

Article

The Gibbs Paradox, the Landauer Principle and the Irreversibility Associated with Tilted Observers

Luis Herrera

Instituto Universitario de Física Fundamental y Matemáticas, Universidad de Salamanca, Salamanca 37007, Spain; lherrera@usal.es; Tel.: +34-923261426

Academic Editors: Lawrence P. Horwitz and Asher Yahalom Received: 10 February 2017; Accepted: 9 March 2017; Published: 11 March 2017

Abstract: It is well known that, in the context of General Relativity, some spacetimes, when described by a congruence of comoving observers, may consist of a distribution of a perfect (non–dissipative) fluid, whereas the same spacetime as seen by a "tilted" (Lorentz–boosted) congruence of observers may exhibit the presence of dissipative processes. As we shall see, the appearance of entropy-producing processes are related to the high dependence of entropy on the specific congruence of observers. This fact is well illustrated by the Gibbs paradox. The appearance of such dissipative processes, as required by the Landauer principle, are necessary in order to erase the different amount of information stored by comoving observers, with respect to tilted ones.

Keywords: tilted spacetimes; irreversibility; dissipative processes

1. Introduction

"Irreversibility is a consequence of the explicit introduction of ignorance into the fundamental laws." M. Born Observers play an essential role in any physical theory. This is particularly true in Thermodynamics and in General Relativity.

Indeed, in this latter theory, it is well known that a variety of line elements may satisfy the Einstein equations for different (physically meaningful) stress-energy tensors (see [1–12] and references therein). This ambiguity in the description of the source may be related, in some cases, to the arbritariness in the choice of the four-velocity in terms of which the energy-momentum tensor is split.

The above-mentioned arbitrariness, in its turn, is related to the well-known fact that different congruences of observers would assign different four-velocities to a given fluid distribution. In this case, we have in mind the situation when one of the conguences corresponds to comoving observers, whereas the other is obtained by applying a Lorentz boost to the comoving observers.

For example, in the case of the zero curvature Friedmann–Robertson–Walker (FRW) model, we have a perfect fluid solution for observers at rest with respect to the timelike congruence defined by the eigenvectors of the Ricci tensor, whereas for observers moving relative to the previously mentioned congruence of observers, it can also be interpreted as the exact solution for a viscous dissipative fluid [4]. It is worth noting that the relative ("tilting") velocity between the two congruences may be related to a physical phenomenon such as the observed motion of our galaxy relative to the microwave background radiation [9].

Thus, zero curvature FRW models as described by "tilted" observers, will exhibit a dissipative fluid and energy–density inhomogeneity, as well as different values for the expansion scalar and the shear tensor, among other differences, with respect to the "standard" (comoving) observers (see [4] for a comprehensive discussion on this example).

The same phenomenon appears in the tilted versions of the Lemaitre–Tolman–Bondi (LTB) [13–15] (see [16]), the Szekeres spacetimes [17,18] (see [19]), and in many other circumstances (see [20–25] and references therein).



fluids (understood as fluids whose energy-momentum tensors present a non-vanishing heat flux contribution), are not necessarily incompatible with reversible processes (e.g., see [26–28]).

In the context of the standard Eckart theory [29], a necessary condition for the compatibility of an imperfect fluid with vanishing entropy production (in the absence of bulk viscosity) is the existence of a conformal Killing vector field (CKV) χ^{α} such that $\chi^{\alpha} = \frac{V^{\alpha}}{T}$ where V^{α} is the four-velocity of the fluid and *T* denotes the temperature. In the context of causal dissipative theories, e.g., [30–35], the existence of such CKV is also necessary for an imperfect fluid to be compatible with vanishing entropy production (see [16]).

However, a much more careful analysis of the problem readily shows that the compatibility of reversible processes and the existence of dissipative fluxes becomes trivial if a constitutive transport equation is adopted, since in this latter case such compatibility forces the heat flux vector to vanish as well. In other words, even if ab initio the fluid is assumed to be imperfect (non–vanishing heat flow vector), the imposition of the CKV and the vanishing entropy production condition may cancel the heat flux once a transport equation is assumed (see [36] for a detailed discussion on this point).

In other words, in the presence of a CKV of the kind mentioned before, the assumption of a transport equation whether in the context of the Eckart–Landau theory, or a causal theory, implies that a vanishing entropy production leads to a vanishing heat flux vector. Therefore, under the conditions above, the system is not only reversible but also non-dissipative.

Furthermore, since neither LTB nor the Szekeres spacetimes admit a CKV, we may safely conclude that the heat flux vector appearing in these cases is associated with truly (entropy-producing) dissipative processes.

The main purpose of this work is to explain the origin of such processes.

2. Comoving and Tilted Observers

Let us consider a congruence of observers which are comoving with a dissipationless dust distribution, then the four-velocity for that congruence, in some globally defined coordinate sytem, reads

$$v^{\mu} = (1, 0, 0, 0). \tag{1}$$

In order to obtain the four-velocity corresponding to the tilted congruence (in the same globally-defined coordinate system), one proceeds as follows.

We have first to perform a (locally defined) coordinate transformation to the Locally Minkowskian Frame (LMF). Denoting by L^{ν}_{μ} the local coordinate transformation matrix, and by \bar{v}^{α} the components of the four velocity in such LMF, we have:

$$\bar{v}^{\mu} = L^{\mu}_{\nu} v^{\nu}. \tag{2}$$

Next, let us perform a Lorentz boost from the LMF associated with \bar{v}^{α} , to the (tilted) LMF with respect to which a fluid element is moving with some, non-vanishing three–velocity.

Then the four-velocity in the tilted LMF is defined by:

$$\tilde{v}_{\beta} = \Lambda^{\alpha}_{\beta} \bar{v}_{\alpha}, \tag{3}$$

where Λ^{α}_{β} denotes the Lorentz matrix.

Finally, we have to perform a transformation from the tilted LMF, back to the (global) frame associated to the line element under consideration. Such a transformation, which obviously only exists locally, is defined by the inverse of L^{ν}_{μ} , and produces the four-velocity of the tilted congruence, in our globally defined coordinate system, say V^{α} .

Let us now consider a given spacetime, which according to comoving observers, is sourced by a dissipationless dust distribution, so that the energy momentum-tensor reads

$$T_{\mu\nu}^{\mathsf{C}} = \mu_{\mathsf{C}} v_{\mu} v_{\nu},\tag{4}$$

where *C* stands for comoving and μ_C denotes the energy density, as measured by the comoving observers.

However, for the tilted congruence we may write

$$T_{\alpha\beta}^{T} = (\mu_{T} + P)V_{\alpha}V_{\beta} + Pg_{\alpha\beta} + \Pi_{\alpha\beta} + q_{\alpha}V_{\beta} + q_{\beta}V_{\alpha},$$
(5)

where *T* stands for tilted, and μ_T , q_{α} , *P* and $\Pi_{\alpha\beta}$ denote the energy density, the heat flux, the isotropic pressure, and the anisotropic tensor, as measured by the tilted observers.

Obviously, both energy-momentum tensors are exactly the same, since the metric is the same and therefore the Einstein tensor is the same. However, the way in which the energy-momentum tensor is split is not the same. This simple fact opens the possibility (for tilted observers) to obtain an energy-momentum tensor which describes a quite different picture from the one obtained by comoving observers. In the same order of ideas, it should be emphasized that the kinematical variables (four-acceleration, expansion scalar, shear tensor, vorticity tensor), being defined in terms of the four-velocity, will also differ from their values as measured by the comoving observers.

Among the differences appearing in the tilted congruence, with respect to the comoving one, there is one which raises the most intriguing question, namely: how it is possible that tilted observers may detect irreversible processes, whereas comoving observers describe an isentropic situation?

As we shall see, the answer to the above question is closely related to the fact that the definition of entropy is highly observer-dependent, as illustrated, for example, by the Gibbs paradox.

3. The Gibbs Paradox, the Landauer Principle and the Definition of Entropy

Entropy is a measure of how much is not known (uncertainty). Also known, although usually overlooked, is the fact that physical objects do not have an intrinsic uncertainty (entropy) (see [37] for an enlightening discussion on this issue).

The "subjective" nature of the concept of entropy is clearly illustrated by the Gibbs paradox. In its simplest form, the paradox appears from the consideration of a box divided by a wall in two identical parts, each of which is filled with an ideal gas (at the same pressure and temperature). Then, if the partition wall is removed, the gases of both parts of the box will mix.

Now, if the gases from both sides are distinguishable, the entropy of the system will rise, whereas if they are identical there is no increase in entropy. This leads to the striking conclusion that irreversibility (and thereby entropy), depends on the ability of the observer to distinguish, or not, the gases from both sides of the box. In other words, irreversibility would depend on our knowledge of physics [38], confirming thereby our previous statement that physical objects are deprived of intrinsic entropy. It can only be defined **after** the number of states that can be resolved by the measurements are established. The anthropomorphic nature of entropy has been brought out and discussed in detail by Jaynes [39].

Let us now turn back to our comoving and tilted observers.

If a given physical system is studied by a congruence of comoving observers, this implies at once that the three–velocity of any given fluid element is automatically assumed to vanish, whereas for the tilted observers this variable represents an additional degree of freedom. In other words, the number of possible states in the latter case is much larger than in the former one.

Since, for the comoving observers, the system is dissipationless, it is clear that the increasing of entropy, when passing to the tilted congruence, should imply the presence of dissipative (entropy producing) fluxes, in the tilted congruence.

It is instructive to take a look on this issue from a different perspective, by considering the transition from the tilted congruence to the comoving one. According to the Landauer principle,

[40] (also referred to as the Brillouin principle [41–45]), the erasure of one bit of information stored in a system requires the dissipation into the environment of a minimal amount of energy, the lower bound of which is given by

$$\triangle E = kT \ln 2, \tag{6}$$

where k is the Boltzmann constant and T denotes the temperature of the environment.

In the above, erasure is just a reset operation restoring the system to a specific state, and is achieved by means of an external agent. In other words, one can decrease the entropy of the system by doing work on it, but then one has to increase the entropy of another system (or the environment).

Thus, the Landauer principle is an expression of the fact that logical irreversibility necessarily implies thermodynamical irreversibility.

Now, when passing from the tilted to the comoving congruence, a decrease of entropy occurs, but we have no external agent, and therefore such a decrease of entropy is accounted for by the dissipative flux observed in the tilted congruence (we recall that in the comoving congruence the system is dissipationless).

The point is, that passing from one of the congruences to the other, we usually overlook the fact that both congruences of observers store different amounts of information. Here resides the clue to resolve the quandary mentioned above, about the presence or not of dissipative processes, depending on the congruence of observers, that carry out the analysis of the system.

Before concluding this section, it is necessary to make three remarks:

- The main issue discussed in this work, namely: the presence or not of dissipative processes, depending on the congruence of observers, that carry out the analysis of the system, will remain for any theory of gravity. However, specific details of the dissipative processes observed by the tilted observers, will depend on the theory of gravity under consideration.
- The discussion about the entropy budget of the universe is of the utmost relevance (see [46] and references therein), because its increase is associated with all possible irreversible processes, on all scales. However, in that reference, as well as in the references therein, the issue under consideration is the estimate of entropy as observed by *one given* congruence of observers. The main point of our work is to stress how (and why) any of these estimates, changes when it is evaluated by *different* congruences of observers.
- It goes without saying that, in the context of a covariant theory of gravity (such as General Relativity), a covariant definition of entropy should be invoked. Such a definition can be found in the context of different relativistic dissipative theories (see for example [30–35]). However, we have not made use of them in the text, which explains why we did not refer to this particular issue.

4. Conclusions

We may summarize the main issues addressed in this paper in the following points:

- Uncertainty (entropy) is highly dependent on the observer.
- Comoving and tilted observers store different amounts of information.
- According to the Landauer principle, erasure of information is always accompanied by dissipation (there is a price for forgetting).
- The detection of dissipative processes by tilted observers in physical systems which are described by comoving observers, such as perfect fluids, becomes intelligible in the light of the three previous comments.

Acknowledgments: This work was partially supported by the Spanish Ministerio de Ciencia e Innovación under Research Projects No. FIS2015-65140-P (MINECO/FEDER).

Conflicts of Interest: The author declares no conflict of interest.

References

- 1. King, A.R.; Ellis, G.F.R. Tilted homogeneous cosmological models. Commun. Math. Phys. 1973, 31, 209–242.
- Tupper, B.O.J. The equivalence of electromagnetic fields and viscous fluids in general relativity. *J. Math. Phys.* 1981, 22, 2666–2673.
- 3. Raychaudhuri, A.K.; Saha, S.K. Viscous fluid interpretation of electromagnetic fields. J. Math. Phys. 1981, 22, 2237–2239.
- 4. Coley, A.A.; Tupper, B.O.J. Zero-curvature Friedmann-Robertson-Walker models as exact viscous magnetohydrodynamic. *Astrophys. J.* **1983**, 271, 1–8.
- 5. Coley, A.A.; Tupper, B.O.J. A new look at FRW cosmologies. Gen. Relativ. Gravit. 1983, 15, 977–983.
- 6. Tupper, B.O.J. The equivalence of perfect fluid space-times and viscous magnetohydrodynamic space-times in general relativity. *Gen. Relativ. Gravit.* **1983**, *15*, 849–873.
- Coley, A.A.; Tupper, B.O.J. Exact viscous fluid FRW cosmologies: The case of general k. *Phy. Lett. A* 1984, 100, 495–498.
- 8. Carot, J.; Ibáñez, J. On the viscous fluid interpretation of some exact solutions. J. Math. Phys. 1985, 26, 2282–2285.
- 9. Coley, A.A. Observations and nonstandard FRW models. Astrophys. J. 1987, 318, 487–506.
- 10. Calvao, M.; Salim, J.M. Extended thermodynamics of Friedmann-Robertson-Walker models in the Landau-Lifshitz frame. *Class. Quantum Gravit.* **1992**, *9*, 127.
- 11. Ellis, G.F.R.; Matravers, D.R.; Treciokas, R. An exact anisotropic solution of the Einstein-Liouville equations. *Gen. Relativ. Gravit.* **1983**, *15*, 931–944.
- 12. Maharaj, S.D.; Maartens, R. Exact inhomogeneous Einstein-Liouville solutions in Robertson-Walker space-times. *Gen. Relativ. Gravit.* **1987**, *19*, 499–509.
- 13. Lemaître, G. The Expanding Universe. Gen. Relativ. Gravit. 1997, 29, 641-680.
- 14. Tolman, R.C. Effect of Inhomogeneity on Cosmological Models. Gen. Relativ. Gravit. 1997, 29, 935–943.
- 15. Bondi, H. Spherically Symmetrical Models in General Relativity. Gen. Relativ. Gravit. 1999, 31, 1783–1805.
- 16. Herrera, L.; Di Prisco, A.; Ibáñez, J. Tilted Lemaitre-Tolman-Bondi spacetimes: Hydrodynamic and thermodynamic properties. *Phys. Rev. D* 2011, *84*, 064036.
- 17. Szekeres, P. Quasispherical gravitational collapse. Phys. Rev. D 1975, 12, 2941–2948.
- 18. Szekeres, P. A class of inhomogeneous cosmological models. Commun. Math. Phys. 1975, 41, 55-64.
- 19. Herrera, L.; Di Prisco, A.; Ibáñez, J.; Carot, J. Vorticity and entropy production in tilted Szekeres spacetimes. *Phys. Rev. D* 2012, *86*, 044003.
- 20. Tsagas, C.G.; Kadiltzoglou, M.I. Peculiar Raychaudhuri equation. Phys. Rev. D 2013, 88, 083501.
- 21. Kumar, R.; Srivastava, S. Bianchi type-V cosmological model with purely magnetic solution. *Astrophys. Space Sci.* **2013**, *346*, 567–572.
- 22. Sharif, M.; Tahir, H. Dynamics of tilted spherical star and stability of non-tilted congruence. *Astrophys. Space Sci.* **2014**, *351*, 619–624.
- 23. Fernandez, J.J.; Pascual-Sanchez, J.-F. Tilted Lemaître model and the dark flow. *Proc. Math. Stat.* **2014**, *60*, 361–364.
- 24. Sharif, M.; Zaeem Ul Haq Bhatti, M. Structure scalars and super-Poynting vector of tilted Szekeres geometry. *Int. J. Mod. Phys. D* 2015, 24, 1550014.
- 25. Yousaf, Z.; Bamba, K.; Zaeem Ul Haq Bhatti, M. Role of tilted congruence and f(R) gravity on regular compact objects. *Phys. Rev. D* 2017, *95*, 024024.
- 26. Stephani, H. Introduction to General Relativity; Cambridge University Press: Cambridge, UK, 1982.
- 27. Bedran, M.L.; Calvao, M.O. Reversibility and spacetime symmetries. Class. Quantum Gravit. 1993, 10, 767.
- 28. Triginer, J.; Pavón, D. On the thermodynamics of tilted and collisionless gases in Friedmann-Robertson-Walker spacetimes. *Class. Quantum Gravit.* **1995**, *12*, 199.
- 29. Eckart, C. The Thermodynamics of Irreversible Processes. III. Relativistic Theory of the Simple Fluid. *Phys. Rev.* **1940**, *58*, 919, doi:10.1103/PhysRev.58.919.
- 30. Müller, I. Zum Paradoxon der Warmeleitungstheorie. Z. Phys. 1967, 198, 329-344.
- 31. Israel, W. Nonstationary irreversible thermodynamics: A causal relativistic theory. *Ann. Phys.* **1976**, *100*, 310–331.

- 32. Israel, W.; Stewart, J. Thermodynamics of nonstationary and transient effects in a relativistic gas. *Phys. Lett. A* **1976**, *58*, 213–215.
- 33. Israel, W.; Stewart, J. Transient relativistic thermodynamics and kinetic theory. Ann. Phys. 1979, 118, 341–372.
- 34. Jou, D.; Casas-Vázquez, J.; Lebon, G. Extended irreversible thermodynamics. Rep. Prog. Phys. 1988, 51, 1105.
- 35. Jou, D.; Casas-Vázquez, J.; Lebon, G. Extended Irreversible Thermodynamics; Springer: Berlin, Geramny, 1993.
- Herrera, L.; Di Prisco, A.; Ibáñez, J. Reversible dissipative processes, conformal motions and Landau damping. *Phys. Lett. A* 2012, 376, 899–900.
- 37. Adami, C. What is Information? Philos. Trans. R. Soc. A 2016, 374, 20150230.
- 38. Bais, F.A.; Farmer, J.D. The Physics of Information. arXiv 2007, arXiv:0708.2837v2.
- 39. Jaynes, E.T. Gibbs vs Boltzmann Entropies. Am. J. Phys. 1965, 33, 391-398.
- 40. Landauer, R. Irreversibility and Heat Generation in the Computing Process. IBM Res. Dev. 1961, 5, 183–191.
- 41. Brillouin, L. The Negentropy Principle of Information. J. Appl. Phys. 1953, 24, 1152, doi:10.1063/1.1721463.
- 42. Brillouin, L. Scientific Uncertainty and Information; Academic: New York, NY, USA, 1964.
- 43. Brillouin, L. Science and Information Theory; Academic: New York, NY, USA, 1962.
- 44. Kish, L.B.; Granqvist, C.G. Energy Requirement of Control: Comments on Szilard's Engine and Maxwell's Demon. *arXiv* **2012**, *arXiv*:1110.0197v7.
- 45. Kish, L.B. Moore's law and the energy requirement of computing versus performance. *Proc. IEEE* **2004**, *151*, 190–194.
- 46. Egan, C.A.; Lineweaver, C.H. A Larger Estimate of the Entropy of the Universe. *Astrophys. J.* 2010, 710, 1825–1834.



© 2017 by the author; licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (http://creativecommons.org/licenses/by/4.0/).