities, and altitudes.

Altitude	LL	LM	ML	MM
120 km	0.047%	0.038%	0.033%	0.029%
160 km	0.006%	0.007%	0.010%	0.010%
200 km	0%	0.003%	0.004%	0.003%
240 km	0%	0%	0%	0.001%
280 km	0%	0%	0%	0%
320 km	0%	0%	0%	0%

From Tables 4–7, it can be seen that the numerical density decreases as altitude increases. Additionally, the molecules’ average speed in the background atmosphere increases as temperature increases. However, the average collision frequency decreases according to Equation (3). Therefore, the lower the altitude, the more intense the collision. At the same altitude, the temperature increases due to intense solar activity or geomagnetic activity, and the numerical density of the background atmosphere increases; thus, the average collision frequency increases. The transmission probability decreases as the average collision frequency increases because the background atmosphere is dense and collisions are intense at 120 km, while the transmission probability increases as the average collision frequency increases because the background atmosphere is thin at 160 km, 200 km, and 240 km. Above 240 km, the background atmosphere is very thin, the average collision frequency is less than 1 Hz, and the transmission probability decreases to 0.



**Figure 7.** Simulation results of the attitude control jet under different environments. The x-axis represents the gradual intensification of solar activity and geomagnetic activity, where F10.7 represents the degree of solar activity and Ap represents the degree of geomagnetic activity. The y-axis represents the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.



**Figure 8.** Simulation results of the attitude control jet in the LL case: (a) 120 km; (b) 320 km.

Based on the storage conditions of the gas used for attitude control in the mission, with a pressure of 42 MPa, a temperature of 303 K, and a volume of 6 dm<sup>3</sup>, the number of gas molecules used for attitude control was calculated to be  $6.77 \times 10^{25}$ . Based on the sampling port area, the average speed of the platform, the overall flight time, and the numerical density of various molecules in the background atmosphere, the number of atmospheric particles entering the sampling port was calculated to be  $1.53 \times 10^{18}$ . Combining Table 12 with the above two calculation results, it was necessary to reduce the attitude control at low

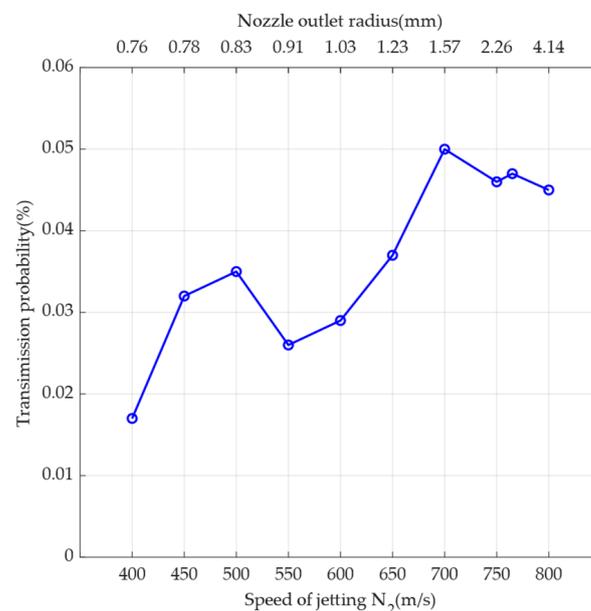
altitudes and decrease the transmission probability. Therefore, the jet-related parameters were optimized to decrease the transmission probability.

### 3.3. The Influence of Nozzle Outlet Cross-Sectional Area on Detection

#### 3.3.1. N<sub>2</sub>

According to the principle of the de Laval nozzle, changing the jet speed must change the outlet cross-sectional area of the nozzle. According to Equations (4) and (5), the speed of sound can be calculated to be about 355 m/s, and the maximum speed can reach 869 m/s after the acceleration of the de Laval nozzle.

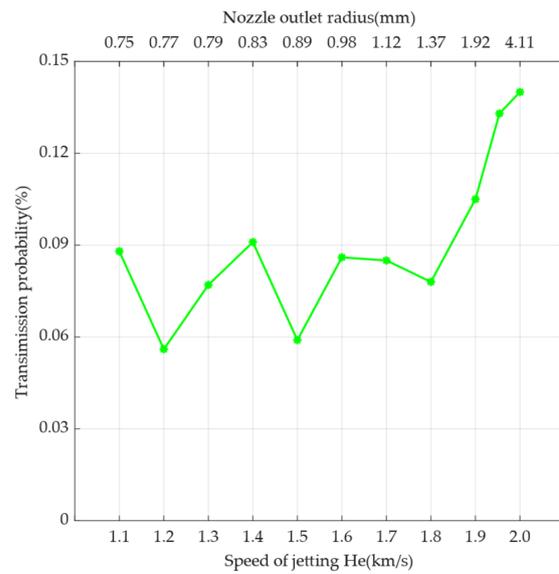
After fixing other parameters and changing the jet speed to 120 km in the LL case, the simulation results are shown in Figure 9. According to Figure 9, it can be observed that the transmission probability shows an overall increasing trend as the speed increases. However, there is a local decrease after 500 m/s and 700 m/s, respectively. From Figure 9, the transmission probability can be reduced by increasing the jet speed to 800 m/s, i.e., by increasing the nozzle outlet cross-sectional area to a radius of 4.14 mm. This can also increase the thrust at the same time to reduce the jet time.



**Figure 9.** Relationship between the speed of jetting N<sub>2</sub> and transmission probability. The simulation results are indicated by circles in the line graph. The primary x-axis at the bottom indicates the speed of jetting N<sub>2</sub>, and the secondary x-axis at the top indicates the nozzle outlet radius at the corresponding speed. The y-axis indicates the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

#### 3.3.2. He

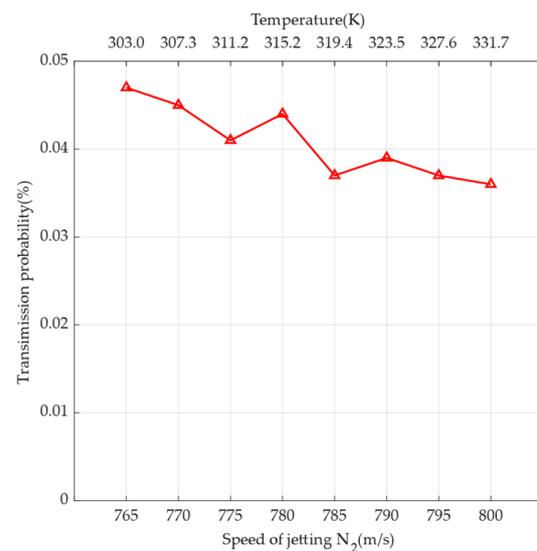
The inert gases N<sub>2</sub> or He are usually chosen as the working gas for cold gas propulsion. If the working gas is changed to He, the speed of sound is about 1022 m/s, and the maximum speed can reach 2051 m/s after acceleration by the de Laval nozzle. The simulation results are shown in Figure 10 for this situation. According to Figure 10, the transmission probability tends to oscillate and then increase as the speed increases; the transmission probability oscillates between 1100 and 1800 m/s and increases after 1800 m/s; the transmission probability of jetting He is higher than the transmission probability of jetting N<sub>2</sub> as a whole. Therefore, N<sub>2</sub> is more suitable for jetting.



**Figure 10.** Relationship between the speed of jetting He and transmission probability. The simulation results are indicated by stars in the line graph. The primary x-axis at the bottom indicates the speed of jetting He, and the secondary x-axis at the top indicates the nozzle outlet radius at the corresponding speed. The y-axis indicates the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

### 3.3.3. N<sub>2</sub> Temperature

According to Equation (4), if the speed of sound is changed, the speed of jetting can be changed while the outlet cross-sectional area remains unchanged. Based on Equations (4) and (5), if the temperature of the gas is increased to 333 K, the mission design of 765 m/s can be increased to 801 m/s. The simulation results for this situation are shown in Figure 11. According to Figure 11, the jet speed increases with increasing temperature, and the overall transmission probability tends to decrease. Therefore, the temperature of the gas can be increased based on engineering needs.

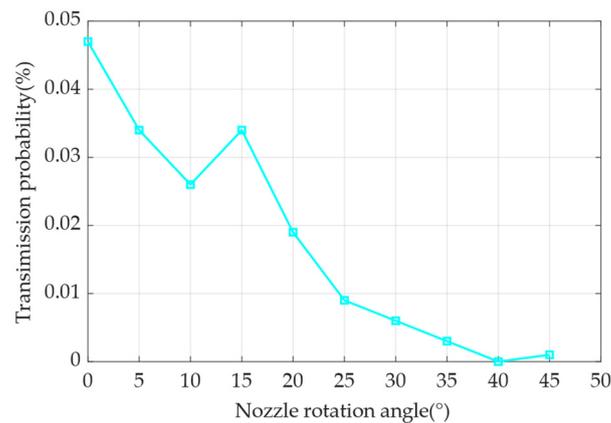


**Figure 11.** Relationship between speed of jetting N<sub>2</sub> and transmission probability. The simulation results are indicated by triangles in the line graph. The primary x-axis at the bottom indicates the speed of jetting N<sub>2</sub>, and the secondary x-axis at the top indicates the temperature at the corresponding speed. The y-axis indicates the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

### 3.4. The Influence of Nozzle Rotation Angle and Outlet Angle on Detection

#### 3.4.1. Rotation Angle

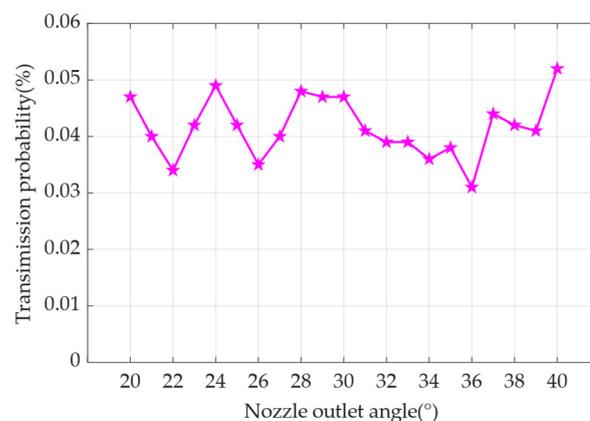
After fixing other parameters and increasing the nozzle rotation angle in the LL case to 120 km, the simulation results are shown in Figure 12. According to Figure 12, the larger the nozzle rotation angle, the smaller the transmission probability. When the nozzle rotation angle is greater than  $25^\circ$ , the transmission probability decreases by an order of magnitude. Additionally, when the nozzle rotation angle increases by more than  $40^\circ$ , the transmission probability basically decreases to zero. From Figure 12, it can be concluded that increasing the nozzle rotation angle can effectively reduce the transmission probability.



**Figure 12.** Relationship between nozzle rotation angle and transmission probability. The simulation results are indicated by squares in the line graph. The x-axis represents the nozzle rotation angle, which is the angle between the outer normal vector of the nozzle outlet and the x-axis in Figure 5. The y-axis represents the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

#### 3.4.2. Outlet Angle

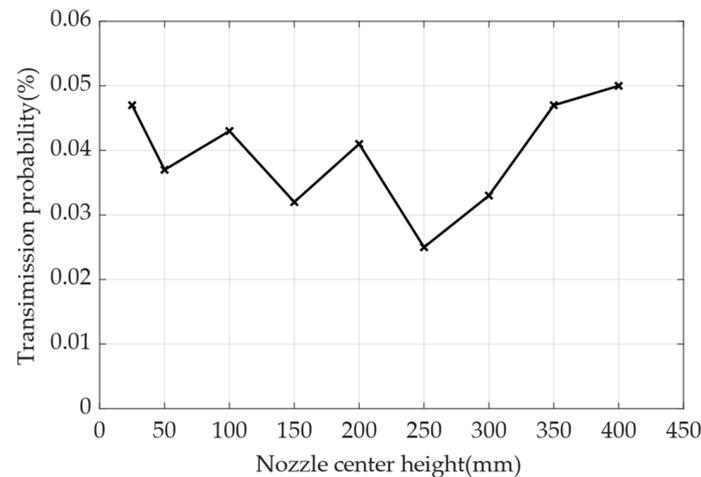
The simulation results are shown in Figure 13, with the other parameters fixed at 120 km in the LL case. According to Figure 13, the transmission probability oscillates as the outlet angle increases. The minimum transmission probability is 0.034% for  $22^\circ$ , and the maximum transmission probability is 0.052% for  $40^\circ$ . From Figure 13, the outlet angle can be increased in the range of  $30^\circ$  to  $36^\circ$ .



**Figure 13.** Relationship between nozzle outlet angle and transmission probability. The simulation results are indicated by pentagrams in the line graph. The x-axis represents the nozzle outlet angle, as shown in Figure 2 and Table 1. The y-axis represents the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

### 3.5. The Influence of Nozzle Center Height on Detection

After fixing other parameters and changing the center height of the nozzle, the simulation results are shown in Figure 14. From Figure 14, it can be seen that the transmission probability with the increase of center height decreases and then increases. The decreasing interval shows oscillating falls, and the minimum transmission probability is 0.025% at 250 mm. Then, in increasing intervals, there is a maximum transmission probability of 0.050% at 400 mm. Therefore, the nozzle center height can be increased to a range of less than 250 mm to reduce the probability of transmission.



**Figure 14.** Relationship between nozzle center height and transmission probability. The simulation results are indicated by x-marks in the line graph. The x-axis represents the nozzle center height, as shown in Figure 5c, which is the location of the nozzle. The y-axis represents the transmission probability, which is the ratio of the number of particles received at the sensor sampling port to the total number of particles.

## 4. Conclusions

In this paper, the collision process between gas molecules from surface materials outgassing and the attitude control jet with the background atmosphere was modeled, simulated, and analyzed using the physical field simulation software COMSOL and the Monte Carlo method in the context of the Meridian Project sounding rocket mission. Using the model, it can be investigated whether the gas molecules of two cases can enter the sampling port of the mass spectrometer sensor for in-situ atmospheric detection, and the methods to reduce the transmission probability can be investigated by simulating and optimizing the parameters related to the jet and obtaining the following conclusions:

1. A simulation model based on COMSOL and Monte Carlo was proposed to simulate and analyze the influence of surface material outgassing and attitude control jet on sounding rocket detection under different solar activity, geomagnetic activity, and altitude;
2. Regardless of medium or low solar activity or medium or low geomagnetic activity, surface material outgassing has little influence on sounding rocket detection. However, a low-altitude attitude control jet has a greater influence on sounding rocket detection, which can be reduced by reducing the number of low-altitude attitude controls and decreasing the transmission probability;
3. According to the simulation, the transmission probability can be reduced by increasing the cross-sectional area of the de Laval nozzle outlet or increasing the gas temperature for attitude control within the allowable range of the project. Increasing the nozzle's rotation angle, the outlet angle within  $36^\circ$ , and the center height within 250 mm can decrease the transmission probability;
4. Since the NRLMSISE-00 model is higher than the actual atmospheric measurement data, the actual transmission probability should be lower than the calculation results of the simulation.

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